

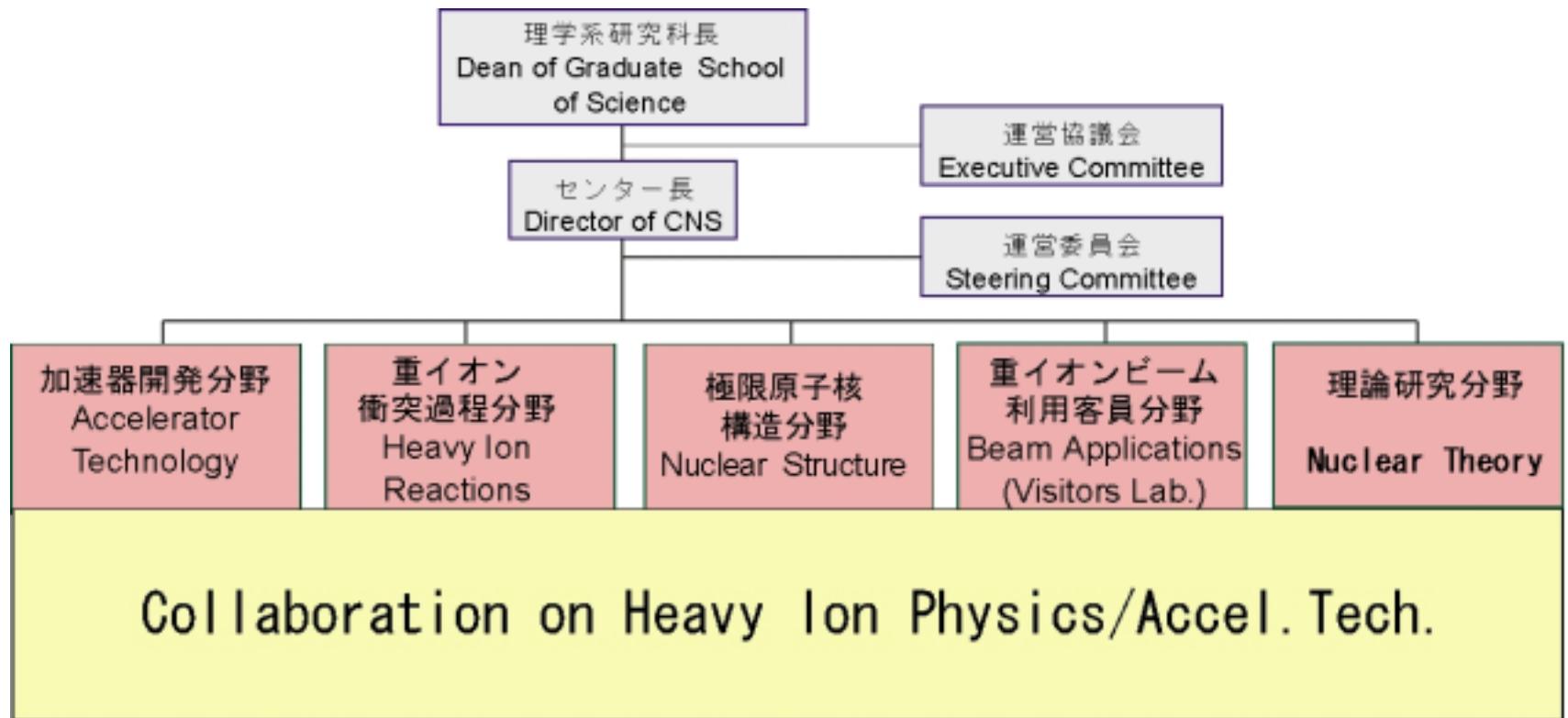
Research Programs at CNS

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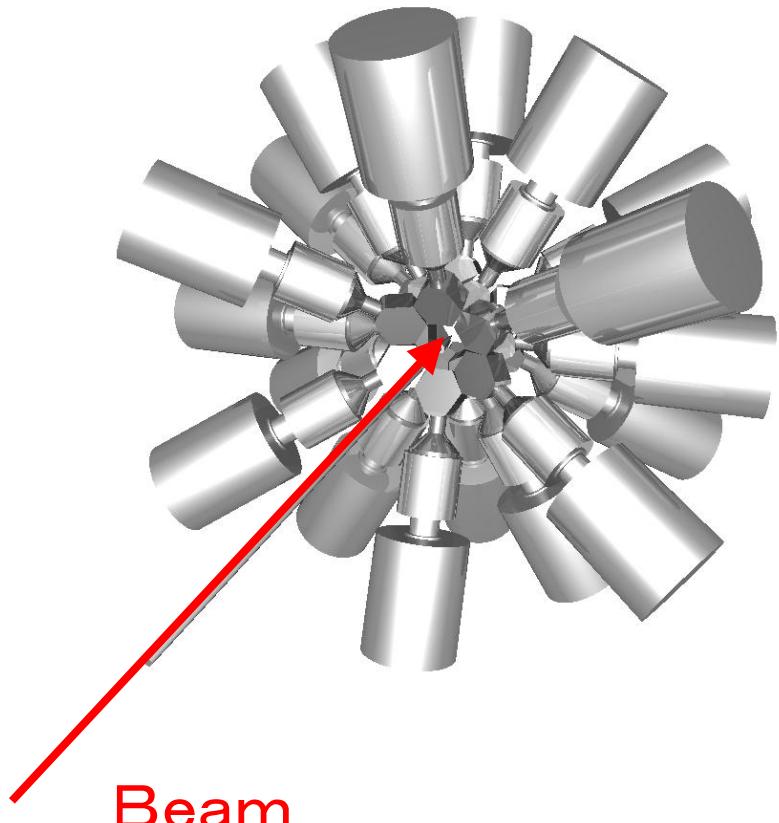
- 1. Outline of CNS**
- 2. Research Opportunities and the Scope**
- 3. Nuclear Astrophysics**
 - Hot pp-chain**
 - Neutron source for the s-process**
 - ANC method for the (n,γ) reaction**

CNS Organization



- Dev. of AVF
ECR
Accel. design
Heavy ion fusion
 - Nuclear structure/
reactions of exotic
nuclei
 - Polarization / Primary beams
- Nuclear Astrophysics
- High Energy Heavy Ion
(RHIC)
- Dev. of AVF
Cluster IS
Laser IS
- Shell Model
Three body

