

Some basic properties and excitations in nuclei  
—  $\beta$ -stable nuclei vs. unstable nuclei.

1) Shape and shell-structure.

One-body potentials and occupied levels;

Effect of neutron excess on density and potential;

Change of shell-structure as  $E_\nu = -10 \text{ MeV} \rightarrow 0 \text{ MeV}$ ;

Deformation (Jahn-Teller effect), and new shell-structure  
at large deformation);

One-particle motion in deformed potentials - Nilsson diagram.

2) Collective modes.

Giant resonances;

Low-energy threshold strength in drip line nuclei;

Model-independent sum-rules.

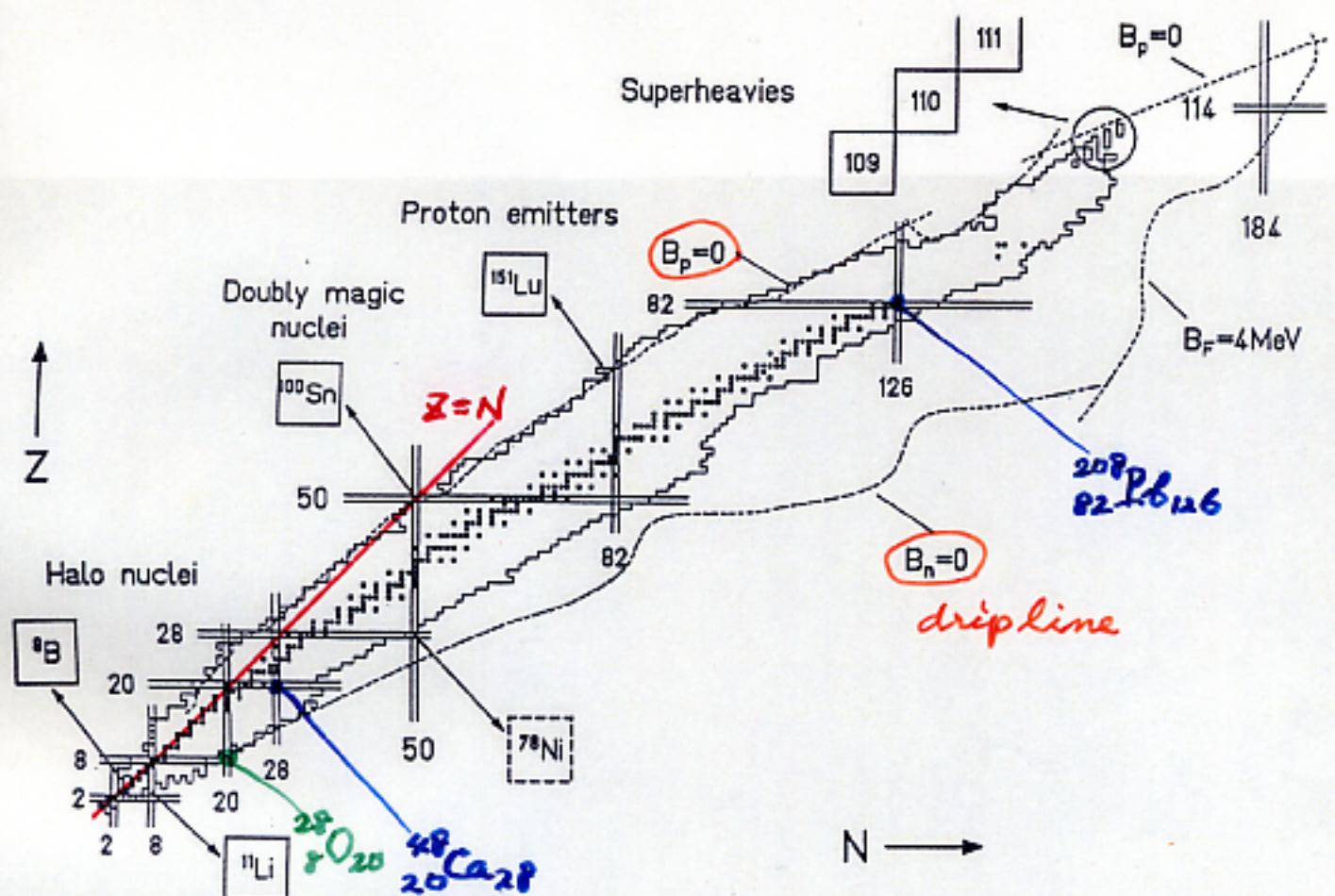


Figure 1 Chart of nuclei with nuclear driplines (dripline) and selected landmarks of exotic nuclear structure. The full lines indicate the present limits of discovered nuclei.

## § Shape and Shell-structure

$\beta$ -stable nuclei :

one-body potential  $\sim$  harmonic-oscillator potential  
+  $(\vec{l} \cdot \vec{s})$  potential  
+ surface effect.

separation energy of nucleons,  $S_n \sim S_p \sim 7-10$  MeV.

magic numbers ( $N$  or  $Z$ ) = 2, 8, 20, 28, 50, 82, 126,

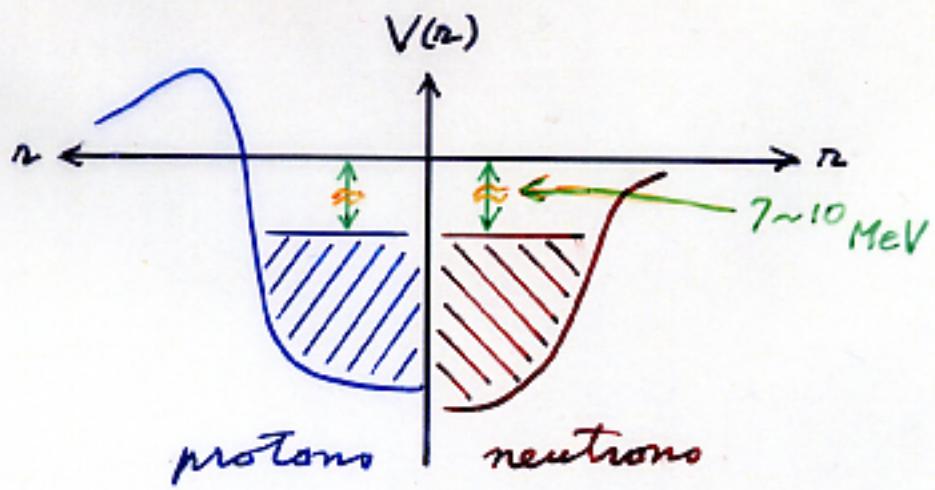
ground states of nuclei with magic numbers are spherical.

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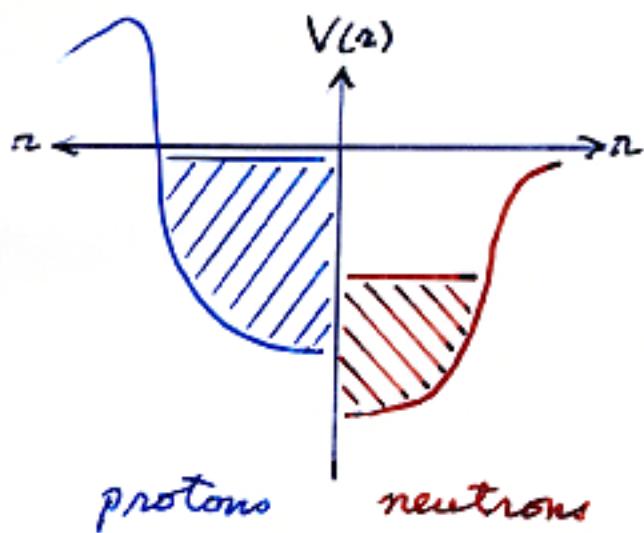
Hartree-Fock approximation (= mean-field approximation of nuclear many-body problems)

$$H = \underbrace{\sum_i t_i}_{\text{kinetic energy}} + \underbrace{\sum_{i < j} v_{ij}}_{\text{two-body interaction}} \rightarrow \sum_i t_i + \underbrace{\sum_i V_i}_{\text{one-body (HF) potential calculated self-consistently.}}$$

