

Weak and Electromagnetic Interactions in Nuclei

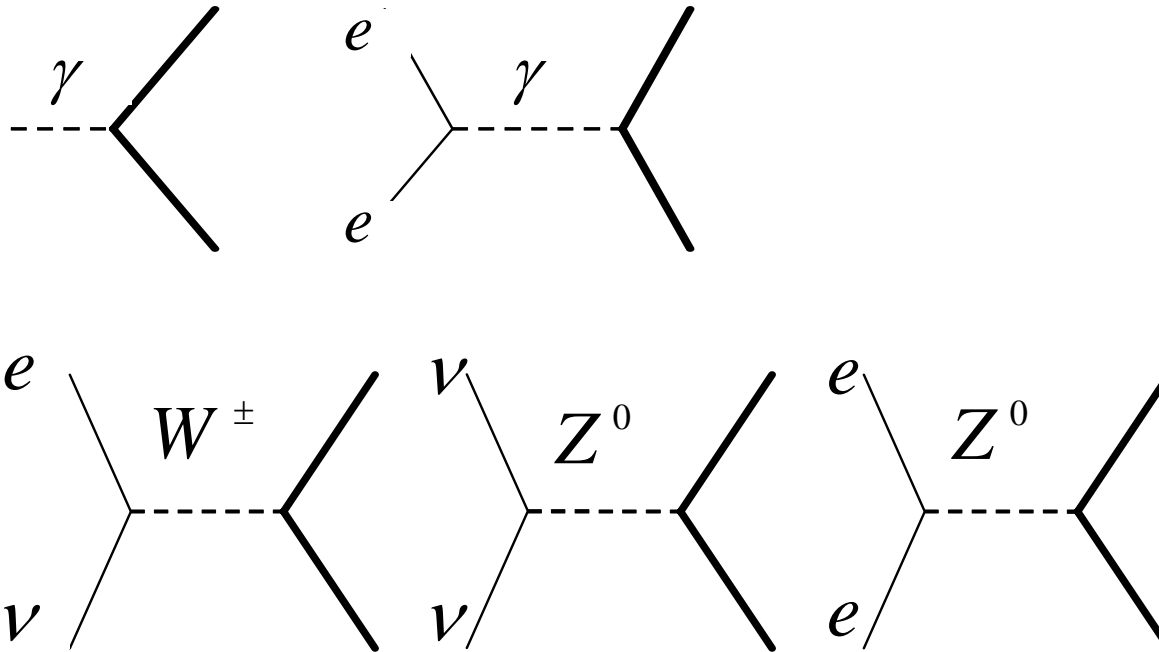
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CSN Summer School
Aug. 28, 2007

Interactions	Situation	Strength (Cross sections)	Handling
EM (electromagnetic)	Known	moderate (enough) $(10^{-36} - 10^{-26} \text{ cm}^2 / \text{sr})$	clear
Weak (W)	Known within the standard model Not well determined part remains at hadron level	weak (small) $(10^{-42} - 10^{-38} \text{ cm}^2)$	clear
Strong (S)	Not completely known at hadron level (need phenomenology)	strong (large) $(\approx 10^{-(27-22)} \text{ cm}^2)$	non-easy (distortion + absorption)

EM and Weak Interactions in Nuclei

- 1. Probe of nuclear structure

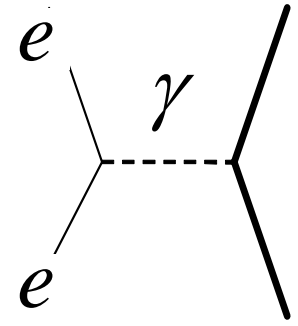


- 2. Probe of weak neutral current

e.g. $\bar{s} \gamma_\mu s$, $\bar{s} \gamma_\mu \gamma_5 s$

EM

$$T_{fi}^{\gamma} = \frac{4\pi\alpha}{q_{\mu}^2} i\bar{e}\gamma_{\mu}e \cdot J_{\mu}^{\gamma}$$



$$J_{\mu}^{\gamma} = i \langle N | \frac{2}{3} \bar{u}\gamma_{\mu}u - \frac{1}{3} \bar{d}\gamma_{\mu}d - \frac{1}{3} \bar{s}\gamma_{\mu}s | N \rangle$$

$$= i \langle N | \frac{1}{6} (\bar{u}\gamma_{\mu}u + \bar{d}\gamma_{\mu}d - 2\bar{s}\gamma_{\mu}s) + \frac{1}{2} (\bar{u}\gamma_{\mu}u - \bar{d}\gamma_{\mu}d) | N \rangle$$

isoscalar (j^8)

isovector (j^3)

$$= i\bar{u}(p') [F_1(q_{\mu}^2)\gamma_{\mu} + \frac{1}{2m} F_2(q_{\mu}^2)\sigma_{\mu\nu}q_{\nu}] u(p)$$

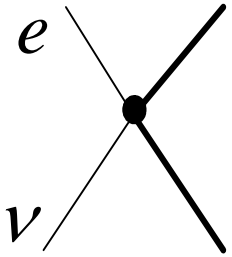
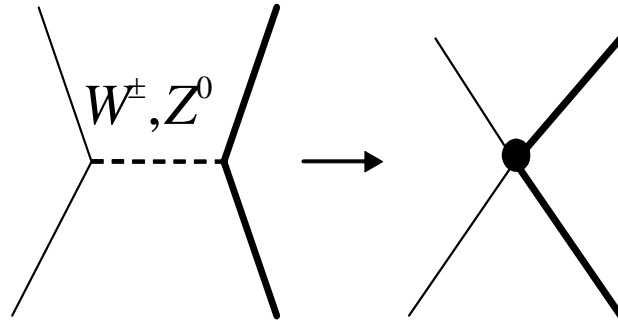
$$= i\bar{u}(p') \frac{1}{2} [(F_1^s + F_1^v \tau_3)\gamma_{\mu} + \frac{1}{2m} (F_2^s + F_2^v \tau_3)\sigma_{\mu\nu}q_{\nu}] u(p)$$

$$F_1^s(0) = F_1^v(0) = 1, \quad F_2^s(0) = -0.12 \quad F_2^v(0) = 3.706$$

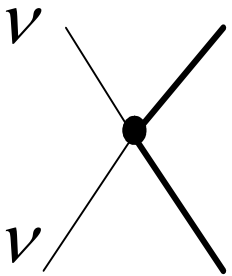
$$\vec{J}^{\gamma} = F_1 \frac{1}{2m} (\vec{p} + \vec{p}') + (F_1 + F_2) \frac{1}{2m} (-i\vec{\sigma} \times \vec{q})$$

$$J_0^{\gamma} = F_1 - (F_1 + F_2) \left(\frac{q^2}{8m^2} - \frac{i}{4m^2} \vec{q} \cdot (\vec{\sigma} \times \vec{q}) \right)$$

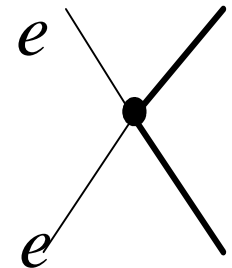
Weak



$$L_{eff}^{(\pm)} = i \frac{G}{\sqrt{2}} \{ [\bar{e} \gamma_\mu (1 + \gamma_5) \nu_e + (e \leftrightarrow \mu)] J_\mu^{(+)} + [\bar{\nu}_e \gamma_\mu (1 + \gamma_5) e + (e \leftrightarrow \mu)] J_\mu^{(-)} \}$$



$$L_{eff}^{(\nu)} = i \frac{G}{\sqrt{2}} [\bar{\nu}_e \gamma_\mu (1 + \gamma_5) \nu_e + (e \rightarrow \mu) + ..] J_\mu^{(0)}$$



$$L_{eff}^{(e)} = -i \frac{G}{\sqrt{2}} [\bar{e} \gamma_\mu (1 + \gamma_5) e - 4 \sin^2 \theta_W \bar{e} \gamma_\mu e] J_\mu^{(0)}$$

Charged Current

$$J_{\mu}^{(+)} = i \langle N | \bar{u} \gamma_{\mu} (1 + \gamma_5) [d \cos \theta_c + s \sin \theta_c] + \bar{c} \gamma_{\mu} (1 + \gamma_5) [-d \sin \theta_c + s \cos \theta_c] | N \rangle$$

$$J_{\mu}^{(+)} = i \cos \theta_c \{ \langle N | \gamma_{\mu} \tau_{\pm} | N \rangle + \langle N | \gamma_{\mu} \gamma_5 \tau_{\pm} | N \rangle \}$$

V(vector)

AV(axial-vector)

$$\langle N | \gamma_{\mu} \tau_{\pm} | N \rangle = \bar{u}(p') [F_1 \gamma_{\mu} + \frac{1}{2m} F_2 \sigma_{\mu\nu} q_{\nu}] \tau_{\pm} u(p) \quad \boxed{\text{CVC}}$$

$$\langle N | \gamma_{\mu} \gamma_5 \tau_{\pm} | N \rangle = \bar{u}(p') [F_A \gamma_5 \gamma_{\mu} - i F_P \gamma_5 q_{\nu}] \tau_{\pm} u(p)$$

$$F_A(0) = -1.26 \quad F_P = \frac{2m F_A}{q^2 + m_{\pi}^2}$$

$$\vec{J} = F_A \vec{\sigma} - (F_1^V + F_2^V) \frac{1}{2m} i \vec{\sigma} \times \vec{q} + F_1^V \frac{1}{2m} (\vec{p} + \vec{p}')$$

$$J_0 = F_1^V + F_A \frac{1}{2m} \vec{\sigma} \cdot (\vec{p} + \vec{p}')$$

Neutral Current

$$\begin{aligned}
 J_\mu^{(0)} &= \frac{i}{2} \langle N | [\bar{u} \gamma_\mu (1 + \gamma_5) u - \bar{d} \gamma_\mu (1 + \gamma_5) d - \bar{s} \gamma_\mu (1 + \gamma_5) s] | N \rangle - 2 \sin^2 \theta_w J_\mu^\gamma \\
 &= \frac{i}{2} \langle N | [(\bar{u} \gamma_\mu u - \bar{d} \gamma_\mu d) + (\bar{u} \gamma_\mu \gamma_5 u - \bar{d} \gamma_\mu \gamma_5 d) - \bar{s} \gamma_\mu s - \bar{s} \gamma_\mu \gamma_5 s] | N \rangle \\
 &\quad \text{isovector-V} \quad \text{isovector-AV} \quad \text{strange-V} \quad \text{strange-AV} \\
 &= V_\mu^3 + A_\mu^3 + V_\mu^S + A_\mu^S - 2 \sin^2 \theta_w J_\mu^\gamma
 \end{aligned}$$

$$V_\mu^3 = \frac{i}{2} \langle N | (\bar{u} \gamma_\mu u - \bar{d} \gamma_\mu d) | N \rangle = \frac{i}{2} \bar{u}(p') [F_1^V \gamma_\mu + \frac{1}{2m} F_2^V \sigma_{\mu\nu} q] \tau_3 u(p)$$

$$A_\mu^3 = \frac{i}{2} \langle N | (\bar{u} \gamma_\mu \gamma_5 u - \bar{d} \gamma_\mu \gamma_5 d) | N \rangle = \frac{i}{2} \bar{u}(p') [F_A \gamma_5 \gamma_\mu - i F_P \gamma_5 q_\nu] \tau_3 u(p)$$

$$V_\mu^S = -\frac{i}{2} \langle N | \bar{s} \gamma_\mu s | N \rangle = -\frac{i}{2} \bar{u}(p') [F_1^S \gamma_\mu + \frac{1}{2m} F_2^S \sigma_{\mu\nu} q] u(p)$$

$$A_\mu^S = -\frac{i}{2} \langle N | \bar{s} \gamma_\mu \gamma_5 s | N \rangle = -\frac{i}{2} G_1^S \bar{u}(p') \gamma_\mu \gamma_5 u(p)$$

$$F_1^S(0) = 0 \quad F_1^S(q_\mu^2 \neq 0), \quad F_2^S, \quad G_1^S \quad \text{not well determined}$$