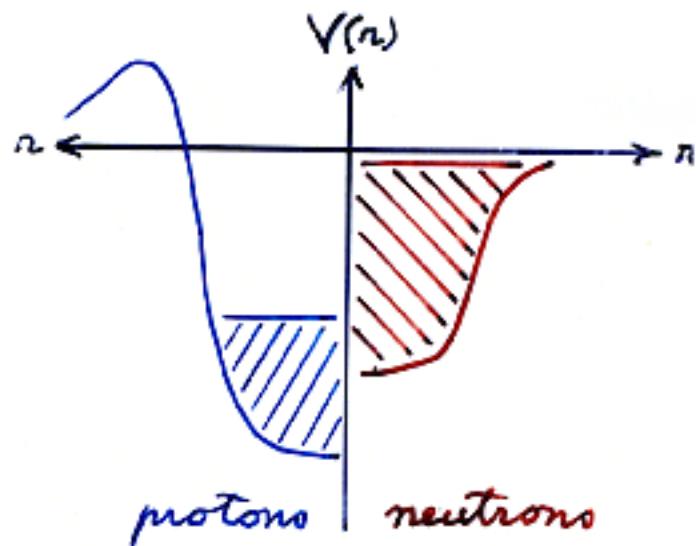
### $\beta$ -stable nuclei

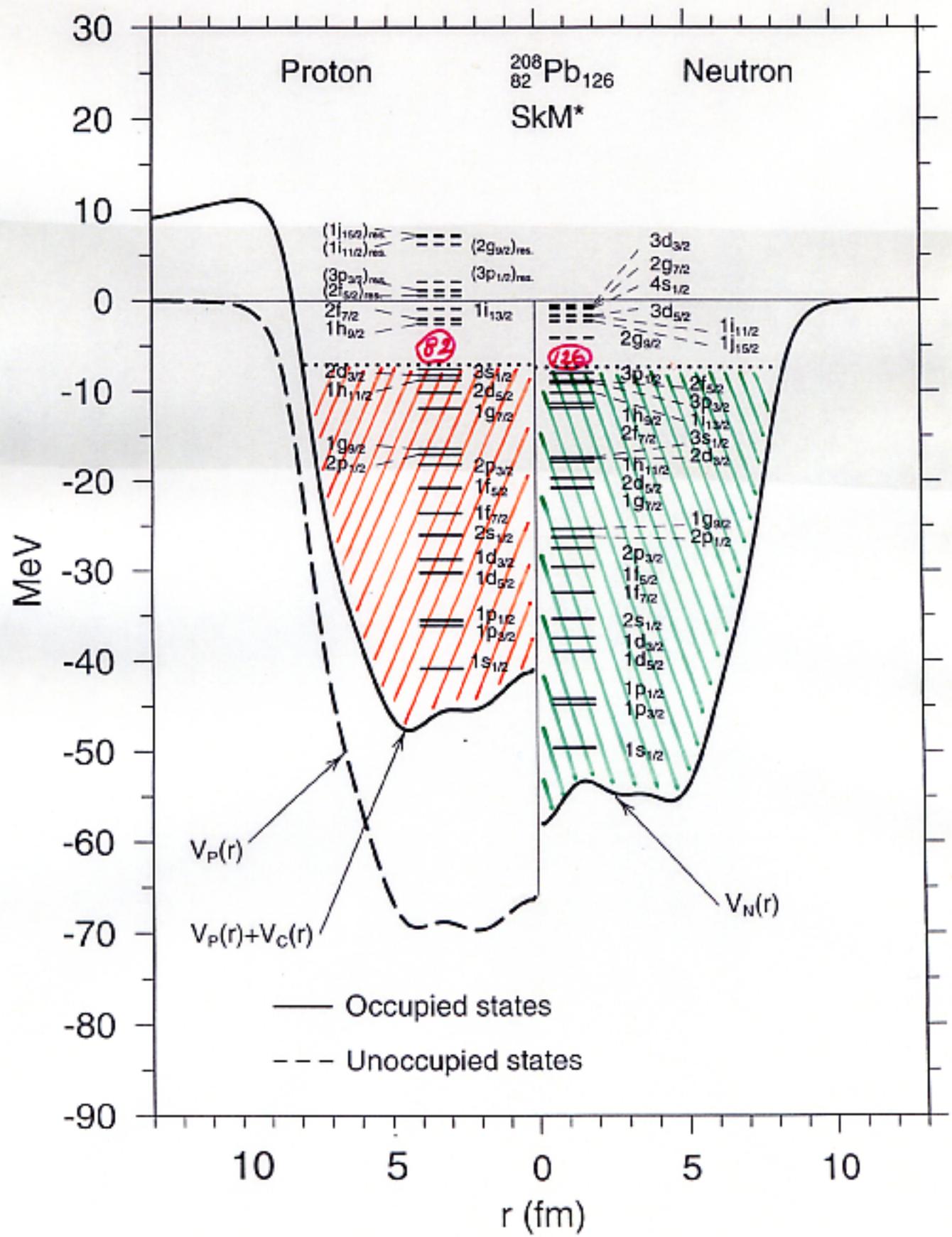


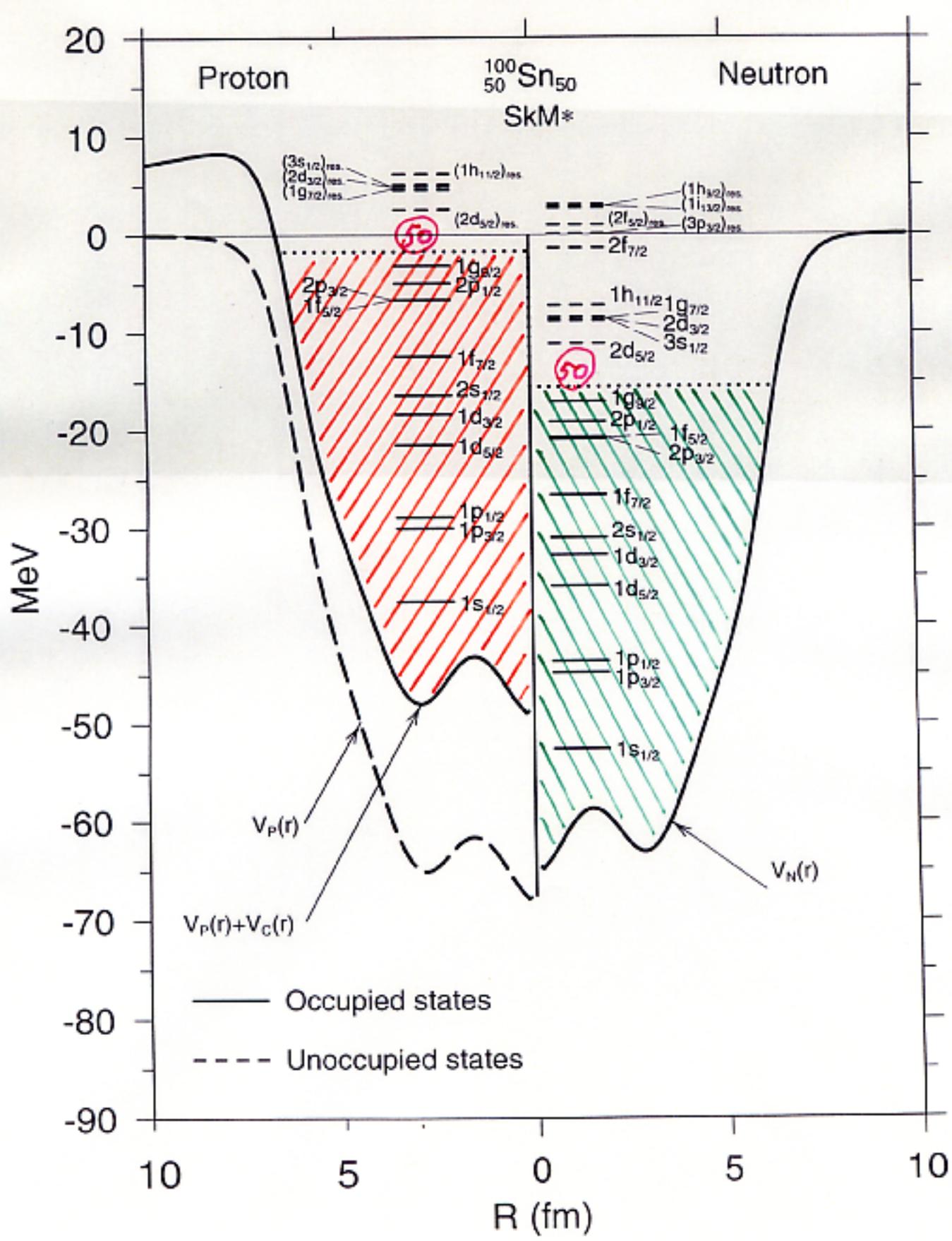
### $p$ drip line nuclei

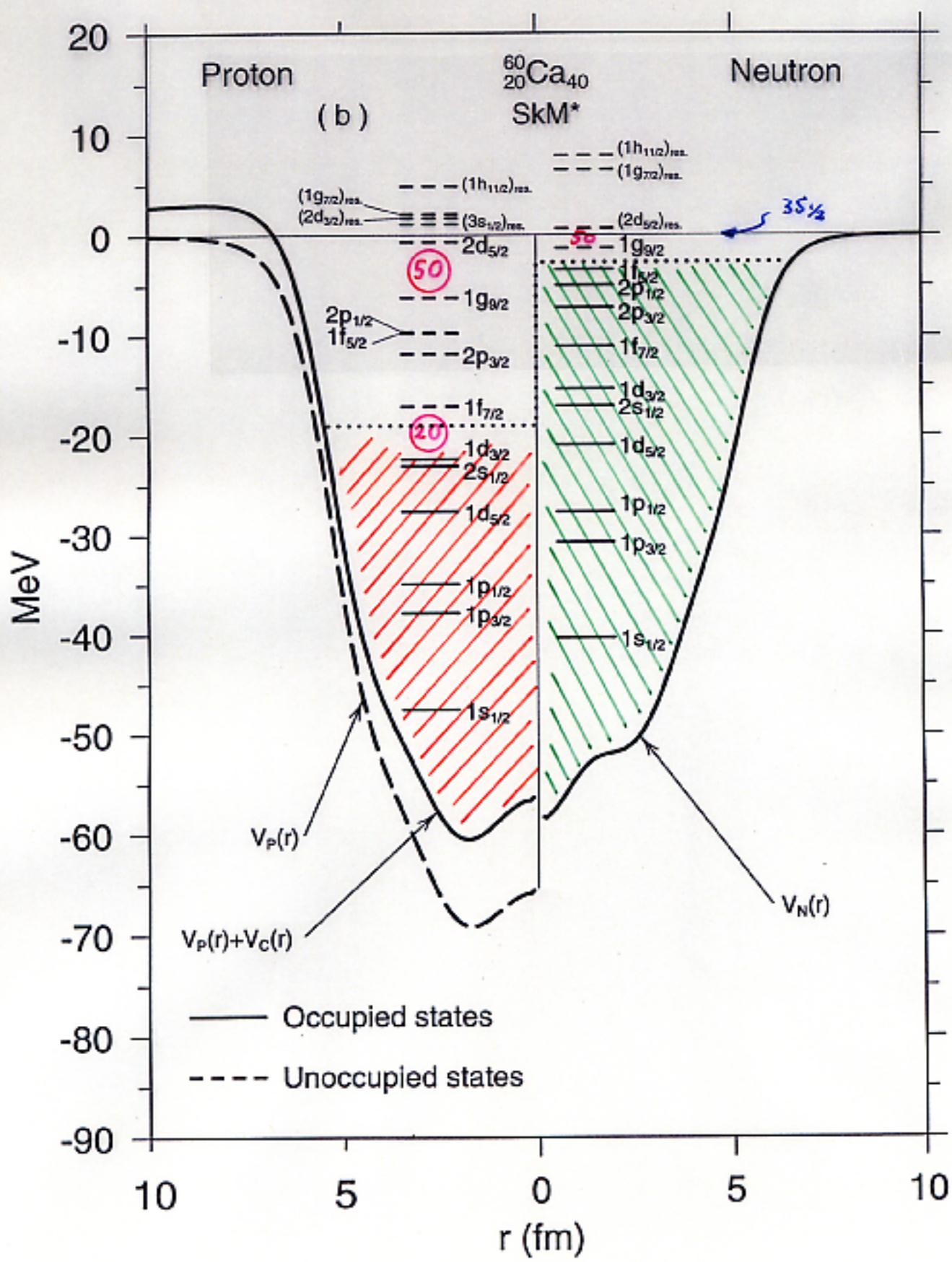


### $n$ drip line nuclei







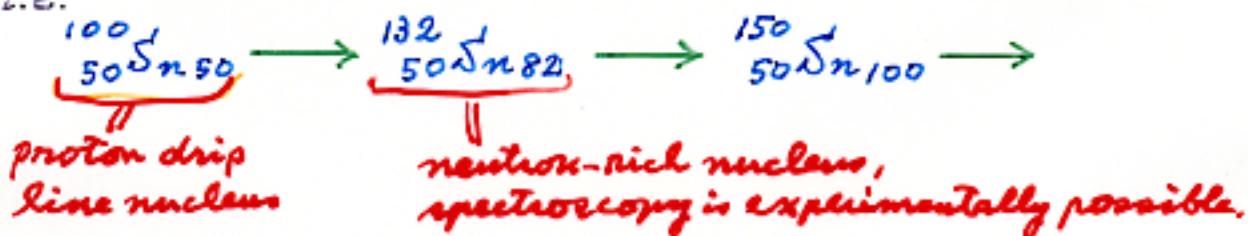


## Effect of neutron excess

e.g.

Keeping  $Z=50$ ,  $N=50 \rightarrow$  larger

i.e.



$$\left\{ \begin{array}{l} V_p(r) \rightarrow \text{deeper}, \quad S_p(r=0) \rightarrow \text{smaller}, \quad R_p \rightarrow \text{larger}, \\ S_n(r=0) \rightarrow \text{larger}, \quad R_n \rightarrow \text{larger}, \quad a_n \rightarrow \text{larger}, \end{array} \right.$$

Dobaczewski, Hamamoto, Nazarewicz and Sheikh,  
Acta Phys. Pol. B25 (1994) 541.