Note

$$j_{\mu}^3 = \frac{1}{2}(\bar{u}\Gamma u - \bar{d}\Gamma d)$$

$$j_{\mu}^8 = \frac{1}{2\sqrt{3}}(\bar{u}\Gamma u + \bar{d}\Gamma d - 2\bar{s}\Gamma s)$$

$$j_{\mu}^0 = \frac{1}{3}(\bar{u}\Gamma u + \bar{d}\Gamma d + \bar{s}\Gamma s)$$

j_{μ}^0 not well determined

Assumption $A_{\mu}^S = V_{\mu}^S = 0$

$$J_{\mu}^{(0)} = A_{\mu}^3 + V_{\mu}^3 - 2\sin^2\theta_W J_{\mu}^{\gamma}$$

Note: vector part $V_{\mu}^{(0)} = V_{\mu}^3 - 2\sin^2\theta_W J_{\mu}^{\gamma}$

C0 $(G_E^{IV} - 2\sin^2\theta_W G_E) < 0 \mid j_0(qr)Y^{(0)} \mid 0 >$

proton: $\frac{1}{2}G_E^p(1 - 4\sin^2\theta_W)\rho_p(r) = \frac{0.08}{2}G_E^p\rho_p(r)$

neutron: $(\frac{1}{2}G_E^p - 2\sin^2\theta_W G_E^n)\rho_n(r) \cong \frac{1}{2}G_E^p\rho_n(r)$

Cross sections for $(\nu, e^-), (\bar{\nu}, e^+), (\nu, \nu), (\bar{\nu}, \bar{\nu})$

$$d\sigma = 2\pi \sum_h |\langle f | H_w | i \rangle|^2 \delta(W_f - W_i) \frac{d^3k}{(2\pi)^3}$$

$$\frac{d\sigma}{d\Omega_{\mp}} = \frac{G^2}{2\pi^2} \frac{1}{\sqrt{(k_1 \cdot p)^2}} \frac{d^3k}{2\varepsilon_2} \eta_{\mu\nu} W_{\mu\nu}$$

$$\eta_{\mu\nu} = \frac{1}{4} \text{Tr}[\gamma_\mu (1 + \gamma_5) \not{k}_1 \gamma_\nu (1 + \gamma_5) \not{k}_2]$$

$$= 2(k_{1\mu} k_{2\nu} + k_{1\nu} k_{2\mu} - (k_1 \cdot k_2) \delta_{\mu\nu} \pm \varepsilon_{\mu\nu\rho\sigma} k_{1\rho} k_{2\sigma})$$

$$W_{\mu\nu} = W_{\mu\nu}^{VV} + W_{\mu\nu}^{VA} + W_{\mu\nu}^{AA}$$

$$W_{\mu\nu}^{VV} = W_1 \delta_{\mu\nu} + \frac{W_2}{M_T^2} \left(p_\mu - \frac{p \cdot q}{q^2} q_\mu \right) \left(p_\nu - \frac{p \cdot q}{q^2} q_\nu \right)$$

$$W_{\mu\nu}^{AA} = W_3 \delta_{\mu\nu} + \frac{W_4}{M_T^2} p_\mu p_\nu + \frac{W_5}{M_T^2} q_\mu q_\nu$$

$$+ \frac{W_6}{M_T^2} (p_\mu q_\nu + p_\nu q_\mu) + \frac{W_7}{M_T^2} (p_\mu q_\nu - p_\nu q_\mu)$$

$$W_{\mu\nu}^{VA} = \frac{W_8}{M_T^2} \varepsilon_{\mu\nu\rho\sigma} p_\rho q_\sigma$$

Current Conservation

$$q_\mu W_{\mu\nu}^{\nu\nu} = W_{\mu\nu}^{\nu\nu} q_\nu = 0, \quad q_\mu \eta_{\mu\nu}^{\nu\nu} = \eta_{\mu\nu}^{\nu\nu} q_\nu = 0$$

$$\frac{d^2\sigma}{d\Omega d\varepsilon_{2\mp}} = \frac{G^2 \varepsilon_2^2}{2\pi^2} \frac{1}{M_T} \left\{ (W_2 + W_4) \cos^2 \frac{\theta}{2} + 2(W_1 + W_3) \sin^2 \frac{\theta}{2} \right. \\ \left. \mp 2 \frac{W_8}{M_T} \sin \frac{\theta}{2} \left(q^2 \cos^2 \frac{\theta}{2} + \vec{q}^2 \sin^2 \frac{\theta}{2} \right)^{1/2} \right\}$$

Multipole expansions

$$\frac{d^2\sigma}{d\Omega \mp} = \frac{G^2 \varepsilon^2}{2\pi^2} \frac{4\pi}{2J_i + 1} \left\{ \cos^2 \frac{\theta}{2} \left| \langle J_f \parallel M_J(q) - \frac{q_0}{q} L_J(q) \parallel J_i \rangle \right|^2 \right. \\ \left. + \left[\frac{q_\mu^2}{2q^2} \cos^2 \frac{\theta}{2} + \sin^2 \frac{\theta}{2} \right] \sum_{J=1}^{\infty} \left[\left| \langle J_f \parallel T_J^{mag}(q) \parallel J_i \rangle \right|^2 \right. \right. \\ \left. \left. + \left| \langle J_f \parallel T_J^{el}(q) \parallel J_i \rangle \right|^2 \right] \mp \frac{\sin \frac{\theta}{2}}{q} \left(q_\mu^2 \cos^2 \frac{\theta}{2} + \vec{q}^2 \sin^2 \frac{\theta}{2} \right)^{1/2} \right. \\ \left. \sum_{J=1}^{\infty} 2 \operatorname{Re} \left[\langle J_f \parallel T_J^{mag}(q) \parallel J_i \rangle \langle J_f \parallel T_J^{el}(q) \parallel J_i \rangle^* \right] \right\}$$

$$M_{JM}(q) = \int d^3x j_J(qx) Y_{JM}(\hat{x}) \rho(\vec{x})$$

$$L_{JM}(q) = \frac{i}{q} \int d^3x \vec{\nabla} [j_J(qx) Y_{JM}(\hat{x})] \vec{J}(\vec{x})$$

$$T_{JM}^{mag}(q) = \int d^3x [j_J(qx) \vec{Y}_{JJ_1}^M(\hat{x})] \vec{J}(\vec{x})$$

$$T_{JM}^{el}(q) = \frac{1}{q} \int d^3x [\vec{\nabla} \times j_J(qx) \vec{Y}_{JJ_1}^M(\hat{x})] \vec{J}(\vec{x})$$

$q \rightarrow 0$

$$M_{JM} = \frac{q^J}{(2J+1)!!} \int d^3x Y_{JM} J_0$$

$$L_{JM} = \frac{1}{i} \frac{q^{J-1}}{(2J+1)!!} \int d^3x Y_{JM} \vec{\nabla} \cdot \vec{J}(\vec{x})$$

$$T_{JM}^{mag} = -\frac{1}{i} \frac{q^J}{(2J+1)!!} \sqrt{\frac{J+1}{J}} \int d^3x \frac{1}{J+1} [\vec{x} \times \vec{J}(\vec{x})] \cdot \vec{\nabla} x^J Y_{JM}$$

$$T_{JM}^{el} = \frac{1}{i} \frac{q^{J-1}}{(2J+1)!!} \sqrt{\frac{J+1}{J}} \int d^3x x^J Y_{JM} \vec{\nabla} \cdot \vec{J}(\vec{x})$$

$$\text{Vector: } T_{1M}^{\text{mag}} = \frac{i\sqrt{2}}{3} \frac{\hbar q}{2mc} \sqrt{\frac{3}{4\pi}} \sum_i (\vec{\ell}_i + \mu_i \vec{\sigma}_i)$$

$$\langle f | T_{JM}^{\text{el}}(q) | i \rangle \xrightarrow{q \rightarrow 0} -\frac{E_f - E_i}{q} \sqrt{\frac{J+1}{J}} \langle f | M_{JM}(q) | i \rangle$$

$$\text{AV: } T_{1M}^{\text{el5}} = \sqrt{2} L_{1M} = \frac{i}{\sqrt{6\pi}} F_A \sum_i \vec{\sigma}_i \tau_{\pm i} \quad \text{GT}$$

$$M_{00}^5 = \frac{1}{\sqrt{4\pi}} F_1 \sum_i \tau_{\pm i} \quad \text{Fermi}$$

Spin-dependent excitations

● Gamow-Teller (1⁺):

● Spin-dipole (0⁻, 1⁻, 2⁻):

$$\begin{matrix} \vec{\sigma} & \tau \\ [\vec{\sigma} \times \vec{r}]^J & \tau_{\pm} \end{matrix}$$

Multipoles

1⁺: E₅1, M1, C₅1, L₅1

2⁻: E₅2, M2, C₅2, L₅2

1⁻: M₅1, E1, C1

0⁻: C₅0, L₅0

Probe of nuclear structure

(e,e) Frois & Papanicolas, Ann. Rev. Nucl. Part. 37 (1987)

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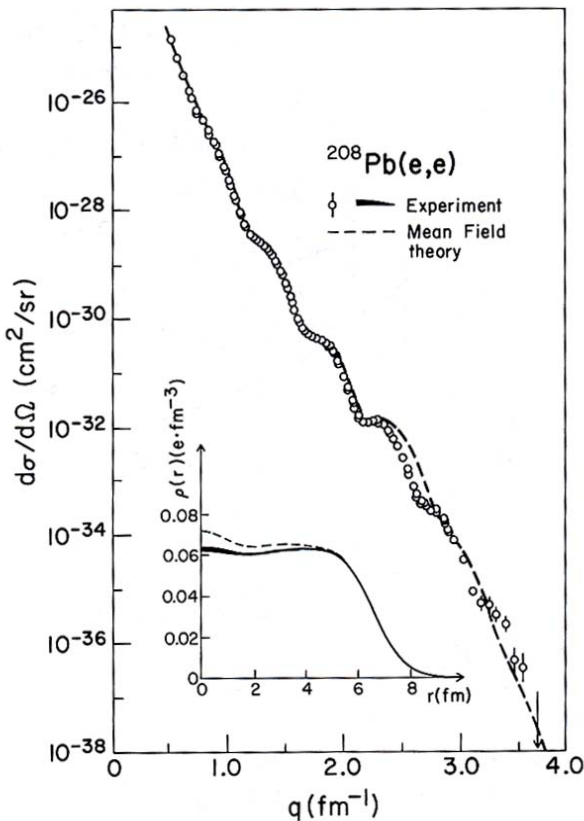


Figure 7 Elastic cross section from ²⁰⁸Pb (47). The dynamic range of the measurements allows the reconstruction of the charge density distribution (insert). The thickness of the line in $\rho(r)$ depicts the experimental uncertainty. The mean field result is that of Dechargé & Gogny (53).