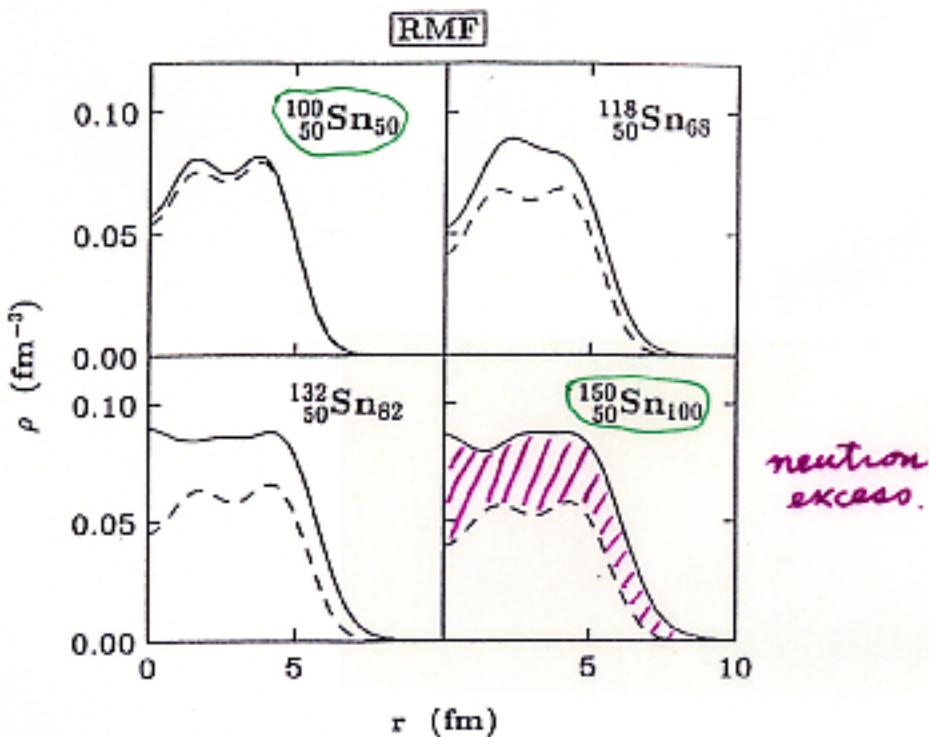
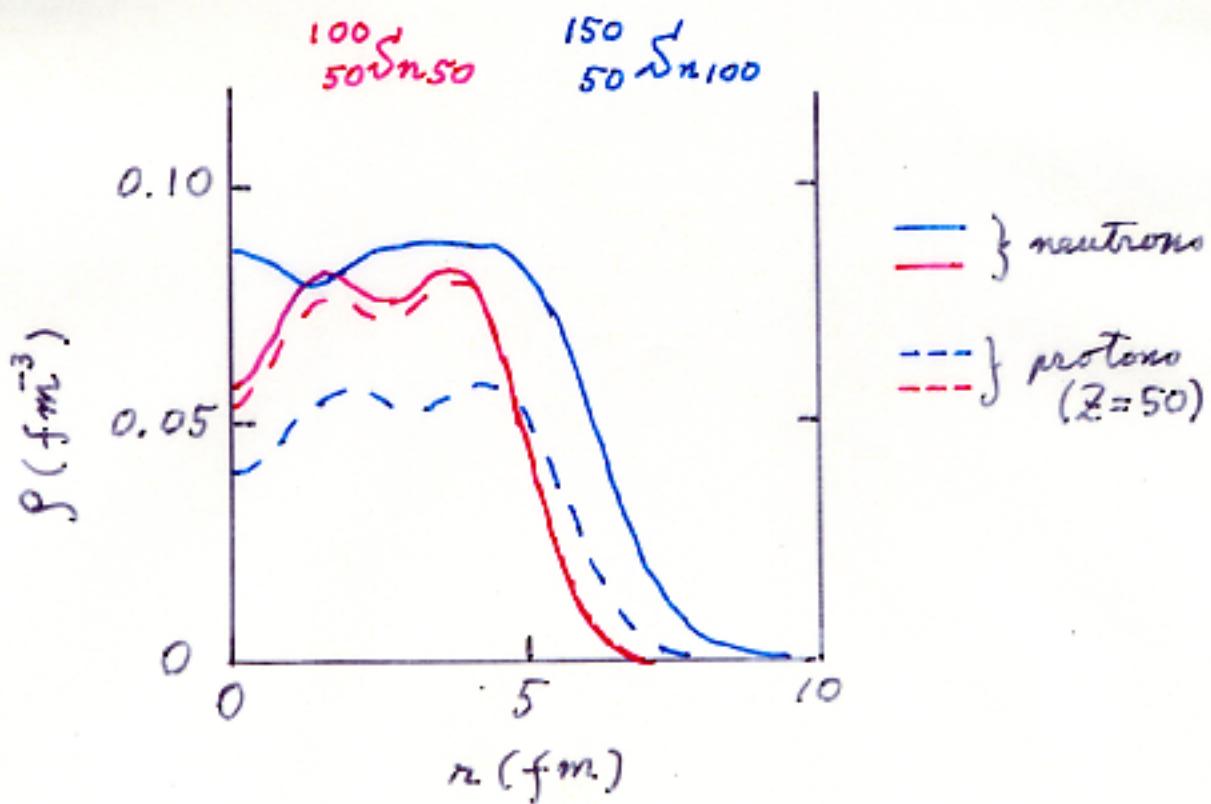
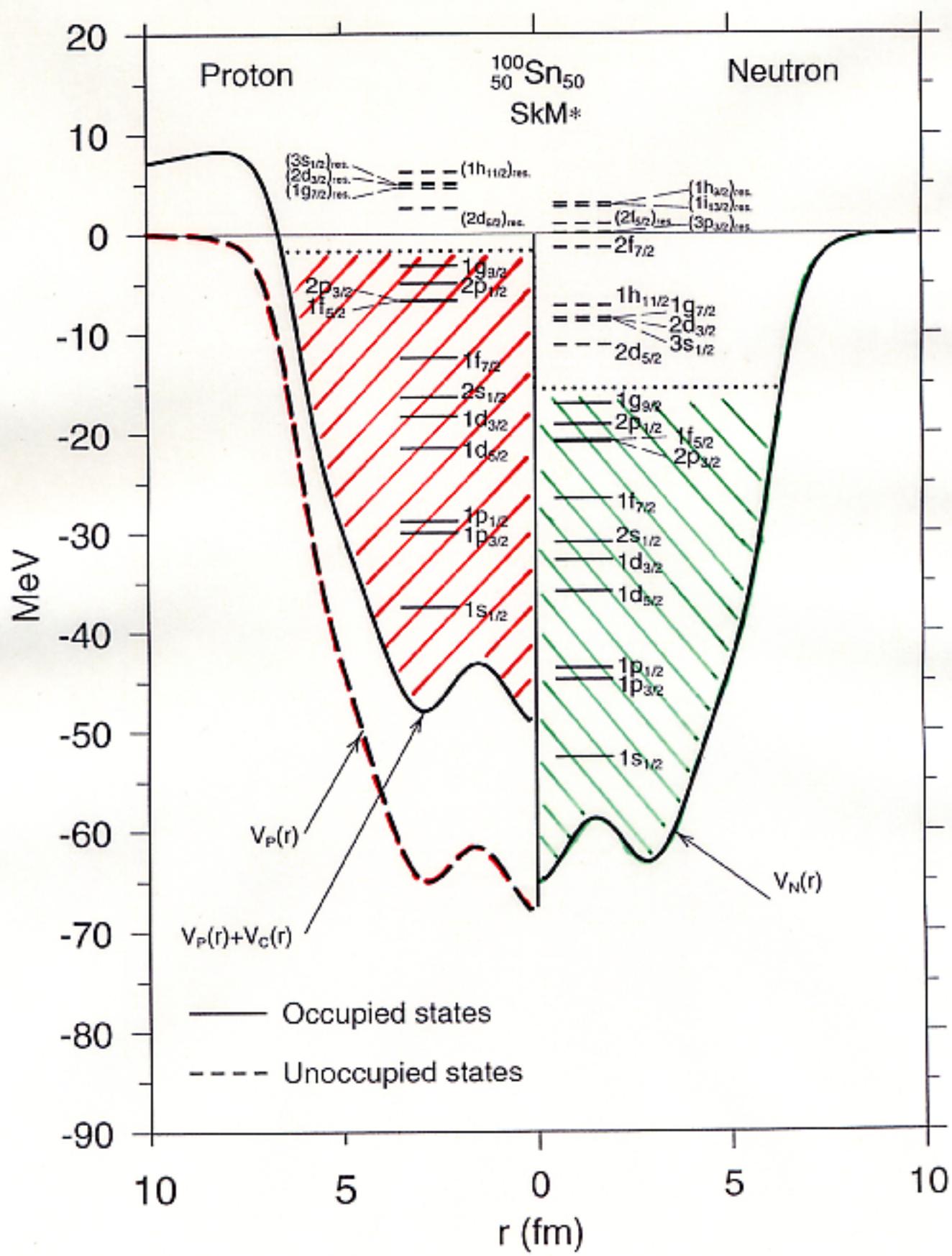
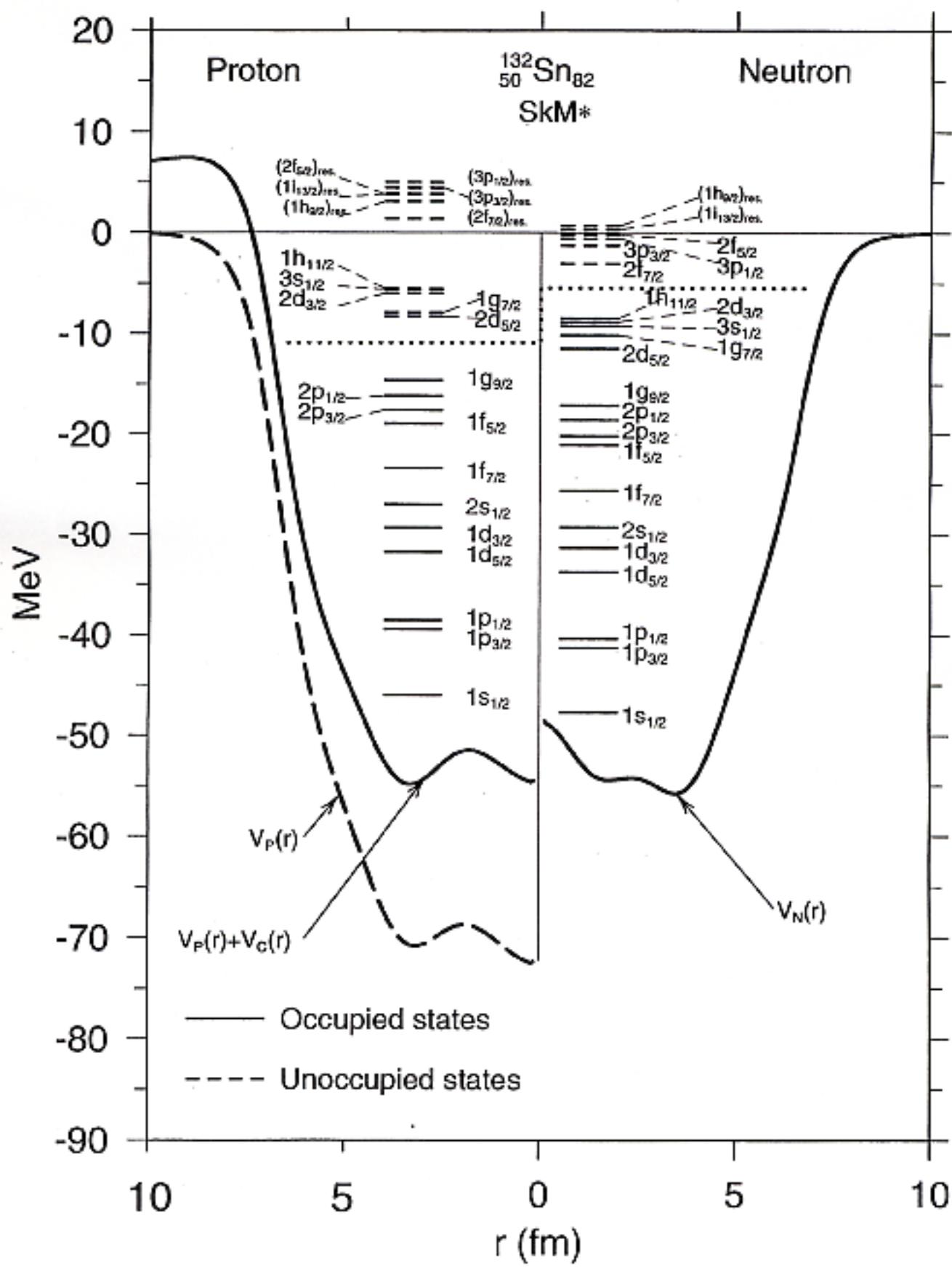
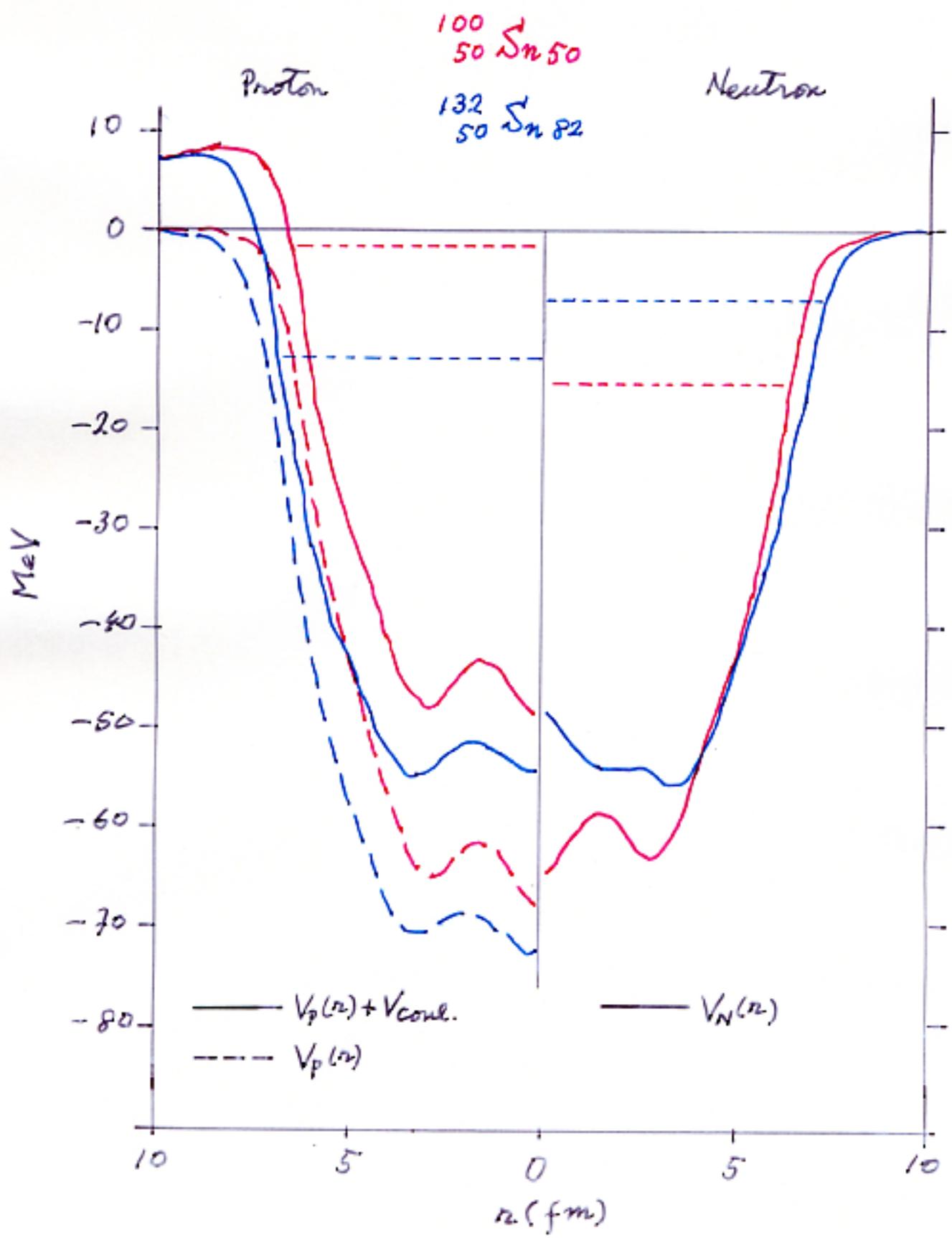


Fig. 3. Single-particle nucleonic densities of the NL-SH RMF model for  $^{100,118,132,150}\text{Sn}$  isotopes (neutrons: solid line, protons: dashed line).









Change of shell-structure, especially for  $E_{\text{nlj}} = -10 \rightarrow 0$  MeV

ex. neutron one-particle energies in Woods-Saxon potential.

Lower centrifugal barrier for particles with smaller  $l$

$\Rightarrow$  wave functions can extend beyond the nucleus  
as  $|E_{\text{nlj}}| \rightarrow 0$ .

$\Rightarrow$  { kinetic energy decreases  $\propto \langle r \frac{dU(r)}{dr} \rangle$   
spin-orbit splitting decreases  $\propto \langle \frac{1}{r} \frac{dU(r)}{dr} \rangle$

$\Rightarrow$  Shell structure may change drastically for  $|E_{\text{nlj}}| \rightarrow 0$ .

virial theorem

For nuclear potentials with realistic diffuseness  
the drastic change of shell-structure occurs for

$$E_{\text{nlj}} \approx [-10 \text{ MeV}] \rightarrow 0$$

Sn or Sp for  $\beta$ -stable nuclei.

Recent exp. data on light nuclei towards n-drip-line show:  
magic numbers 8 and 20 disappear, while  
16 is a new magic number.