Charge density

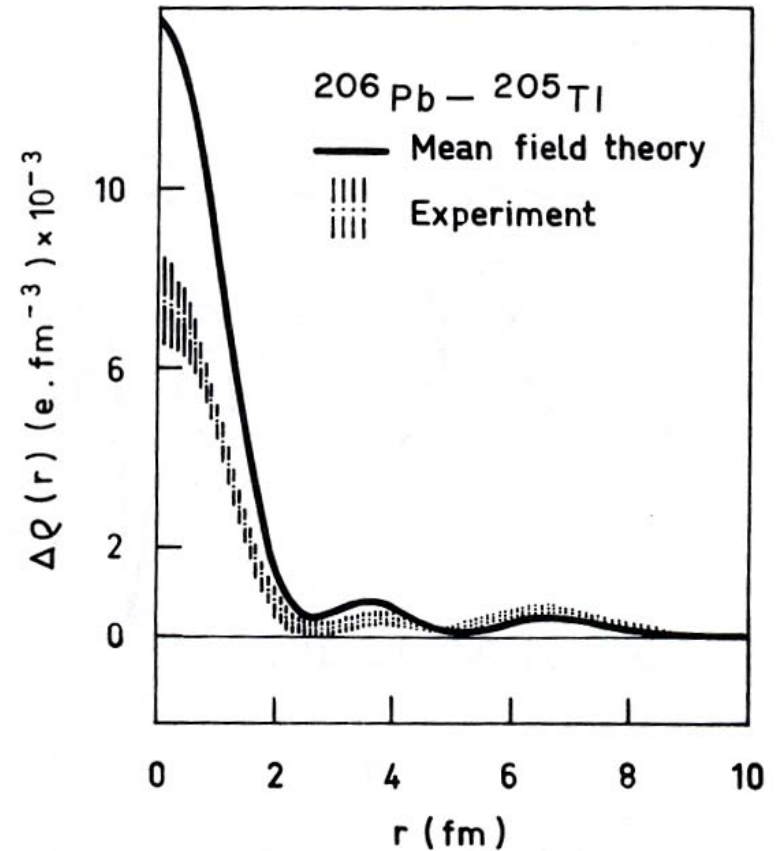


Figure 11 The charge density difference of ²⁰⁶Pb and ²⁰⁵Tl (63, 64).

Charge density of $2s_{1/2}$ orbit

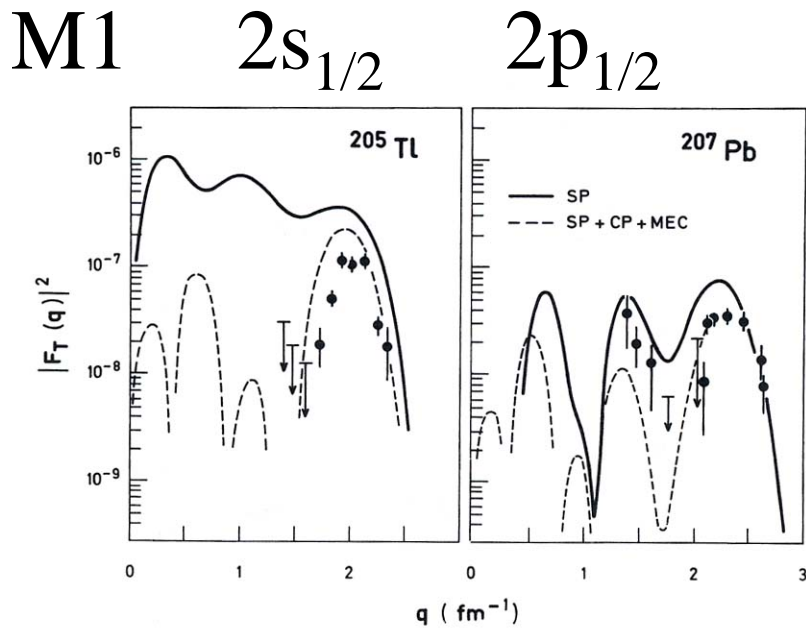


Figure 12 The elastic magnetic form factors of ^{207}Pb and ^{205}Tl (65). The solid curve depicts the single-particle predictions (53). The dashed curve includes in addition the effects of core

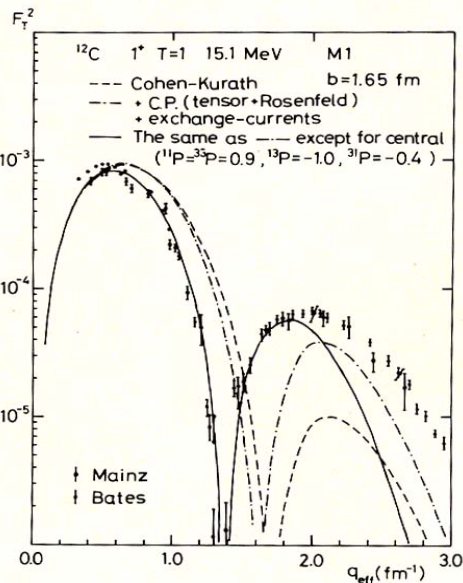
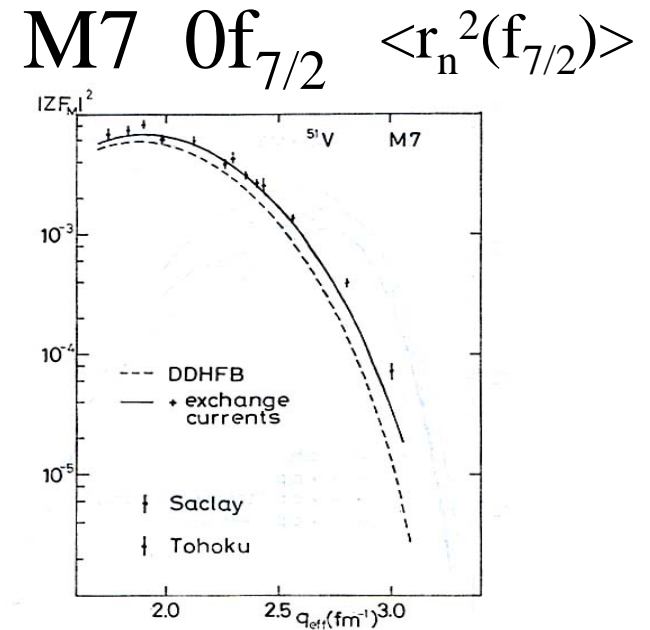


Fig. 2. M1 form factors in $^{12}\text{C}(e,e')^{12}\text{C}$ (1^+ , $T=1$, 15.1 MeV) scattering.

Core-Polarization

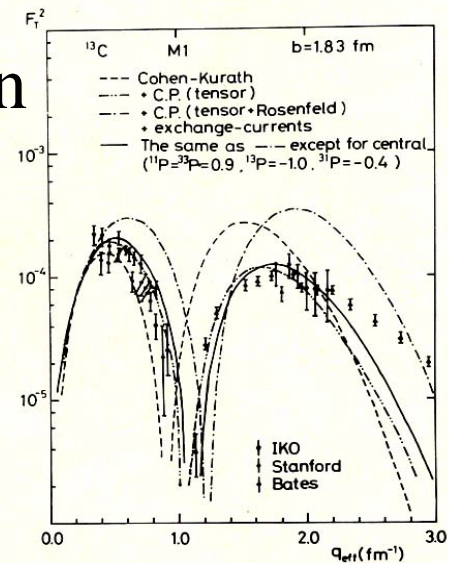


Fig. 1. M1 form factors in $^{13}\text{C}(e,e)^{13}\text{C}$ scattering.