References :

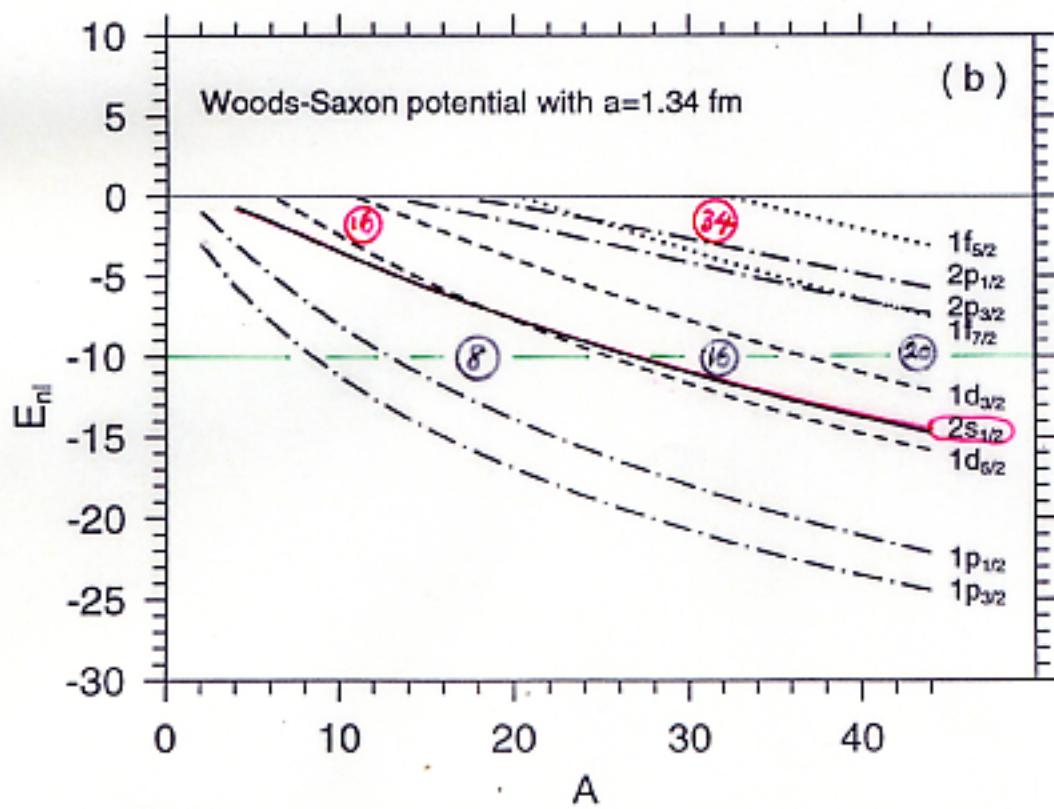
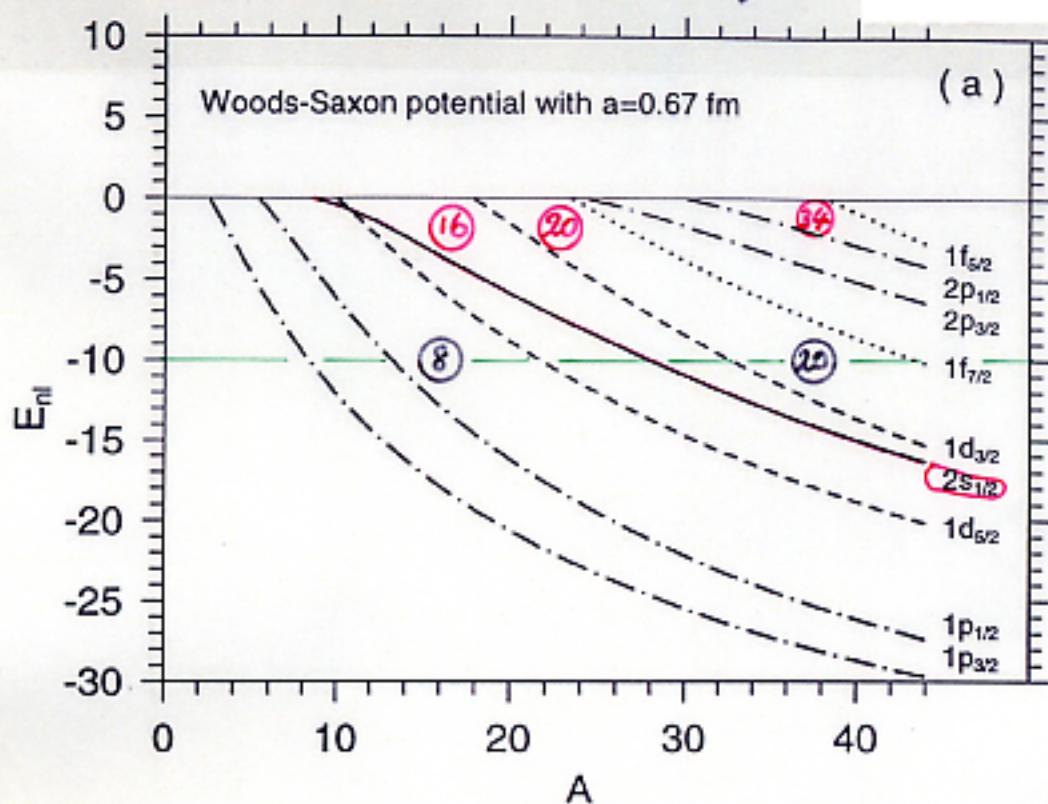
$^{32}_{12}\text{Mg}_{20}$  T. Motobayashi et al, Phys. Lett. B346 (1995) 9.

D. Guillemaud-Mueller et al, Nucl. Phys. A426 (1984) 37.

$^{12}_{4}\text{Be}_{8}$  H. Iwasaki et al, Phys. Lett. B481 (2000) ? ;  
Phys. Lett. B491 (2000) 8.

$N=16$  A. Ozawa et al, Phys. Rev. Lett. 84 (2000) 5493.

### Neutron one-particle energies

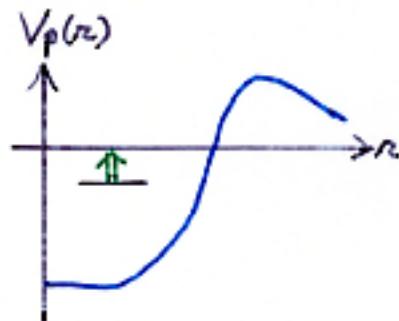


figs taken from

I. Hamamoto, S.V. Lukyanov and X.Z. Zhang,  
Nucl. Phys. A683 (2001) 255.

Close to p-drip-line :

( In medium-heavy nuclei the large Coulomb barrier hinders wave-functions from extending beyond the nucleus.



$N \approx Z$  nuclei with  $A \approx 80$

Ex.  $^{80}_{40}\text{Zr}_{40}$  is deformed.

( OBS ) all nuclei with  $N$  (or  $Z$ ) = 40 around  $\beta$ -stability line are spherical.

# Deformation

Some nuclei are deformed in the ground states.

ex.

$^{12}_6\text{C}_6$ ,  $^{20}_{10}\text{Ne}_{10}$ , rare-earth ( $90 \leq N \leq 112$ ) nuclei, ....,

## 1) Jahn-Teller effect

In the presence of many particles outside of closed-shell in spherical potentials, [the degeneracy in quantum spectra] → possibility of gaining energy by breaking away from spherical symmetry.

ex. medium-size deformation of ground states of rareearth nuclei.

## 2) New shell-structure (and new magic numbers) at large deformation.

ex. superdeformation { fission isomers in actinide nuclei.  
high-spin states in rare-earth nuclei.  
.....

strongly deformed bands in closed-shell (of spherical potentials) nuclei.

{ excited states in  $^{40}\text{Ca}_{20}$   
ground states of  $N=Z$  nuclei around  $A \sim 76$ .

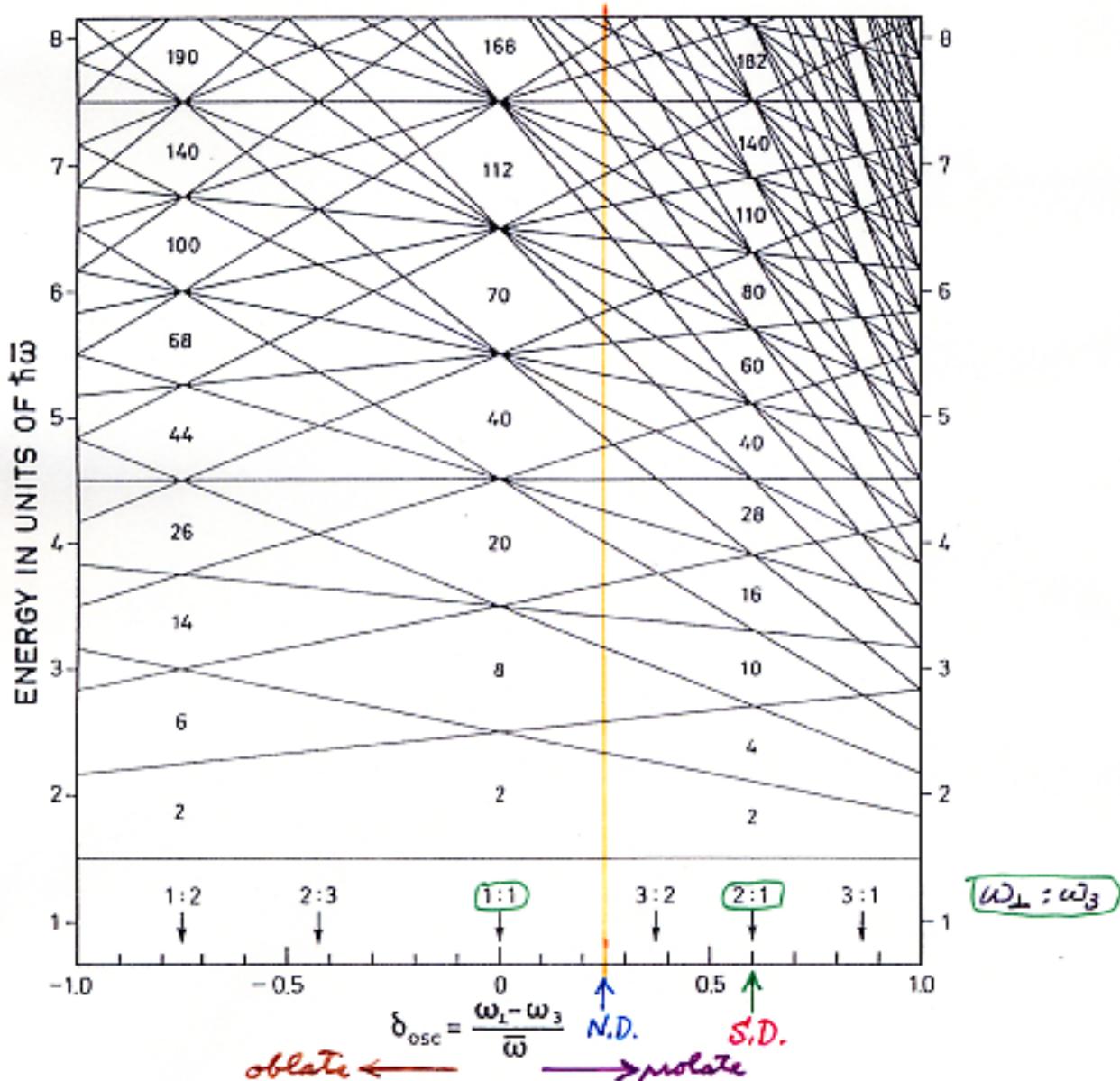
$^{72}_{26}\text{Kr}_{36}$ ,  $^{76}_{38}\text{Sr}_{38}$ ,  $^{80}_{40}\text{Zn}_{40}$ ,

Anisotropic oscillator potential

$$V = \frac{1}{2}M(\omega_3^2 x_3^2 + \omega_{\perp}^2(x_1^2 + x_2^2)),$$

$$\epsilon(n_3, n_{\perp}) = (n_3 + \frac{1}{2})\hbar\omega_3 + (n_{\perp} + 1)\hbar\omega_{\perp}$$

where  $n_3, n_{\perp} = 0, 1, 2, 3, \dots$



**Figure 6-48** Single-particle spectrum for axially symmetric harmonic oscillator potentials. The eigenvalues are measured in units of  $\bar{\omega} = (2\omega_{\perp} + \omega_3)/3$ , and the deformation parameter  $\delta_{\text{osc}}$  is that defined by Eq. (5-11). The arrows mark the deformations corresponding to the indicated rational ratios of frequencies  $\omega_{\perp} : \omega_3$ .

A. Bohr and B.R. Mottelson, Nuclear Structure, vol II.

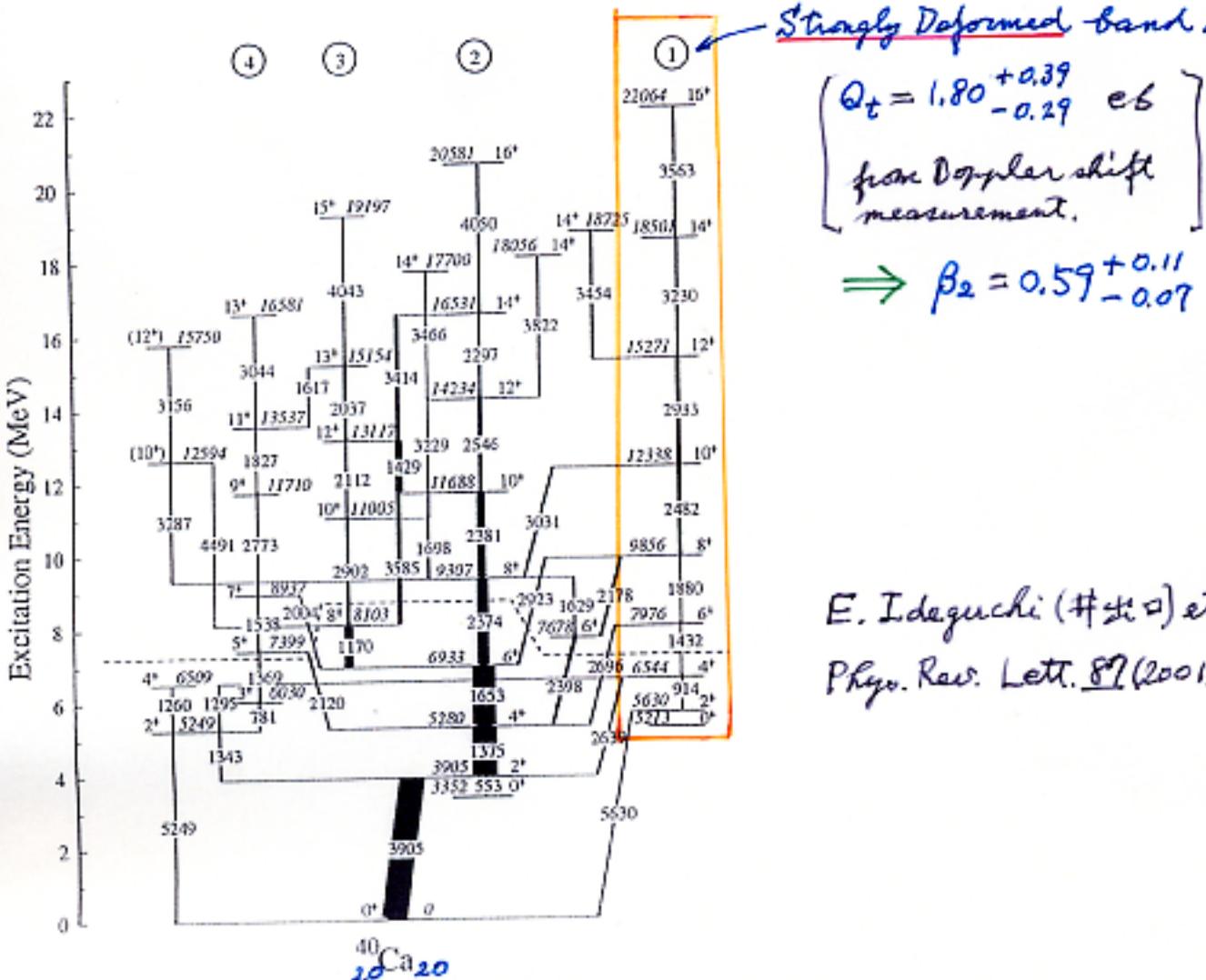


FIG. 1. Partial level scheme of  $^{40}\text{Ca}$ ; the energy labels are given in keV, and the widths of the arrows are proportional to the relative intensities of the  $\gamma$  rays. Only the levels below the dashed line were known prior to this work.