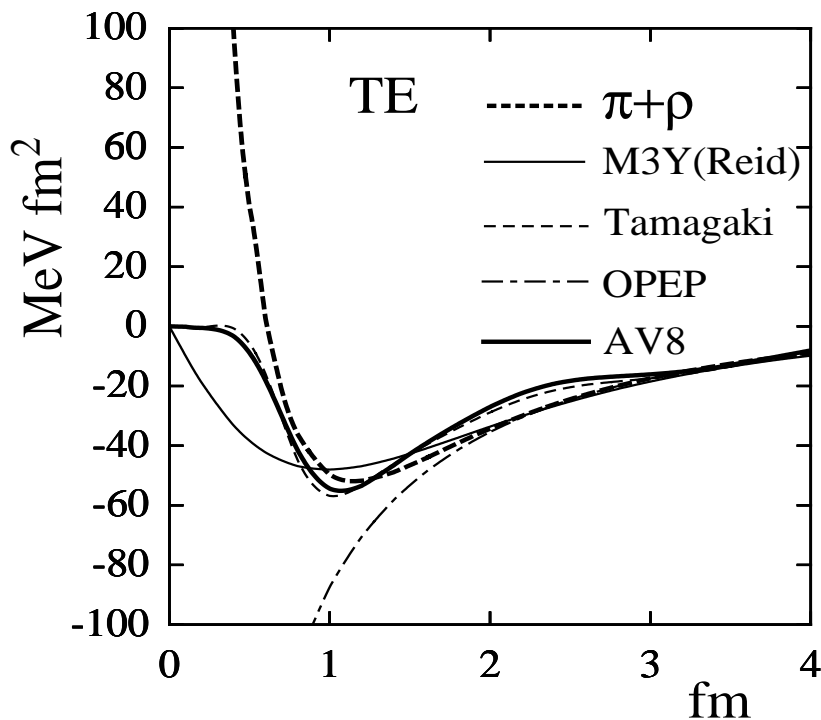
Shell-model interactions

$$V = V_C + V_{Tensor} + V_{LS}$$

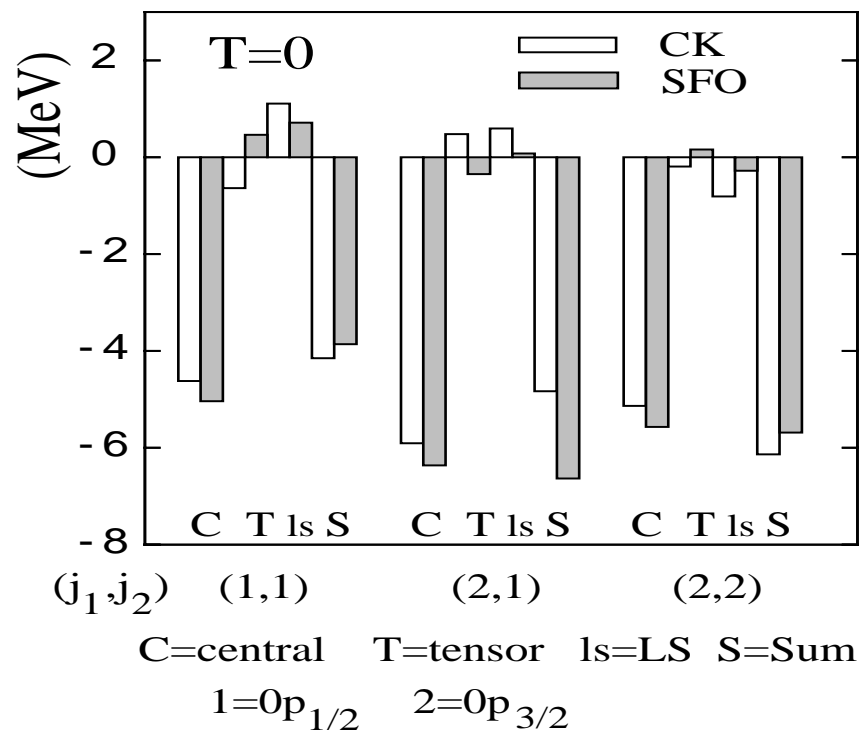
$$V_T = \vec{\tau}_1 \cdot \vec{\tau}_1 \{ 3(\vec{\sigma}_1 \cdot \hat{r})(\vec{\sigma}_2 \cdot \hat{r}) - \vec{\sigma}_1 \cdot \vec{\sigma}_2 \} Y(r)$$

Monopole terms in p-shell

$$\bar{V}_{j_1 j_2}^T = \sum_J (2J+1) \langle j_1 j_2 : JT | V | j_1 j_2 : JT \rangle / \sum_J (2J+1)$$



Monopole matrix elements



Effects of Tensor Force on Shell Evolution

Otsuka, Suzuki, Fujimoto, Grawe, Akaishi, PRL 69 (2005)

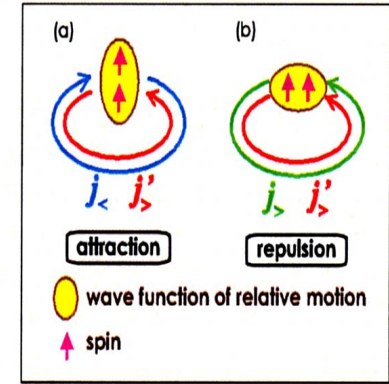
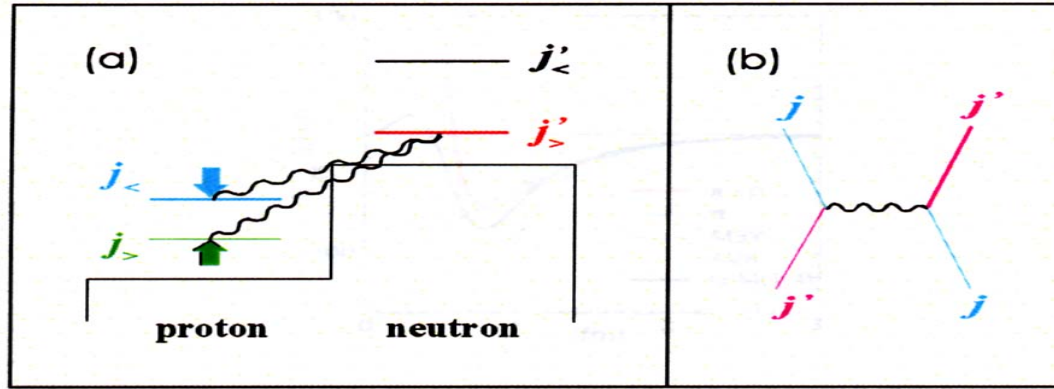
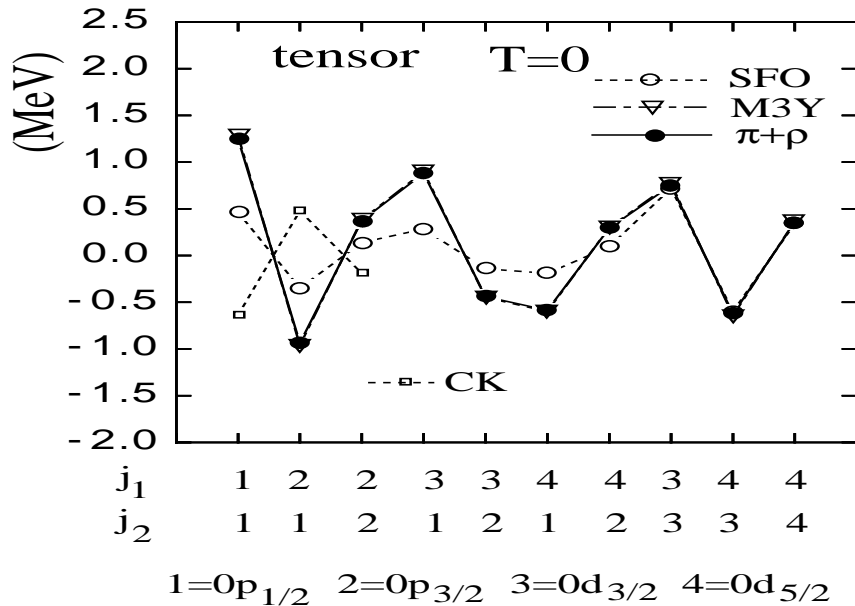


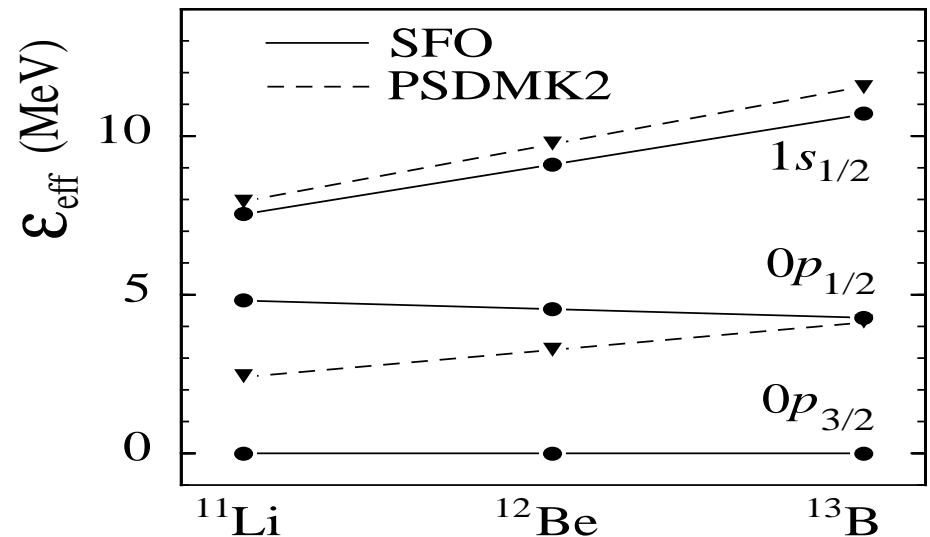
FIG. 1: (a) Schematic picture of the monopole interaction produced by the tensor force between a proton in $j_{<}$ and a neutron in $j_{>}$. (b) Exchange processes contributing to the monopole interaction of the tensor force.

FIG. 2: Intuitive picture of the tensor force acting two nucleons on orbits j and j' .

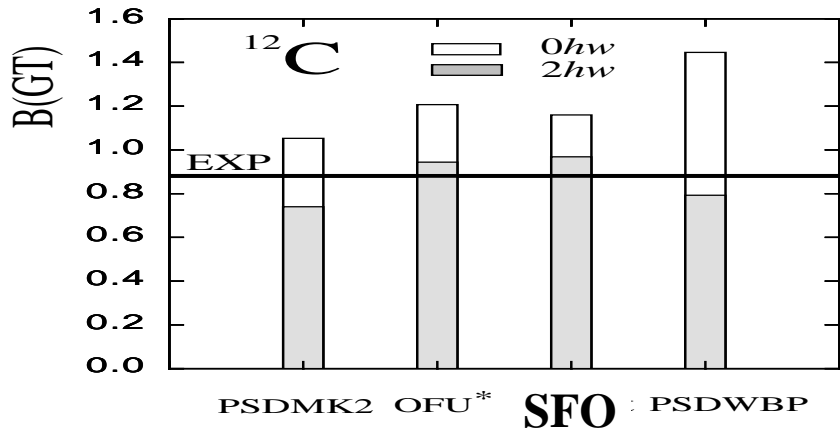
Monopole matrix elements



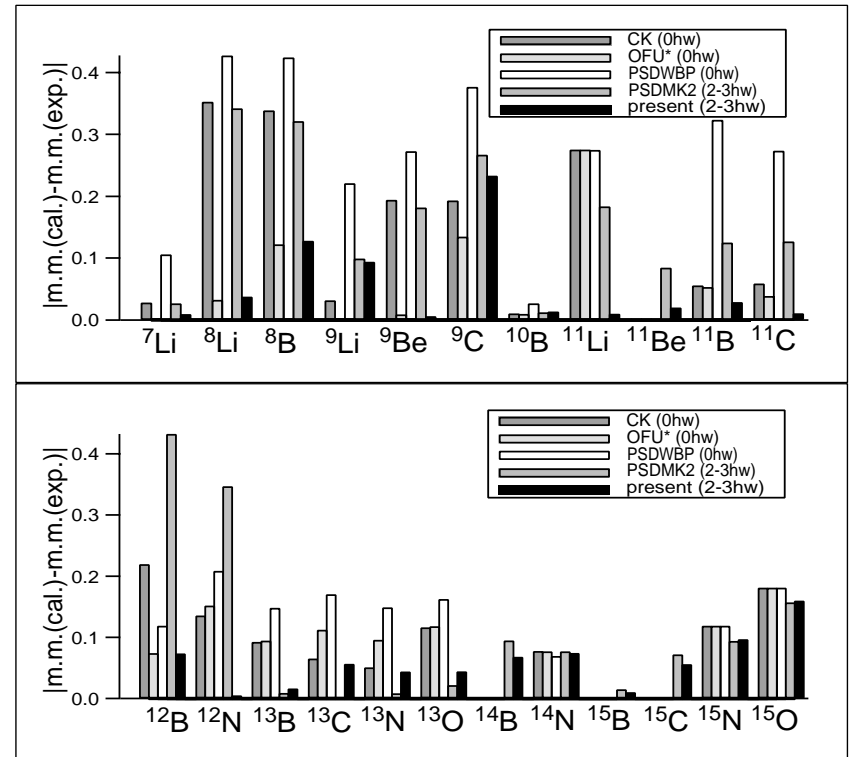
Effective Single-Particle Energy for N=8 Nuclei