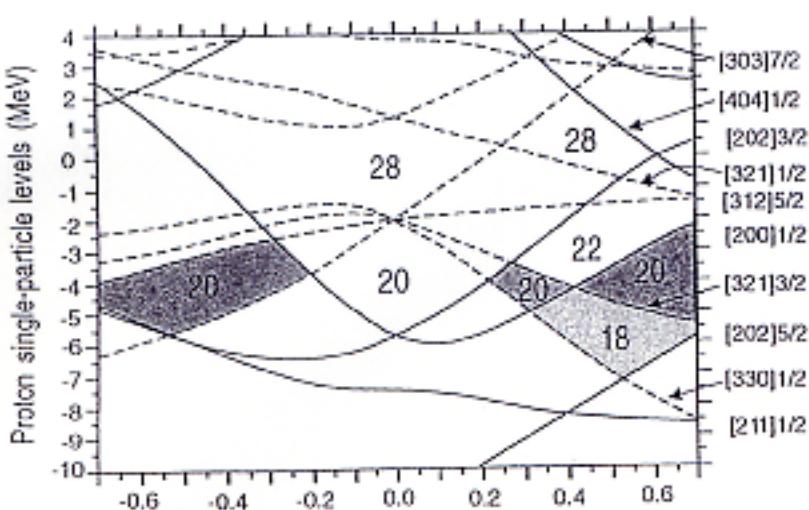


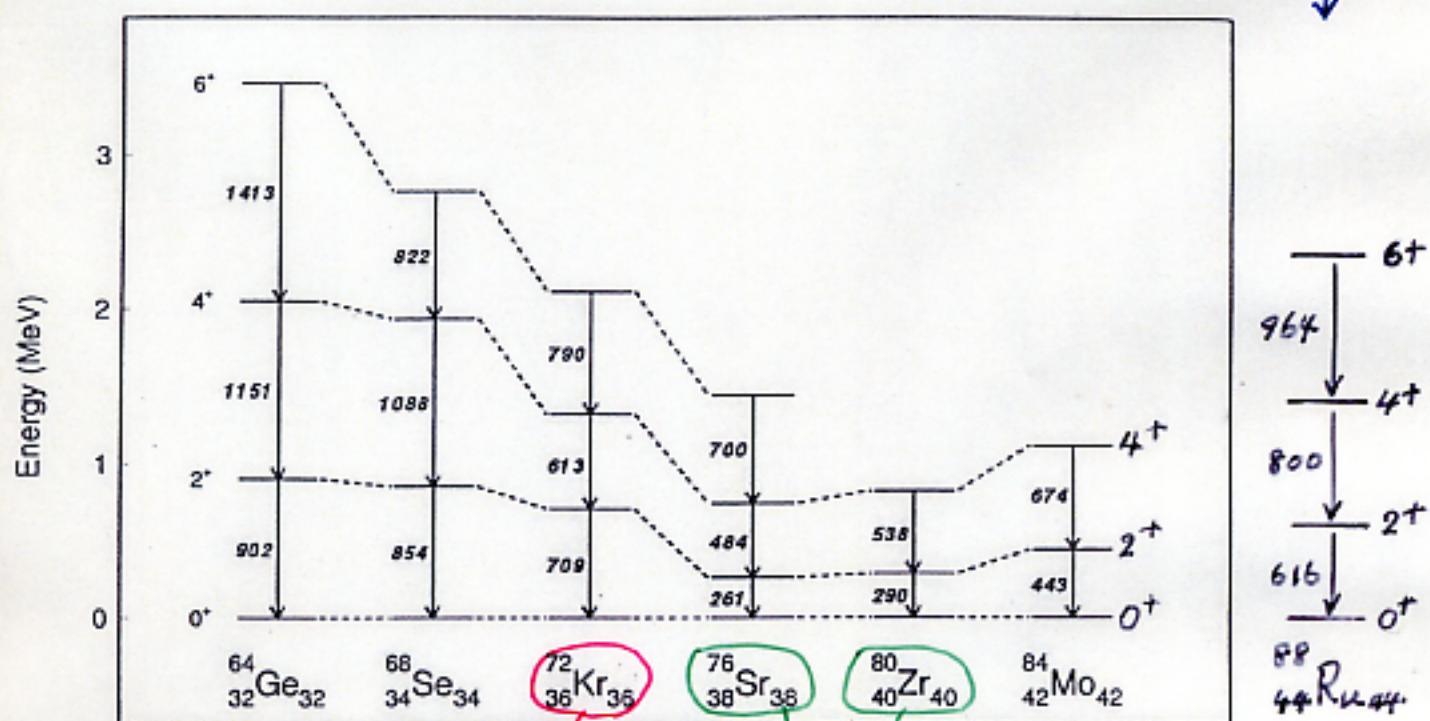
FIG. 4. Woods-Saxon orbitals as a function of the deformation,  $\beta_2$ , for particle numbers near  $N = Z = 20$ . The solid and dashed lines are for positive and negative parity orbitals, respectively. Here  $\beta_4 = 0$  was assumed. The dark shaded areas indicate the gaps responsible for the larger deformation of band 1 and the smaller deformation of band 2 in  $^{40}\text{Ca}$ .

E. Ideguchi (#210) et al.,  
Phys. Rev. Lett. 87 (2001) 212501

from  $^{58}\text{Ni}(^{32}\text{S}, 2n)$   $^{88}\text{Ru}$

at GASP

$N = Z$  nuclei



The evolution of the yrast levels in  $N=Z$  even-even nuclei.

Ovalate?      Prolate?

Fig.10.4 Observed excitation spectra of even-even  $N=Z$  nuclei. The energy differences,  $E(2^+) - E(0^+)$ ,  $E(4^+) - E(2^+)$  and  $E(6^+) - E(4^+)$ , are expressed in units of keV. The figure is taken from D.Bucurescu et al., Phys.Rev.C56(1997) 2497.

$^{80}_{40}\text{Zr}_{40}$  is deformed.

[ OBS ]  
 all nuclei with  $N$  ( $\text{or } Z$ ) = 40  
 along  $\beta$ -stability line are spherical. ]

# One-particle motion in deformed potential.—Nilsson diagram.

Spectroscopy of deformed nuclei is much **simpler** than that of spherically-vibrating nuclei, since the major part of two-body interaction is already taken into account in the deformation.

**Nilsson diagram** for axially-symmetric quadrupole deformation:  
One-particle orbitals are named by **asymptotic quantum numbers**  $[N n_z \Lambda \Omega]$ .

$\Omega (\Leftarrow j_z)$  : always **good** quantum number.

$N, n_z, \Lambda (\Leftarrow l_z)$  : become good quantum number for large  $\beta_2$ .

Each orbital has double degeneracy with  $\pm \Omega$ .

Low-lying states in deformed odd- $A$  nuclei may well be interpreted in terms of one-particle

$[N n_z \Lambda \Omega]$  orbitals, where  $\Omega = K (\Leftarrow I_z)$ , together with rotational spectra

$$E(I) = A [I(I+1) + \text{---} (-1)^{I+\frac{1}{2}} (I+\frac{1}{2}) \cdot \delta(K, \frac{1}{2})]$$

+ -----

decoupling parameter, which depends on  $(N, n_z, \Lambda)$ .

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## [home work]

Try to understand experimental information on  $^{33}_{12}\text{Mg}_{21}$ , using Nilsson diagram.  
deformed nucleus.

Slopes are determined by  $(N, n_z)$ .

Selection rules in  $[N n_z \Lambda \Omega]$  for various operators.

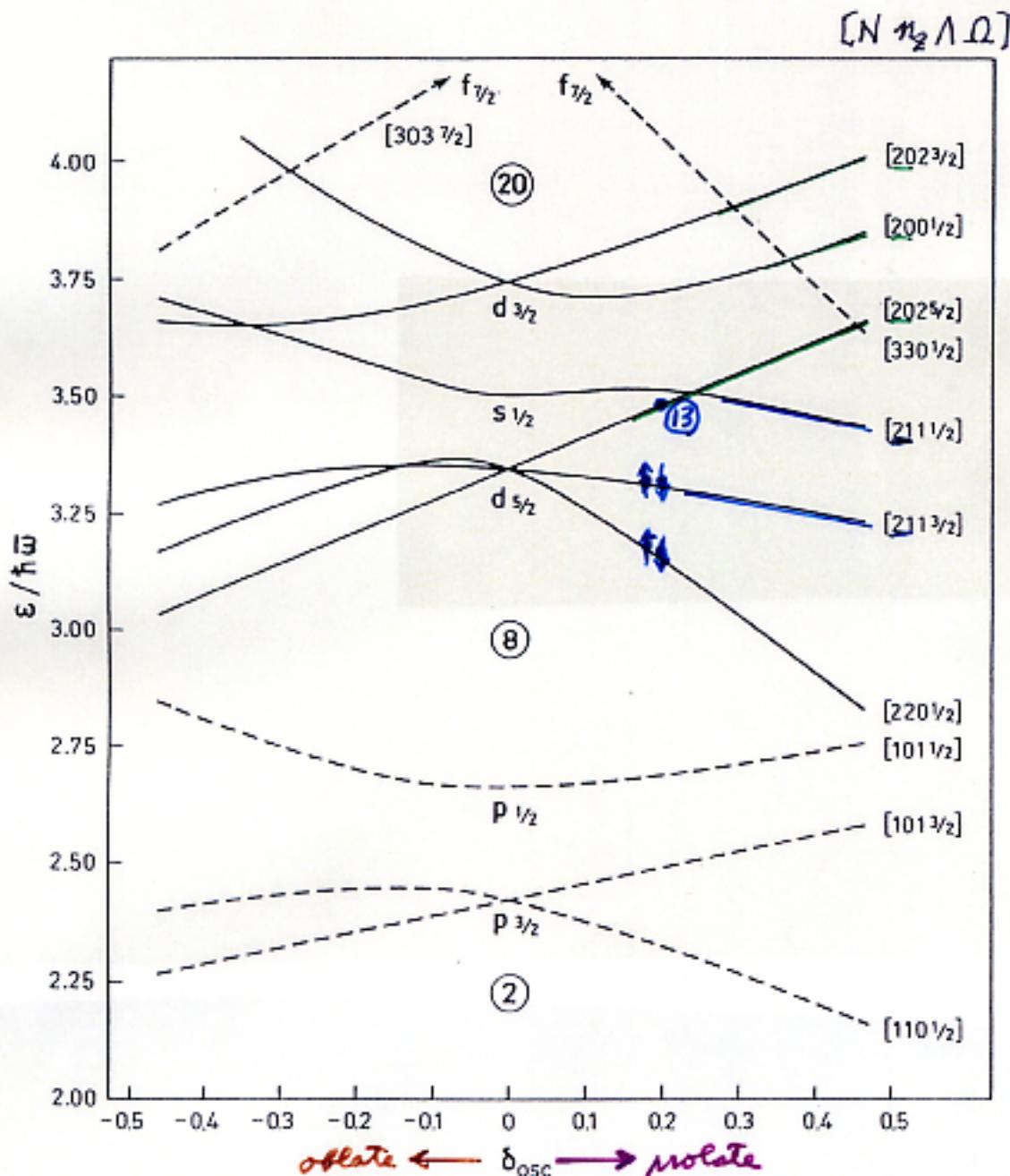
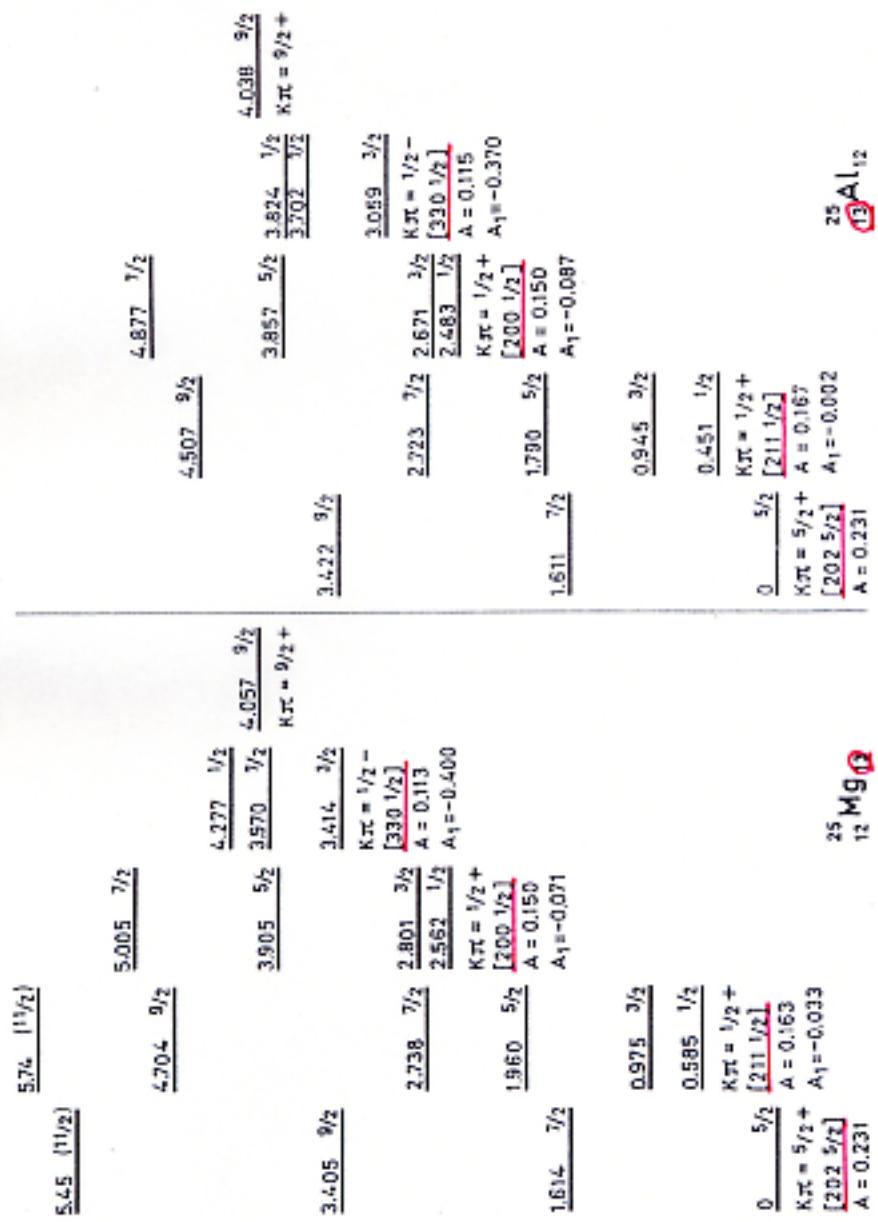


Figure 5-1 Spectrum of single-particle orbits in spheroidal potential ( $N$  and  $Z < 20$ ). The spectrum is taken from B. R. Mottelson and S. G. Nilsson, *Mat. Fys. Skr. Dan. Vid. Selsk.* 1, no. 8 (1959). The orbits are labeled by the asymptotic quantum numbers  $[N n_z \Lambda \Omega]$  referring to large prolate deformations. Levels with even and odd parity are drawn with solid and dashed lines, respectively.

[A. Bohr and B. R. Mottelson,  
Nuclear Structure, vol II.]



**Figure 5-15** Spectra of  $^{25}\text{Mg}$  and  $^{25}\text{Al}$ . The recognition of rotational band structure in the  $A=25$  system followed from the extensive series of experiments at Chalk River (see the survey by A. E. Litherland, H. McManus, E. B. Paul, D. A. Bromley, and H. E. Gove, *Can. J. Phys.* 36, 378, 1958). The figure is based on the experimental data summarized by Endt and van der Leun, 1967, and by Litherland, 1968. The tentative assignment of  $I=11/2$  for the 5.45 and 5.74 MeV states in  $^{25}\text{Mg}$  is based on the data of S. Hinds, R. Middleton, and A. E. Litherland, *Proceedings of the Rutherford Jubilee Intern. Conf.*, p. 305, ed. J. B. Birks, Heywood and Co., London 1961. Energies are in units of MeV.

The 13<sup>th</sup> nucleon occupies the orbital [ $Nm_3 \Lambda \Omega$ ].

## § Collective modes

8/10>

(The best collective motion in nuclei is **rotation** of deformed nuclei.)

Giant Resonances (   : experimentally well-established )

1) Charge-exchange ( $t_{\pm}$ ) modes.

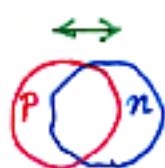
	operator	$I^{\pi}$ of the operator
Fermi (IAS')	$\sum_i t_{\pm}(i)$	$0^+$
Gamow-Teller (GTGR)	$\sum_i t_{\pm}(i) \sigma_{\mu}(i)$	$1^+$
Spin dipole (SDR)	$\sum_i t_{\pm}(i) r_i (Y_1 0)^I_i$	$0^-, 1^-, 2^-$
Spin monopole* (SGMR)	$\sum_i t_{\pm}(i) r_i^2 \sigma_{\mu}(i)$	$1^+$

2) ( $t_1$  or  $t_2$ ) modes

$t_1$ IS $t_2$ IV	operator	$I^{\pi}$ of the operator	Observed energy for $A \gtrsim 60$
IV dipole (IVGDR)	$\sum_i t_2(i) r_i Y_{1\mu}(i)$	$1^-$	$79 \text{ fm}^{-1/3} \text{ MeV}$
IS' quadrupole (IS'GQR)	$\sum_i r_i^2 Y_{2\mu}(i)$	$2^+$	$58 \text{ fm}^{-1/3} \text{ MeV}$
IS' monopole* (IS'GMR)	$\sum_i r_i^2$	$0^+$	$80 \text{ fm}^{-1/3} \text{ MeV}$
IV quadrupole (IVGQR)	$\sum_i t_2(i) r_i^2 Y_{2\mu}(i)$	$2^+$	
IS' dipole* (IS'GDR)	$\sum_i r_i^3 Y_{1\mu}(i)$	$1^-$	

\*: compression mode

IVGDR



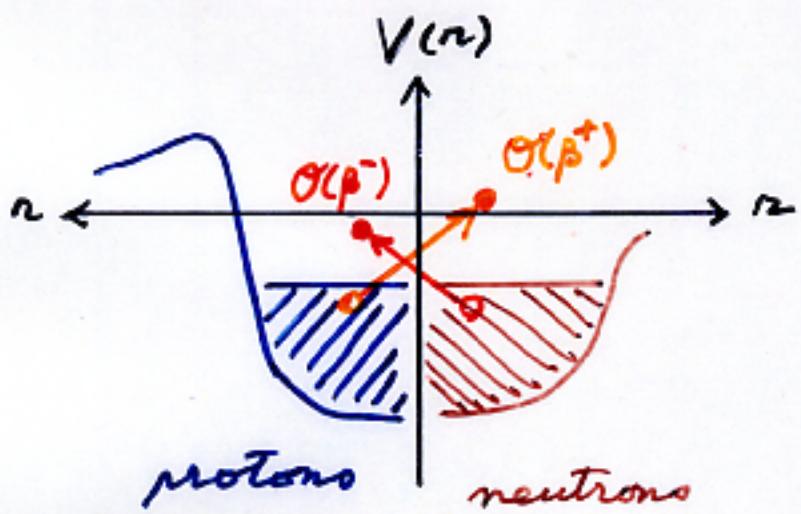
IS'GQR



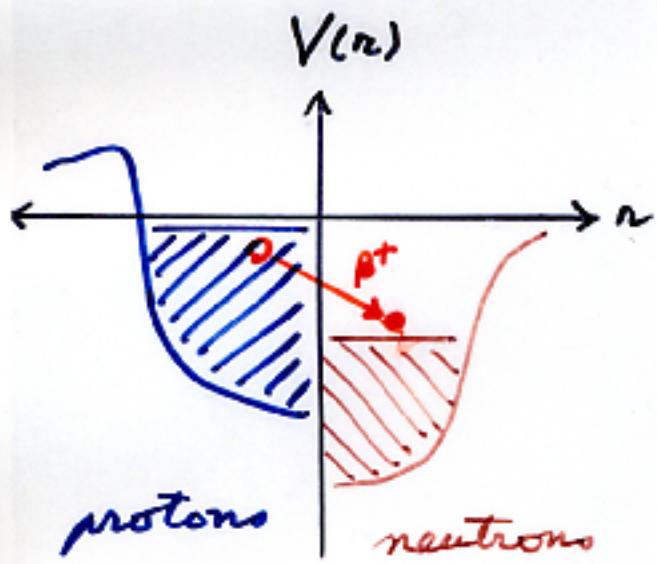
IS'GMR



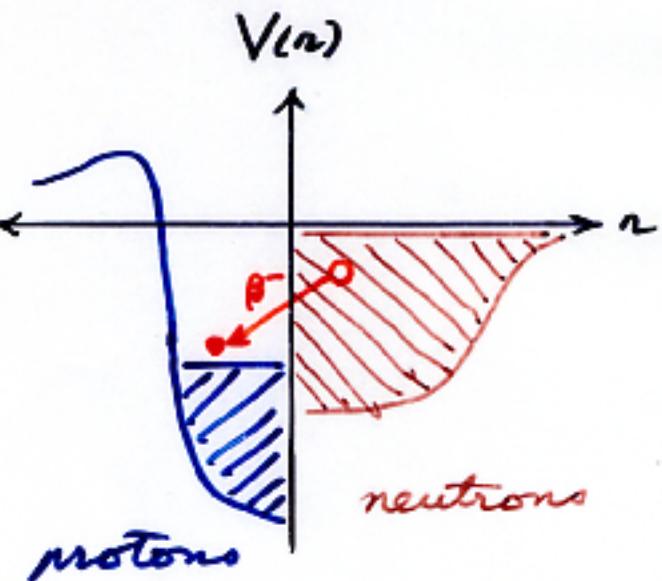
### $\beta$ -stable nuclei



### $p$ drip line nuclei



### $n$ drip line nuclei



## § Giant resonances and low-energy threshold strength.

### Quadrupole mode

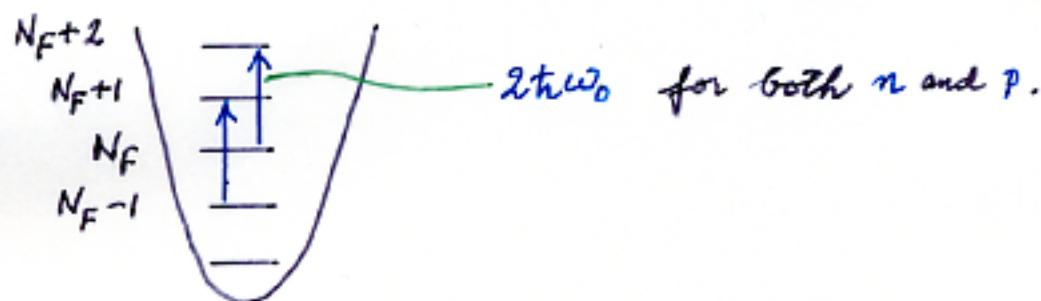
$$Q = \sum_i n_i^2 Y_{2\mu}(i) \quad \text{for isoscalar}$$