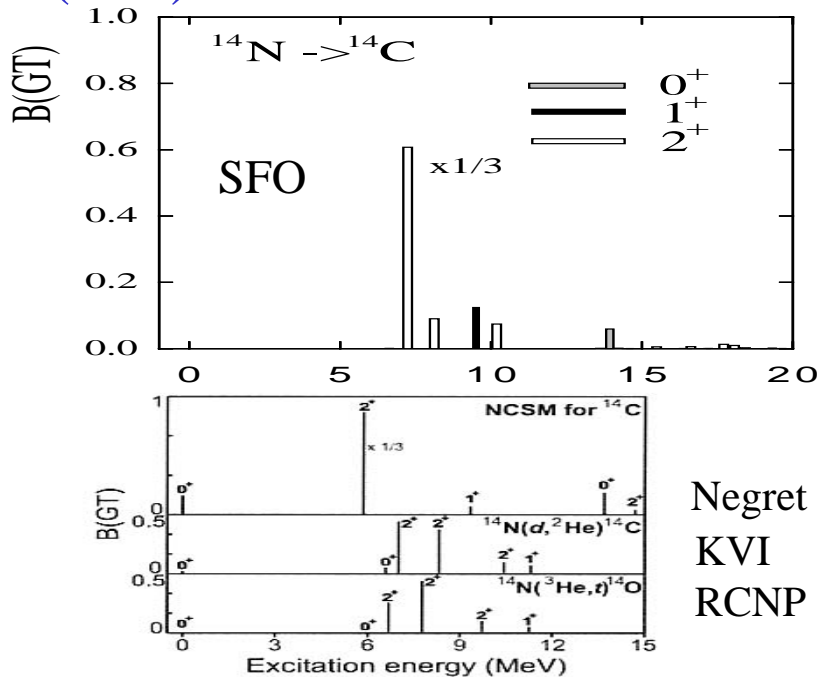
B(GT) values for $^{12}\text{C} \rightarrow ^{12}\text{N}$



Magnetic moments of p-shell nuclei



B(GT) values for $^{14}\text{N} \rightarrow ^{14}\text{C}$



present = SFO Suzuki, Fujimoto, Otsuka, PR C67 (2003)

Negret et al., PRL 97 (2006)
KVI
RCNP

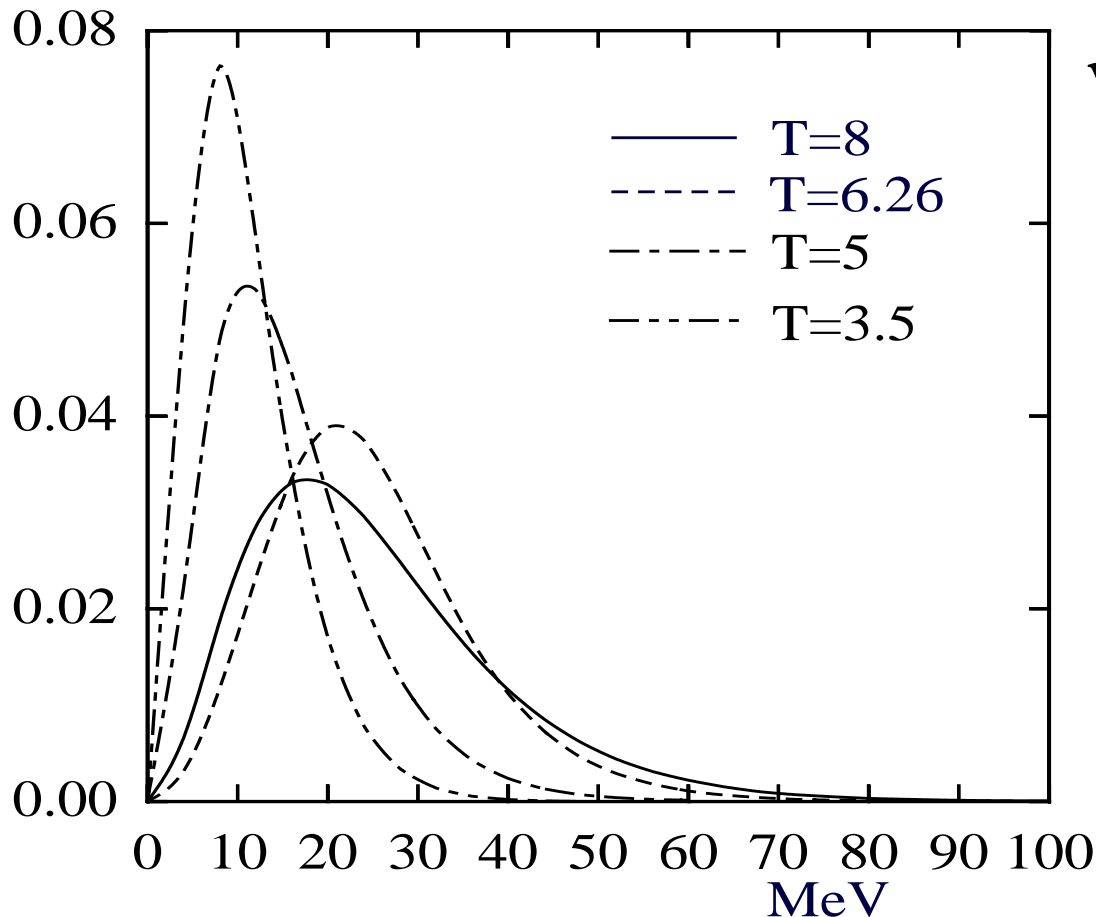
**SFO*: $g_A^{\text{eff}}/g_A = 0.95$
B(GT: ^{12}C) fitted to experiment**

FIG. 3. Experimental B(GT) distributions, compared to the theoretical result of Aroua *et al.* [14], where the B(GT) to the 2^+ state was scaled down by a factor of 3.

Supernovae ν Spectra

$\sigma \propto E^2$
 $\langle E \rangle$ & tail part

$$f(E) = N \frac{E^2 / T^3}{1 + \exp[E/T - \alpha]}$$



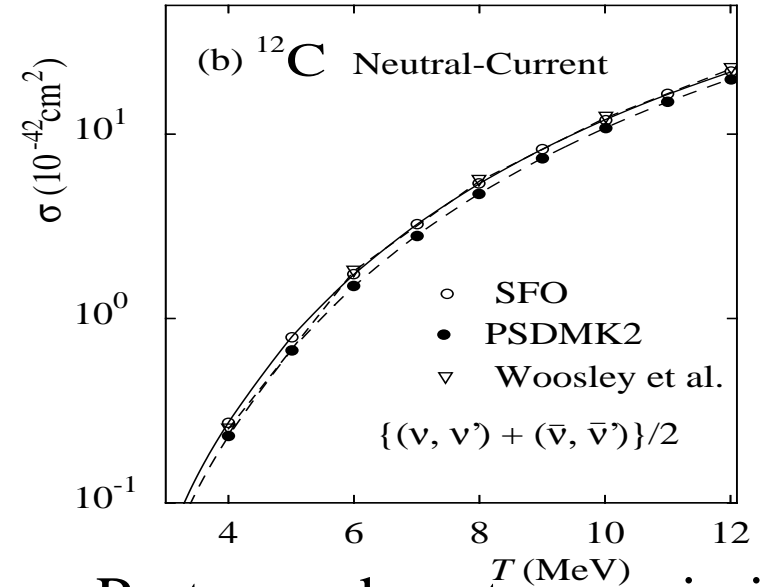
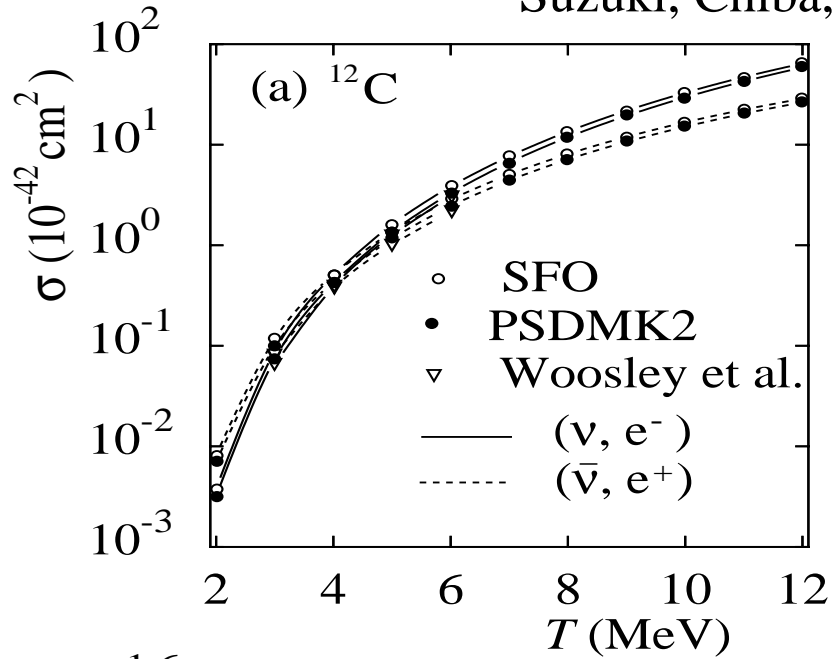
$\nu_{\mu}, \nu_{\tau} : \langle E \rangle = 25 \text{ MeV}$
 $(T, \alpha) = (8 \text{ MeV}, 0)$
 $(6.26 \text{ MeV}, 3)$

$\nu_e : \langle E \rangle = 11 \text{ MeV}$
 $(T, \alpha) = (3.5 \text{ MeV}, 0)$

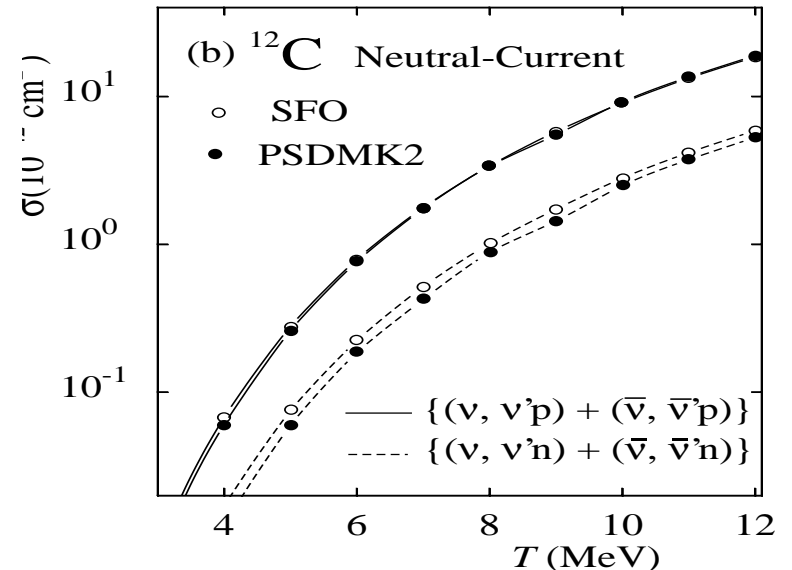
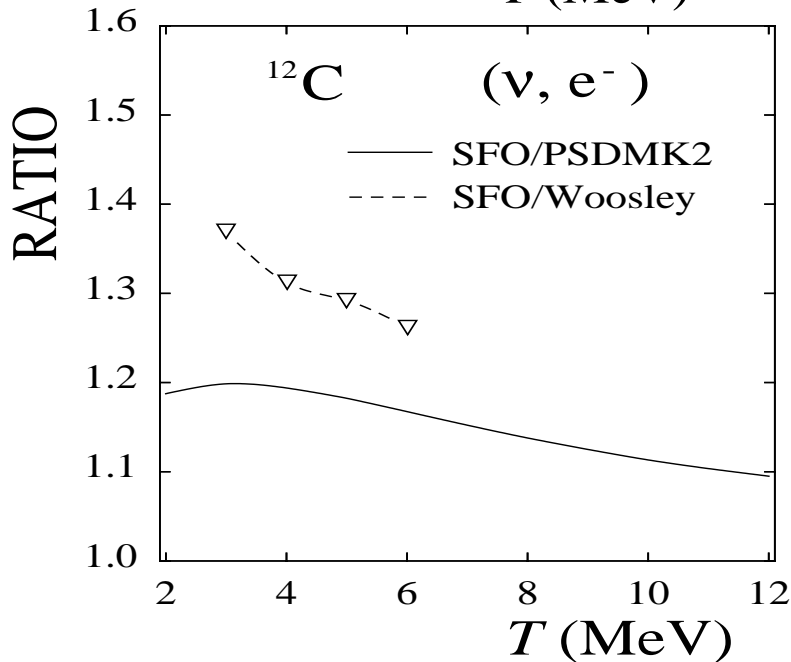
$\bar{\nu}_e : \langle E \rangle = 16 \text{ MeV}$
 $(T, \alpha) = (5 \text{ MeV}, 0)$

Cross sections for Supernova Neutrinos with temperature T

Suzuki, Chiba, Yoshida, Kajino, Otsuka, PR C74 (2006)

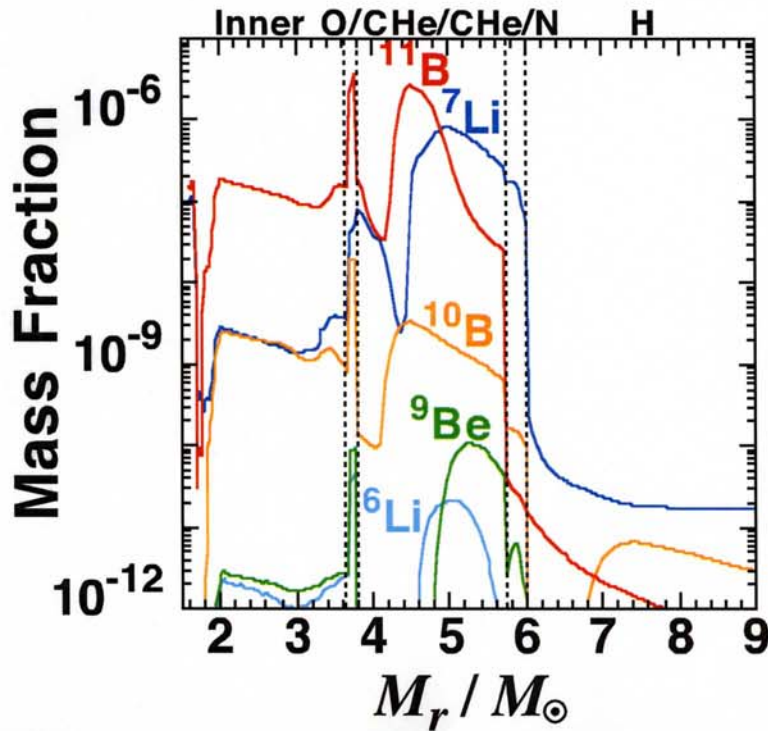


Proton and neutron emissions
BR: Hauser-Feshbach model

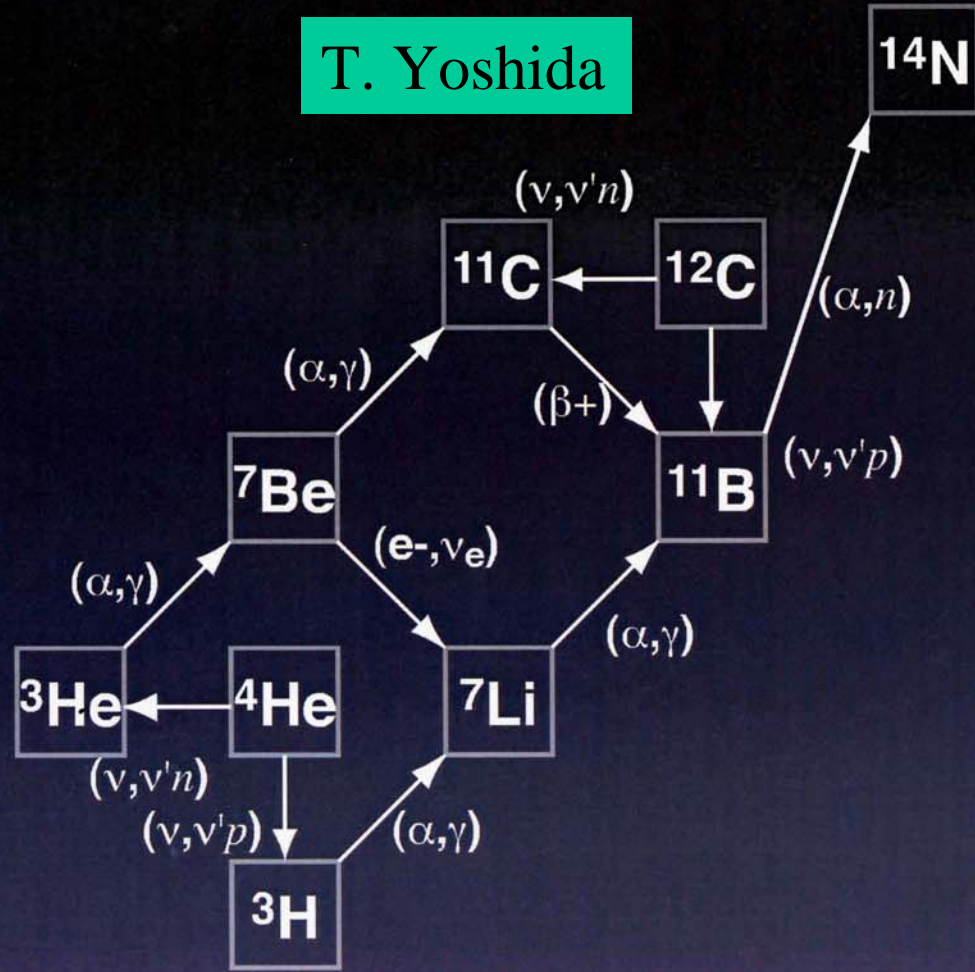


Light Element Abundances and Nucleosynthesis Processes

T. Yoshida



$E_\nu=300$ foe, $\tau_\nu = 3$ s, $T_{\nu\mu,\tau} = 8$ MeV

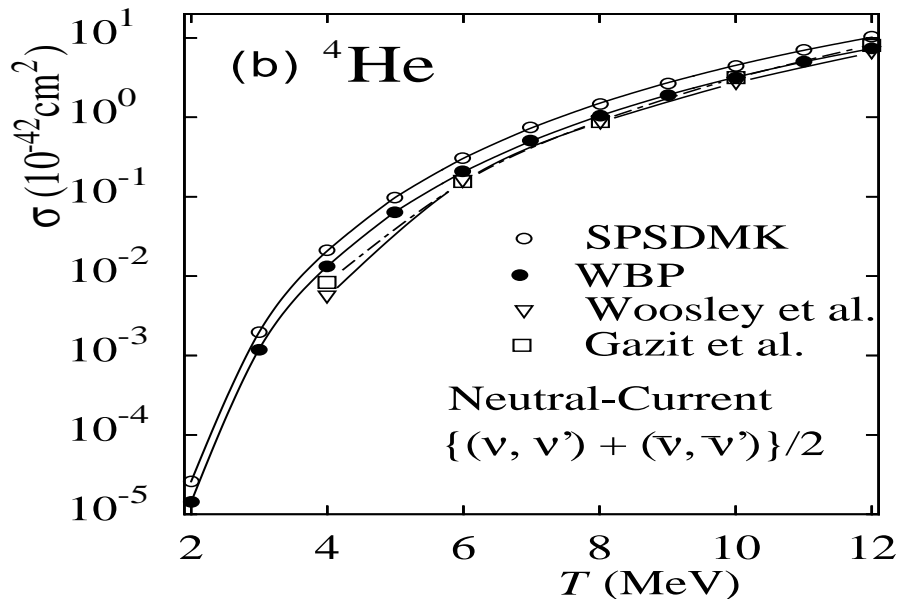
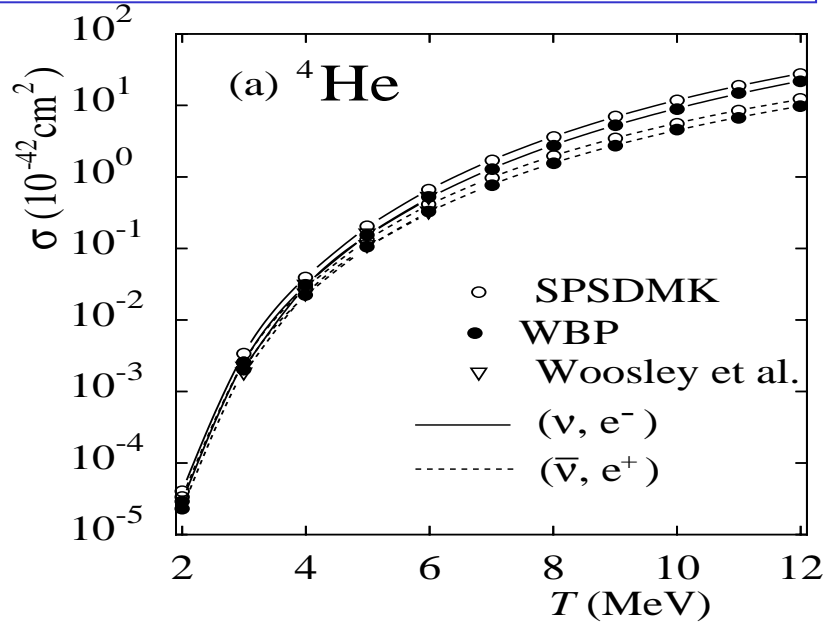