$$Q = \sum_i T_2^{(i)} n_i^2 Y_{2\mu}(i) \quad \text{for isovector}$$

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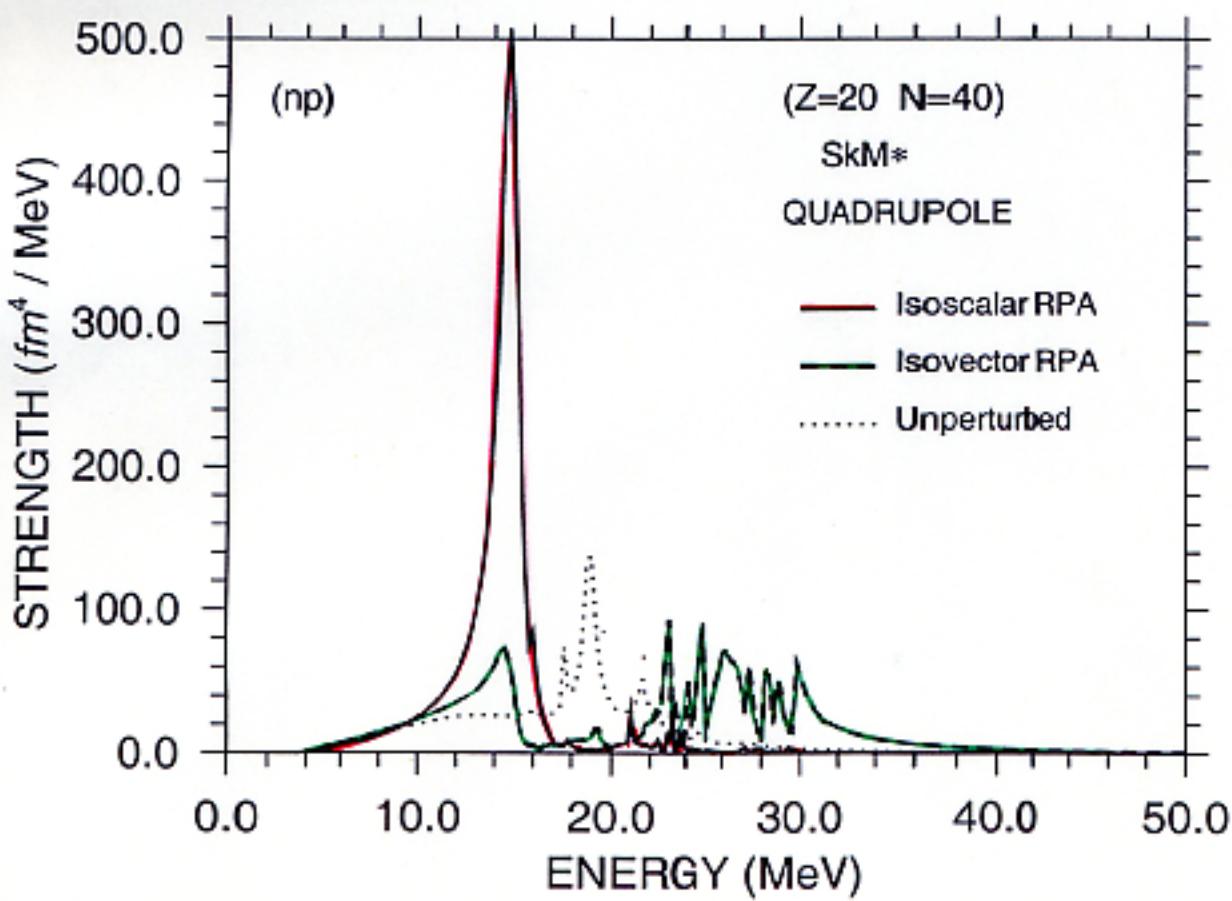
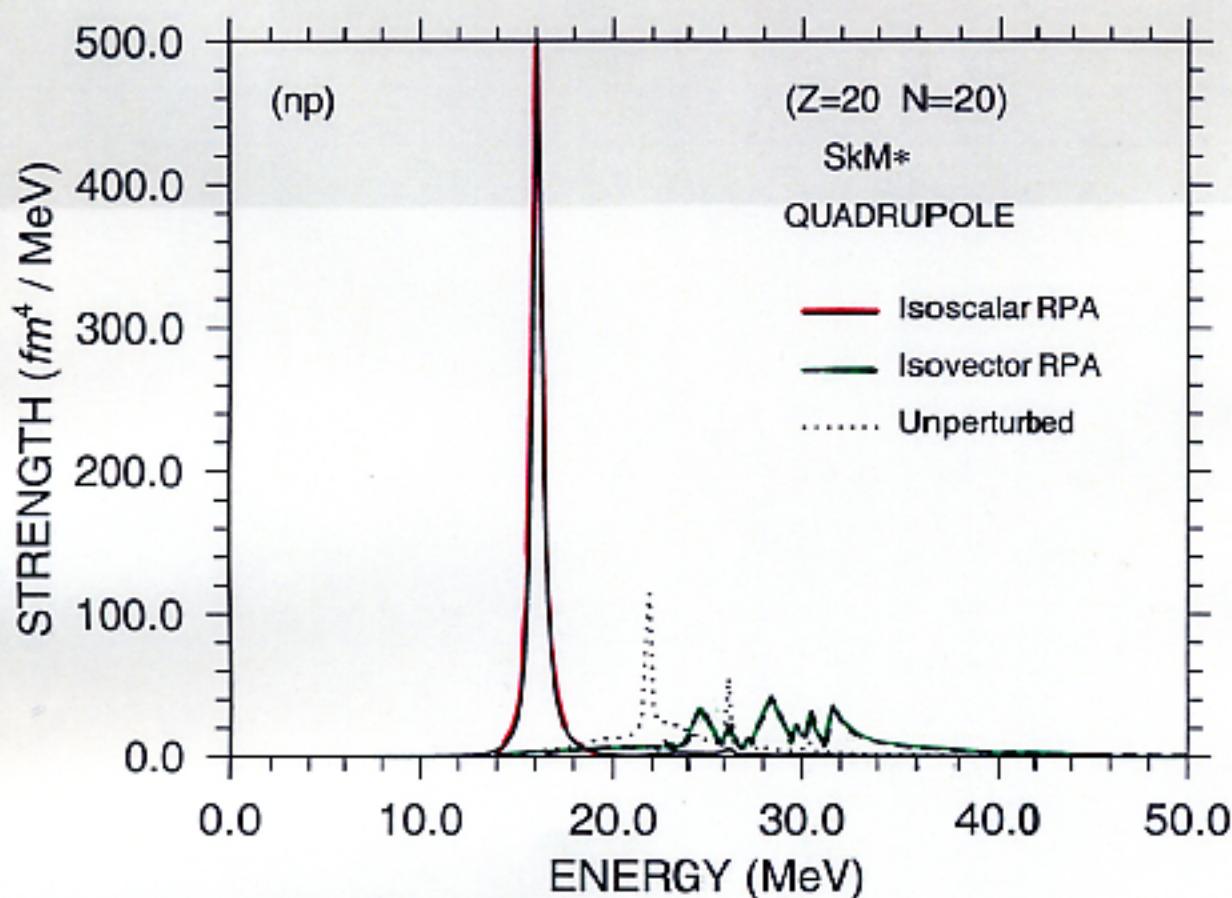
ex.  
 $l=5$  closed-shell nuclei,  $^{40}_{20}\text{Ca}_{20}$  &  $^{60}_{20}\text{Ca}_{40}$  :

In the harmonic-oscillator model,  
unperturbed P-h excitations at  $2\hbar\omega_0$ .  
(no low-lying excitations!)

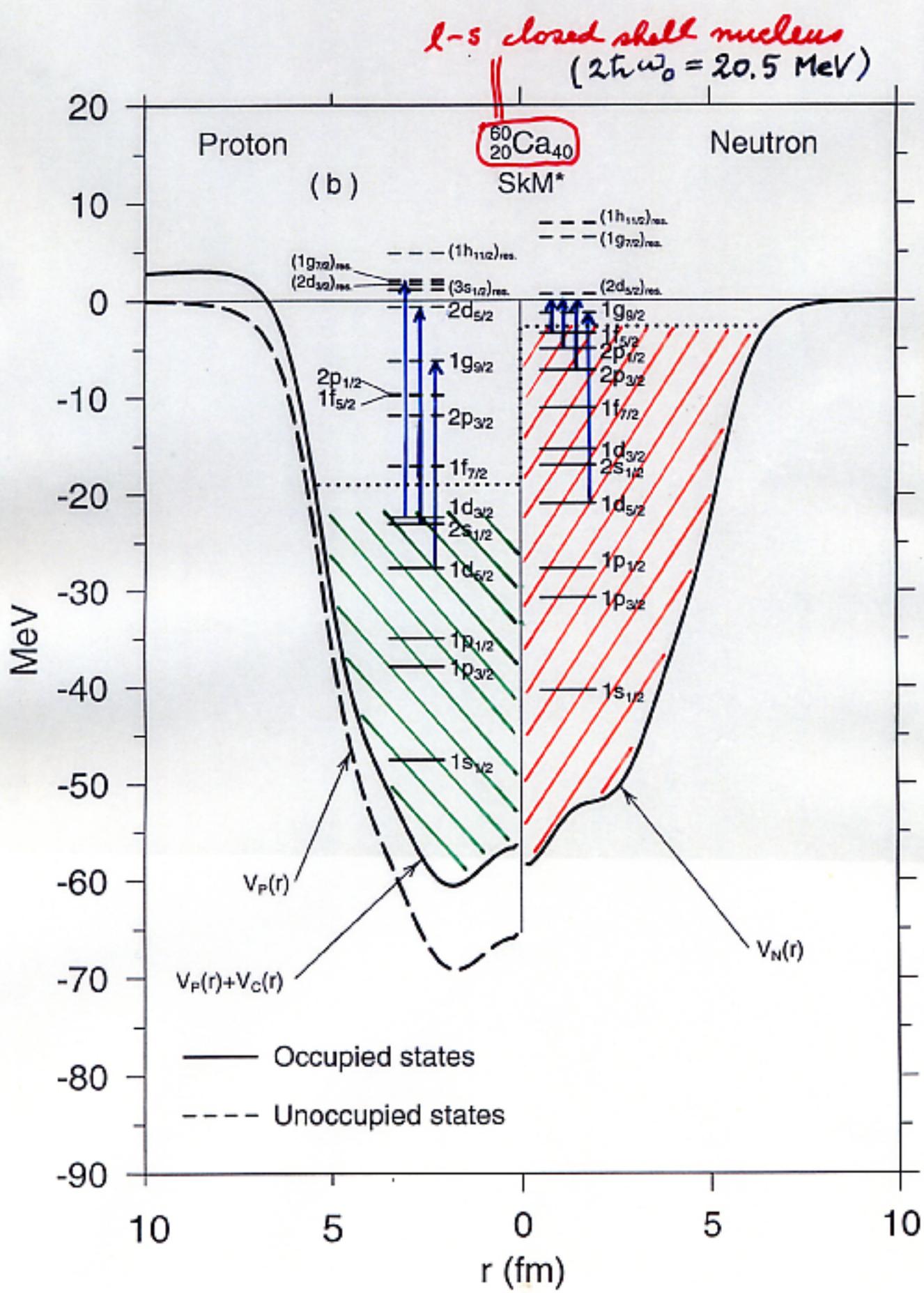


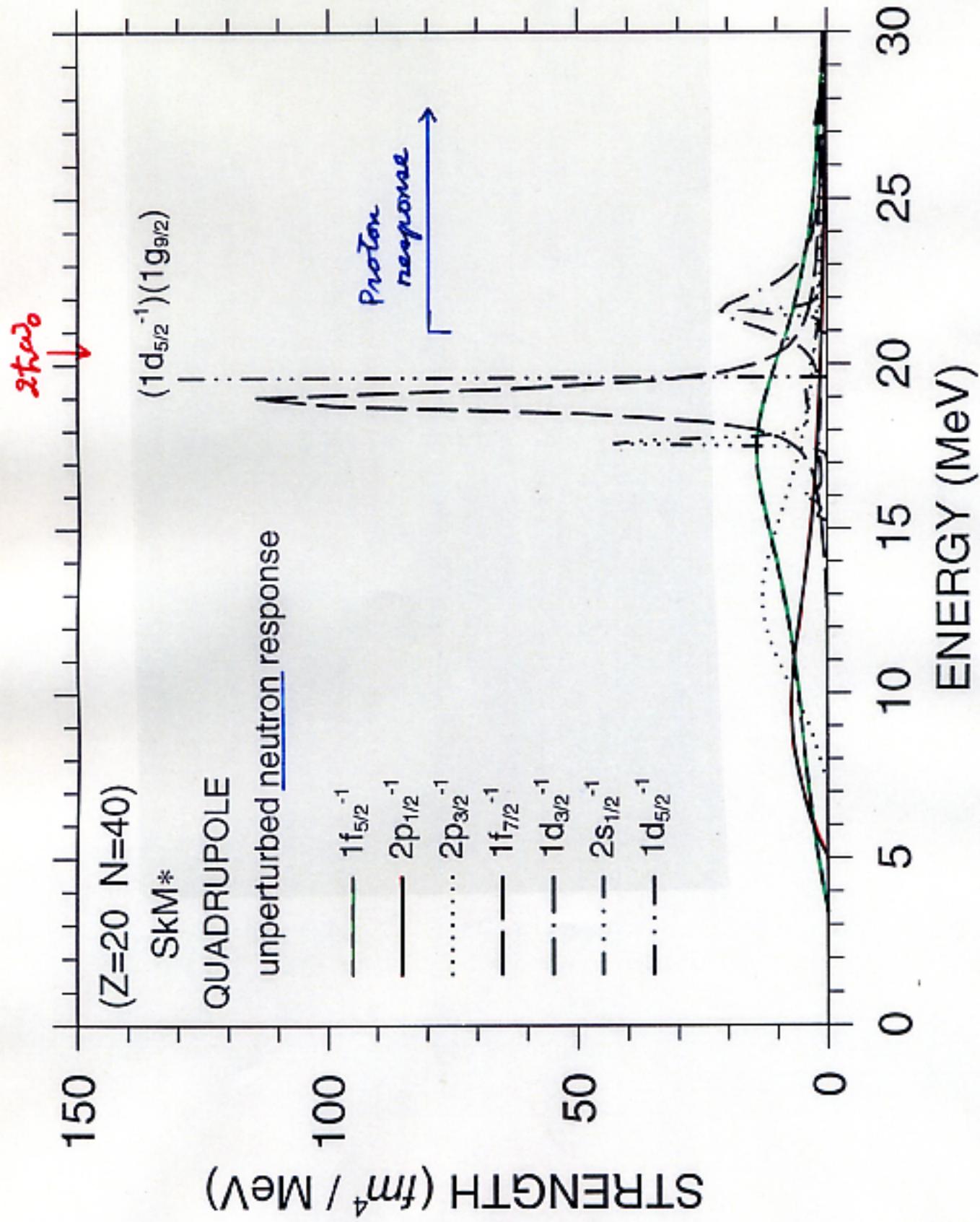
Quadrupole ( $2^+$ ) excitation

self-consistent HF+RPA with Skyrme interaction

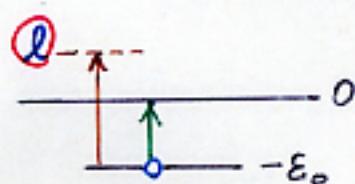
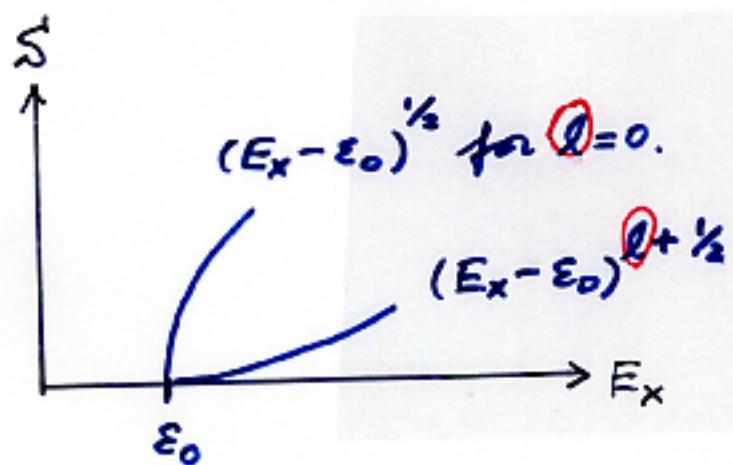