${}^7\text{Li}$ & ${}^{11}\text{B}$ production in He/C layer

${}^4\text{He}(\nu, \nu'p){}^3\text{H}$, ${}^4\text{He}(\nu, \nu'n){}^3\text{He}$, ${}^{12}\text{C}(\nu, \nu'p){}^{11}\text{B}$

${}^4\text{He}(\bar{\nu}_e, e^+n){}^3\text{H}$, ${}^4\text{He}(\nu_e, e^-n){}^3\text{He}$

ν - ^4He reaction cross sections



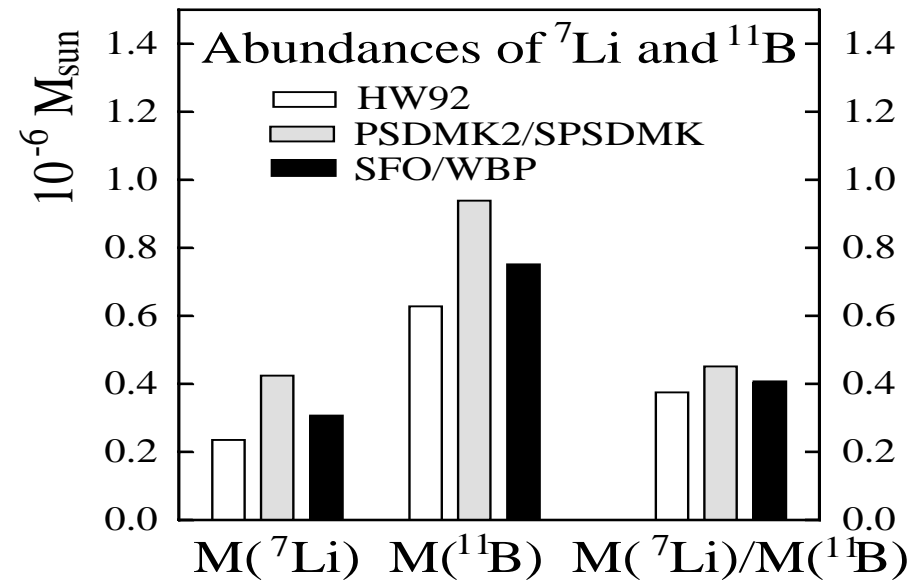
Abundances of ^7Li and ^{11}B produced in supernova explosion processes

$M=16.2 M_{\odot}$ (SN 1987A)

$T_{\nu_e} = 3.2 \text{ MeV}$, $T_{\bar{\nu}_e} = 5.0 \text{ MeV}$

$T_{\nu_{\mu}, \nu_{\tau}} = 6.0 \text{ MeV}$

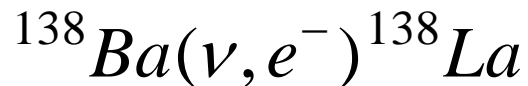
$(\nu, \nu' p), (\nu, \nu' n)$ $\nu = \nu_{\mu, \tau}, \bar{\nu}_{\mu, \tau}$



cf. Woosley-Haxton: Sussex potential by Elliott et al.

Nucleosynthesis through neutrino-induced reactions

- Production of rare elements by ν - reactions



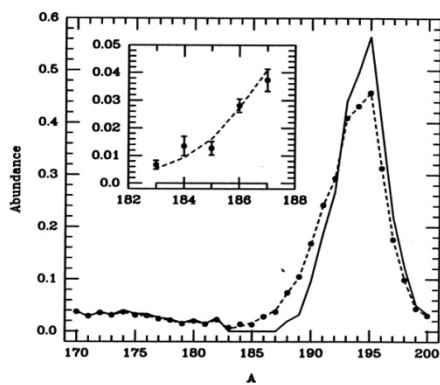
Calculation by Heger et al.

GT exp. RCNP ($^3\text{He}, t$)

More GT strength than RPA
(Heger et al.)

Byelikov et al., PRL 98 (2007)

- Role of ν in r-process nucleosynthesis



N=82, 126 regions
With ν -induced n
emission \rightarrow solar
abundances

Figure 5. Effects of neutrino-induced neutron emission for the region near the abundance peak at $A \sim 195$ [48, 50]. The abundances before and after neutrino-induced neutron emission (following freeze-out of the r-process) are given by the solid and dashed curves, respectively. The filled circles (some with error bars) give the solar r-process abundances. The region with $A = 183-187$ is highlighted in the inset.

Qian, Haxton, Langanke,
Vogel, PR C55 (1997)

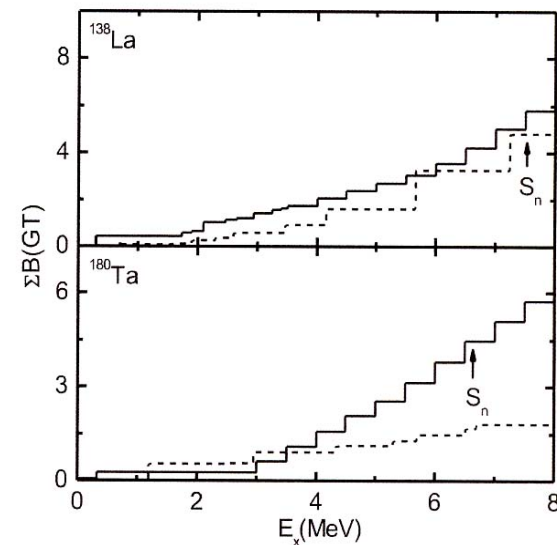


FIG. 4. Summed $B(\text{GT})$ strength in ^{138}La and ^{180}Ta as a function of excitation energy. The neutron emission thresholds (S_n) are indicated by arrows. Solid lines: present experiment. Dashed lines: RPA calculations from Ref. [2].

Probe of Weak Neutral Current

$$G_1^S (\langle N | \bar{s} \gamma_\mu \gamma_5 s | N \rangle) \text{ via } A(\nu, \nu') A(T=0)$$

$$\frac{d\sigma}{d\Omega} \propto |V^S + A^S - 2 \sin^2 \theta_W J^\gamma|^2 \approx |A^S|^2 \propto (G_1^S)^2$$

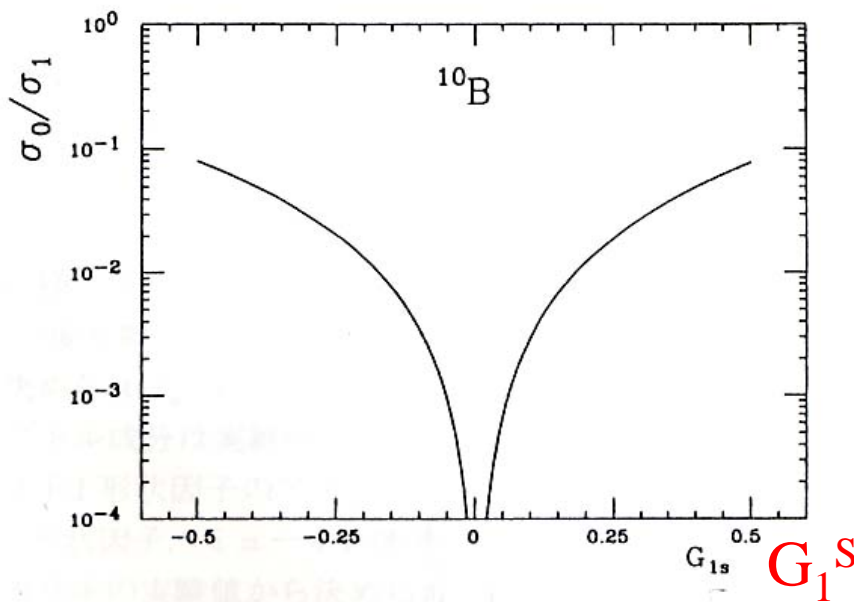


Fig. 5 Ratio of the total cross section of $^{10}\text{B}(\nu, \nu')$ $^{10}\text{B}(2^+, T=0, 3.59\text{MeV})$ over that of $^{10}\text{B}(\nu, \nu')$ $^{10}\text{B}(2^+, T=1, 5.16\text{MeV})$.

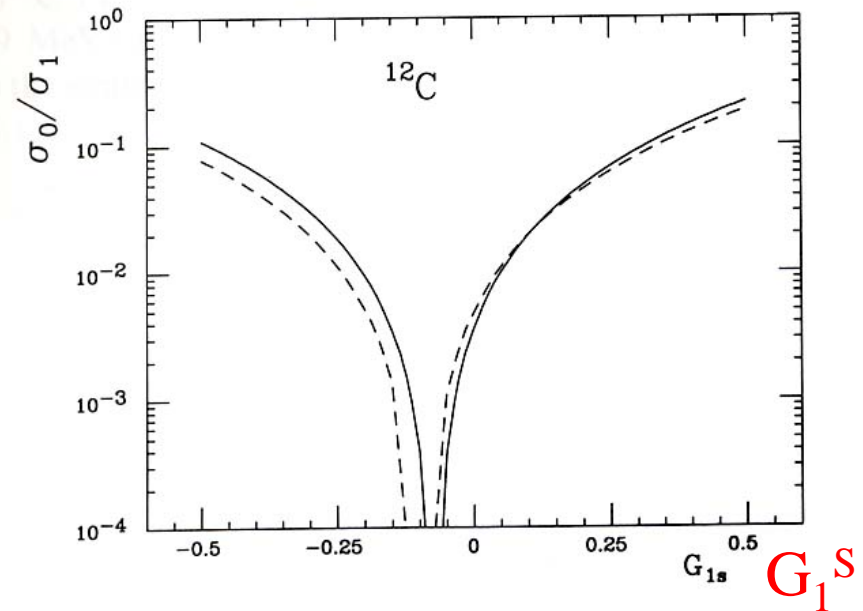


Fig. 1. Ratios of the total cross section of $^{12}\text{C}(\nu, \nu')^{12}\text{C}(1^+, 12.71\text{MeV})$ over that of $^{12}\text{C}(\nu, \nu')^{12}\text{C}(1^+, 15.11\text{MeV})$. The solid curve is obtained by using the Cohen-Kurath wave function, while the dashed curve includes the effects of the admixtures of $p^6(sd)^2$ components.

Polarized electron scattering

$$A(\vec{e}, e)A$$

$$A = \frac{d\sigma_{\rightarrow} - d\sigma_{\leftarrow}}{d\sigma_{\rightarrow} + d\sigma_{\leftarrow}} \propto \text{Re} \left\{ \begin{array}{c} e' \quad N \quad e' \quad N \dagger \\ \gamma^\mu \quad \frac{\gamma}{4\pi\alpha} \quad \gamma^\mu \otimes \gamma^\mu (a+b\gamma^5) \quad Z^0 \quad \gamma^\mu, \gamma^\mu\gamma^5 \\ \uparrow \downarrow \quad \uparrow \downarrow \quad \uparrow \downarrow \quad \uparrow \downarrow \\ e \quad N \quad e \quad N \end{array} \right\}$$

$$R - L = \frac{1}{2}(1 - \gamma_5) - \frac{1}{2}(1 + \gamma_5) = -\gamma_5$$

$$A \propto \text{Re} \left\{ \text{Tr} [k_2 \gamma_\mu \gamma_5 k_1 \gamma_\nu (1 - 4 \sin^2 \theta_W + \gamma_5) J_\mu^\gamma J_\nu^{W*}] \right\}$$

$$\approx \text{Re} \left\{ (1 - 4 \sin^2 \theta_W) J_\mu^\gamma J_\mu^{A*} + J_\mu^\gamma J_\mu^{V*} \right\}$$

$$\approx J_\mu^\gamma \left\{ 0.08 \langle N | \bar{s} \gamma_\mu \gamma_5 s | N \rangle + \langle N | \bar{s} \gamma_\mu s | N \rangle \right\}$$

$$F_2^S$$

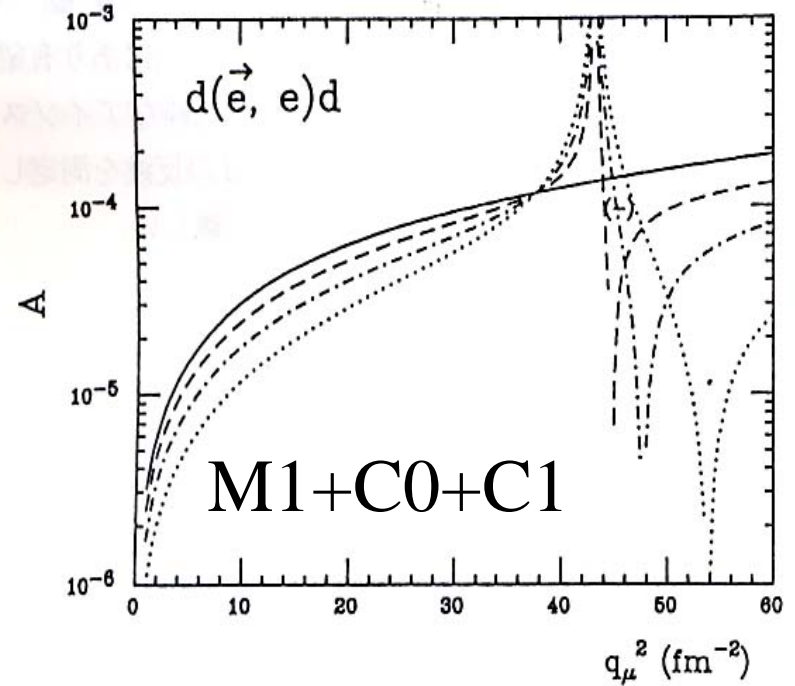
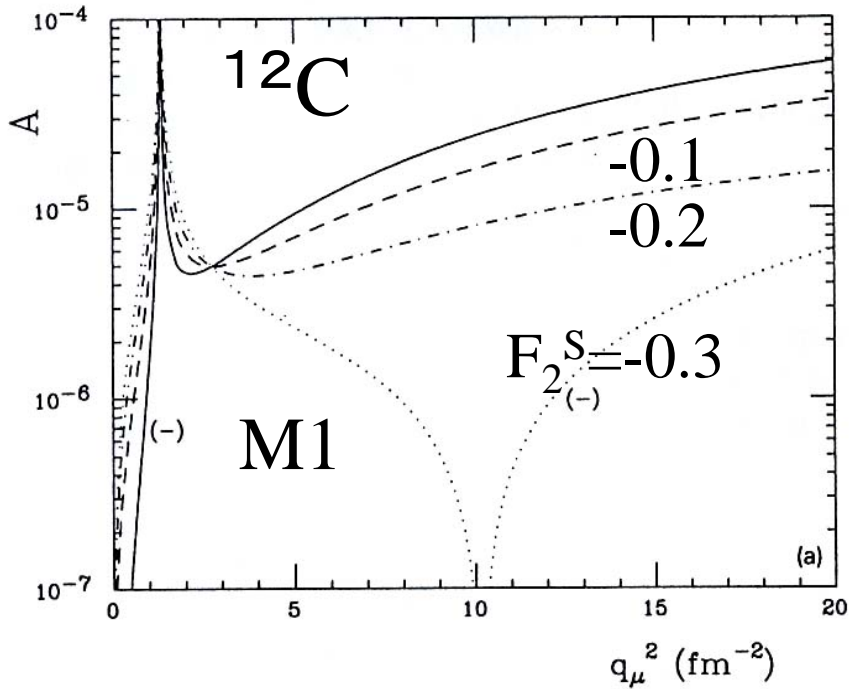
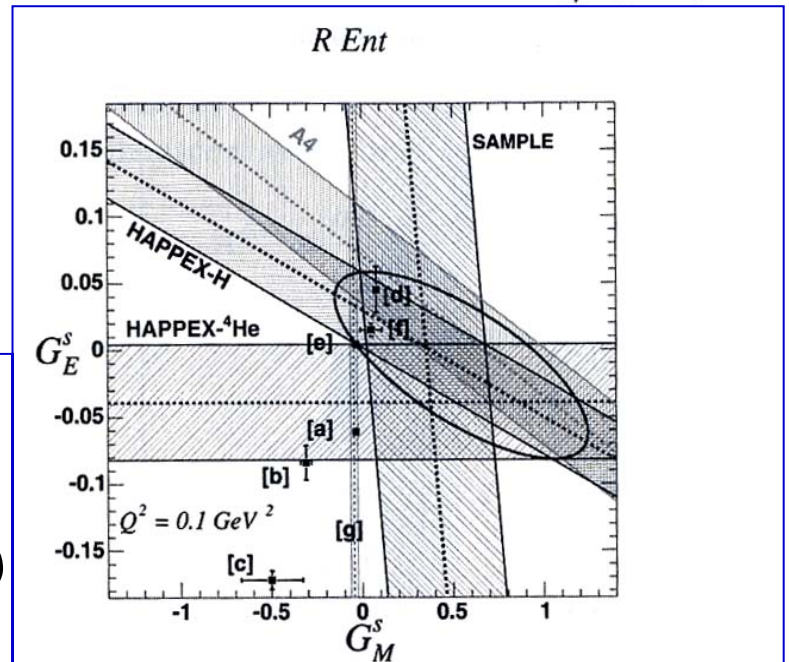


Fig. 3a

Calculated asymmetries of the $^{12}\text{C}(e^-, e')^{12}\text{C}(1^+, 12.71\text{MeV})$ reaction at $E_e = 1\text{ GeV}$. Adopted values of $F_2^s(0)$ are 0.0 (solid), -0.1 (dashed), -0.2 (dash-dotted) and -0.3 (dotted curve). Vector form factors of the nucleon are taken from ref. 26). $G_1^s(0)$ is fixed to be 0.0.

SAMPLE etc
Ent, in *EM Int. and
Hadron Structure (2007)*



$A(\vec{e}, e)A$ Isoscalar Coulomb scatt. $0^+ \rightarrow 0^+$

$$A \approx \frac{G q_\mu^2}{4\pi\alpha\sqrt{2}} \frac{\rho_n(q)}{\rho_p(q)}$$

$$(G_E^S = F_1^S - \frac{q_\mu^2}{4m^2} F_2^S \quad (F_1^S(0) = 0))$$

$$A_S \equiv \pi\alpha\sqrt{2}A / (G q_\mu^2 \sin^2 \theta_w)$$

$$\approx \frac{1}{4\sin^2 \theta_w} \frac{N}{Z} \left\{ 1 - \frac{q^2}{6} (\langle r_n^2 \rangle - \langle r_p^2 \rangle) \right\}$$

Ca isotopes

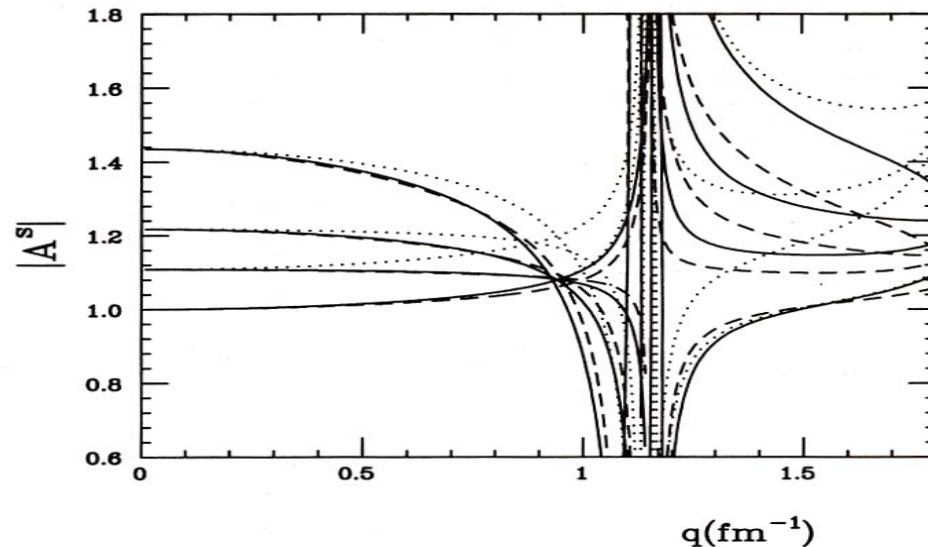


FIG. 1. Calculated asymmetries for Ca isotopes. Solid and dashed curves are obtained by using the SGII-HF and HF-T wave functions, respectively. The dotted curve is obtained by using the modified HF-T wave functions which reproduce the experimental values of r_n and r_p .

Beyond standard model

Neutrino oscillations

$$\alpha = e, \mu, \tau$$

$$|\nu_\alpha\rangle = \sum_a U_{\alpha a}^* |\nu_a\rangle \quad a=1,2,3$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{13}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{13}} & 0 & c_{13} \end{pmatrix}$$

$$\times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$c_{ij} = \cos \theta_{ij}, \quad s_{ij} = \sin \theta_{ij}$$

Present status of parametrization

$$1.5 \times 10^{-3} < |\Delta m_{23}^2| < 3.4 \times 10^{-3} \text{ eV}^2$$

$$\theta_{23} = 45 \pm 8^\circ \text{ (90\%)} \quad (\text{K2K})$$

$$|\Delta m_{12}^2| = 8.2_{-0.5}^{+0.6} \times 10^{-5} \text{ eV}^2$$

$$\theta_{12} = 32.3_{-2.4}^{+2.7} (1\sigma) \quad (\text{KamLAND, SN})$$

$$\sin^2 2\theta_{13} < 0.1 \quad (\text{Chooz})$$

ν oscillations in supernova explosion

→ Possible constraint on lower limit of θ_{13}