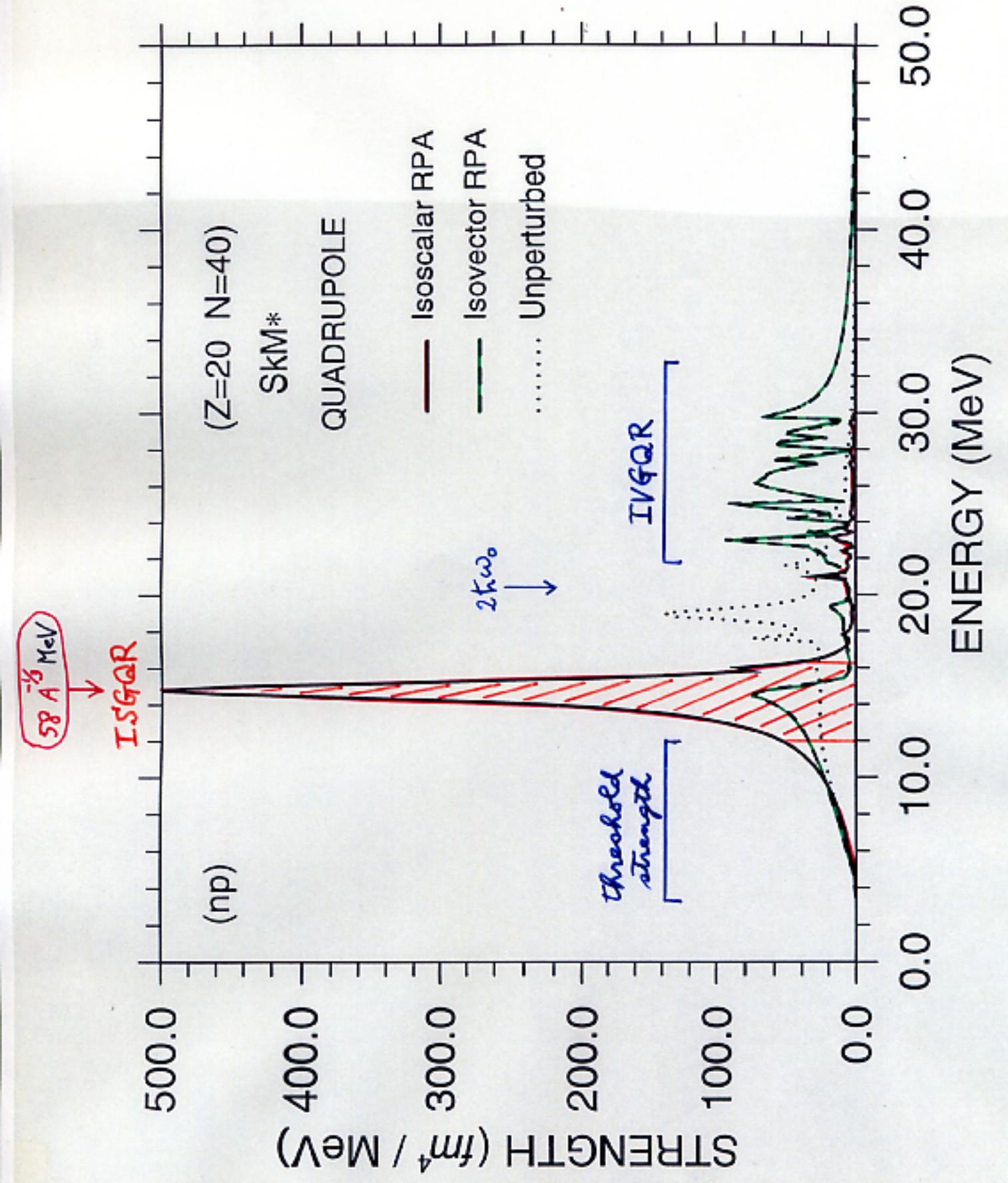
## Quadrupole ( $2^+$ ) excitation





For neutrons





## Model-independent SUM-RULES

Giant Resonances carry a major part of sum-rules.

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1) Charge-exchange ( $t_{\pm}$ ) modes on  $|0\rangle$  ;

$$\left[ \sum_i t_+(i), \sum_j t_-(j) \right] = 2 \sum_k t_z(k)$$

$$2 \langle 0 | \sum_i t_z(i) | 0 \rangle \\ = N - Z,$$

(a) Fermi transitions ;  $\sigma = \sum_i t_{\pm}(i)$ ,

$$\sum_m |\langle m | \sum_i t_-(i) | 0 \rangle|^2 - \sum_n |\langle n | \sum_i t_+(i) | 0 \rangle|^2 = N - Z$$

(b) Gamow-Teller ;  $\sigma_{ji} = \sum_i t_{\pm}(i) \tilde{\sigma}_{ji}(i)$ ,

$$\sum_{ji} \sum_m |\langle m | \sum_i t_-(i) \tilde{\sigma}_{ji}(i) | 0 \rangle|^2 - \sum_{ji} \sum_n |\langle n | \sum_i t_+(i) \tilde{\sigma}_{ji}(i) | 0 \rangle|^2 \\ = 3(N - Z),$$

$$\sum_j \tilde{\sigma}_{ji}^2 = 3$$

(c)  $\sigma = \sum_i t_{\pm}(i) f(n_i)$ ,

$$\sum_m |\langle m | \sum_i t_-(i) f(n_i) | 0 \rangle|^2 - \sum_n |\langle n | \sum_i t_+(i) f(n_i) | 0 \rangle|^2 \\ = N \cdot \langle (f(n))^2 \rangle_n - Z \cdot \langle (f(n))^2 \rangle_p.$$

ex. If  $|\langle n | \sum_i t_+(i) \tilde{\sigma}_{ji}(i) | 0 \rangle|^2 = 0$  due to a sufficient neutron excess,

then,

$$\sum_j \sum_m |\langle m | \sum_i t_-(i) \tilde{\sigma}_{ji}(i) | 0 \rangle|^2 = 3(N - Z).$$



2) ( $t_1$  or  $t_2$ ) modes

For one-body operator  $F_{\lambda\mu} = \sum_i f(r_i) \cdot Y_{\lambda\mu}(i)$

$$[H, F_{\lambda\mu}] \Rightarrow [T, F_{\lambda\mu}], \quad \text{if } [v_{ij}, F_{\lambda\mu}] \neq 0$$

kinetic energy

2-body interaction does  
not explicitly depend on  
momenta of particles.

Then,

Energy-weighted sum-rule

$$S(F) \equiv \sum_{\mu} \sum_{\alpha} (E_{\alpha} - E_0) |\langle \alpha | F_{\lambda\mu} | 0 \rangle|^2$$

$$= \frac{1}{2} \sum_{\mu} \langle 0 | [F_{\lambda\mu}, [H, F_{\lambda\mu}]] | 0 \rangle$$

$$\Rightarrow \frac{1}{2} \sum_{\mu} \langle 0 | [F_{\lambda\mu}, [T, F_{\lambda\mu}]] | 0 \rangle$$

$$= \sum_{\mu} \langle 0 | \frac{\hbar^2}{2M} \sum_{ik} |\nabla_k F_{\lambda\mu}(\vec{r}_{ik})|^2 | 0 \rangle$$

$$= \frac{2\lambda+1}{4\pi} \frac{\hbar^2}{2M} A \left( \left( \frac{df}{dn} \right)^2 + \lambda(\lambda+1) \left( \frac{f}{n} \right)^2 \right)$$

expectation value  
in the ground state

ex.  $F_{\lambda\mu} = n^{\lambda} Y_{\lambda\mu}$

$$E\lambda \sim (\frac{1}{2} - t_2) \cdot 0$$

$$S(F) = \frac{\lambda(2\lambda+1)^2}{4\pi} \frac{\hbar^2}{2M} A \langle n^{2\lambda-2} \rangle$$

$$S(E\lambda)_d = \frac{\lambda(2\lambda+1)^2}{4\pi} \frac{\hbar^2}{2M} Z e^2 \langle n^{2\lambda-2} \rangle_{\text{nucleon}} \quad \text{for } \lambda \geq 2.$$

ex. monopole mode  $F = n^2$

$$S(E\lambda)_d = \frac{2\hbar^2}{M} Z e^2 \langle n^2 \rangle_{\text{nucleon}}$$

$\frac{N}{A}$  for protons  
 $-\frac{Z}{A}$  for neutrons

ex. dipole mode

$$F(E1)_{1\mu} = e \sum_k \left[ \left( \frac{1}{2} - t_2 \right) n Y_{1\mu} \right]_k \rightarrow e \sum_k \left[ \left( \frac{N-Z}{2A} - t_2 \right) n Y_{1\mu} \right]_k$$

(E1 operator must be referred to center of mass motion.)

$$S(E1)_d = \frac{9}{4\pi} \frac{\hbar^2}{2M} e^2 \left[ Z \left( \frac{N}{A} \right)^2 + N \left( \frac{-Z}{A} \right)^2 \right]$$

$$= \frac{9}{4\pi} \frac{\hbar^2}{2M} e^2 \frac{NZ}{A} = 14.8 \frac{NZ}{A} e^2 \text{ fm}^2 \text{ MeV}$$

The motion of a particle is associated with a recoil of the rest of the nucleus, since the total center of mass remains at rest.

— This is especially important in  $E1$  transitions.

$$e \rightarrow (e)_{E1} = \left[ \left( \frac{1}{A} - t_2 \right) - \frac{Z}{A} \right] e = \begin{cases} \frac{N}{A} e & \text{for protons.} \\ -\frac{Z}{A} e & \text{for neutrons.} \end{cases}$$

$\frac{N-Z}{2A} - t_2$

ex. I.S quadrupole modes,  $F_{2\mu} = r^2 Y_{2\mu}$

$$S(r^2 Y_{2\mu}) = \frac{2 \cdot 5^2}{4\pi} \frac{\hbar^2}{2M} A \cdot \langle r^2 \rangle, \quad \langle r^2 \rangle : \text{ground-state expectation-value}$$

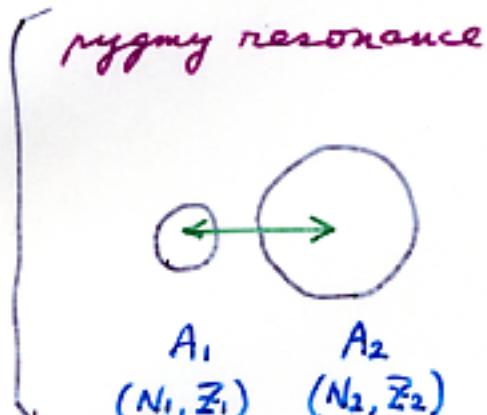
Threshold strength, which comes from loosely-bound nucleons, has extra contribution to  $\langle r^2 \rangle$ .

Thus,  $S(r^2 Y_{2\mu})$ -values are larger for nuclei with threshold strength.

ex.  $S(E1)$  is unique.

Both possible threshold strength and so-called pygmy resonance are a part of

$$S(E1)_{\alpha} = \frac{9}{4\pi} \frac{\hbar^2}{2M} e^2 \frac{NZ}{A}$$



The contribution to  $S(E1)$  is

$$\frac{9}{4\pi} \frac{\hbar^2}{2M} e^2 \frac{(Z_1 A_2 - Z_2 A_1)^2}{A_1 A_2}$$

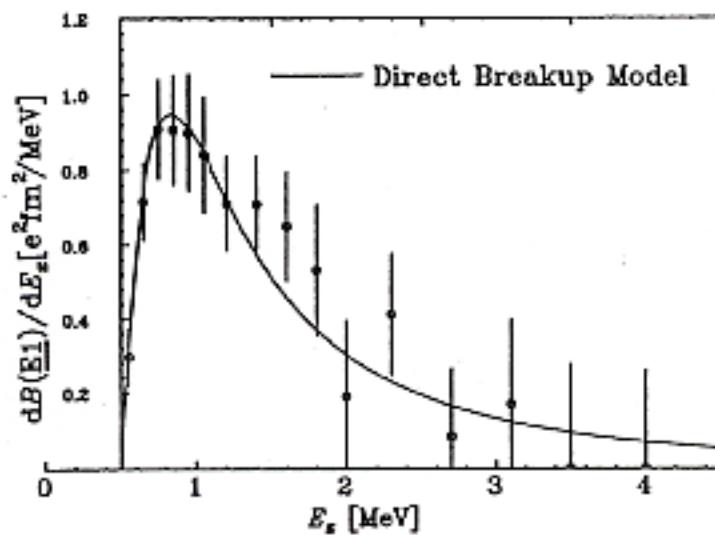
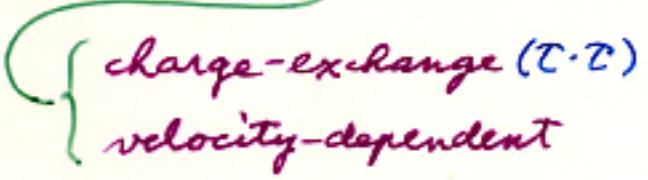


Figure 1. Dipole strength distribution of  $^{11}\text{Be}$ . The solid line represents the direct break up model (eq.(1)), after taking the detector resolution and the post acceleration into consideration.

One-neutron halo nucleus,  $^{11}\text{Be}\gamma$ .

$$S(E\lambda) = S(E\lambda)_{cl} (1 + \alpha)$$



component in the  
nucleonic interactions.

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velocity-dependent component of **S**teyne interactions  
( $\propto \delta(\vec{r}_1 - \vec{r}_2)$ ) do not contribute to  $S(v=0)$  such as  
 $S(r^2 Y_{20})$ .