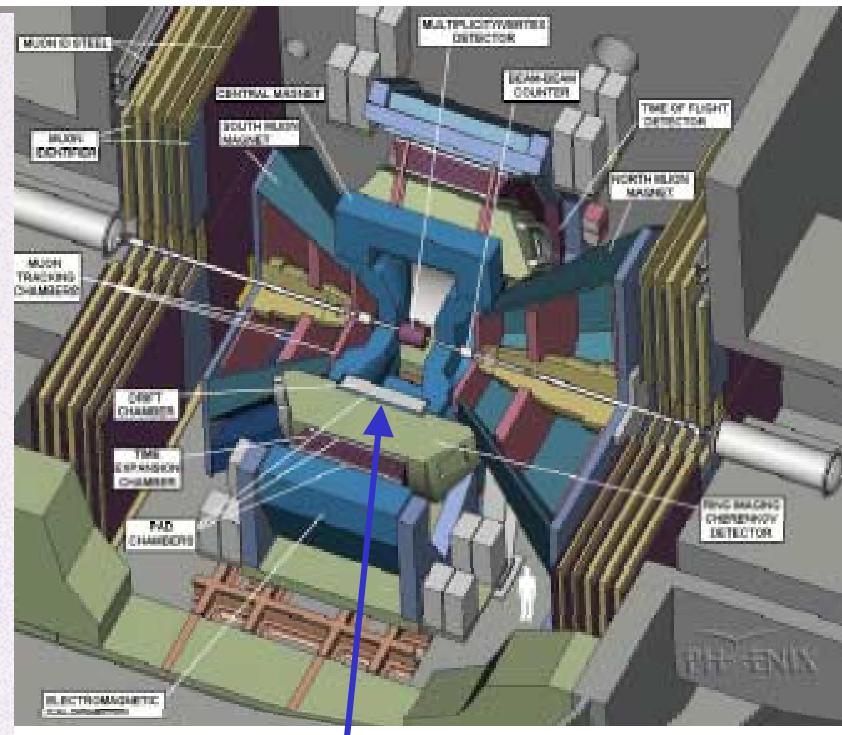
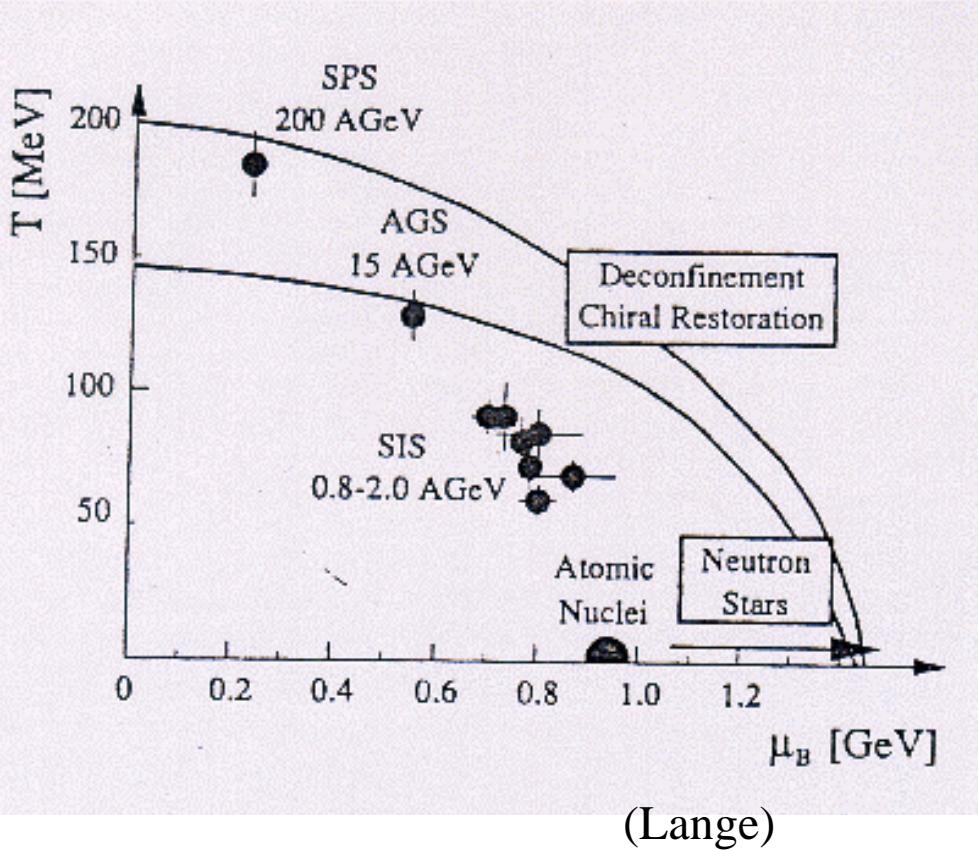
CNS Ge Detector Array



- 18 position-sensitive Ge detectors
- Position resolution $\approx 1\text{mm FWHM}$
→ Doppler Shift Correction
- Total detection efficiency = 5 %
for 1 MeV gamma rays

Physics of Unstable Nuclei
cluster, magicity, halo, G.R.,

PHENIX Experiment at RHIC(@BNL)



PHENIX実験装置

RICH detector / CNS

High density matter ?
Quak-gluon plasma ?

基幹実験装置 : Facilities

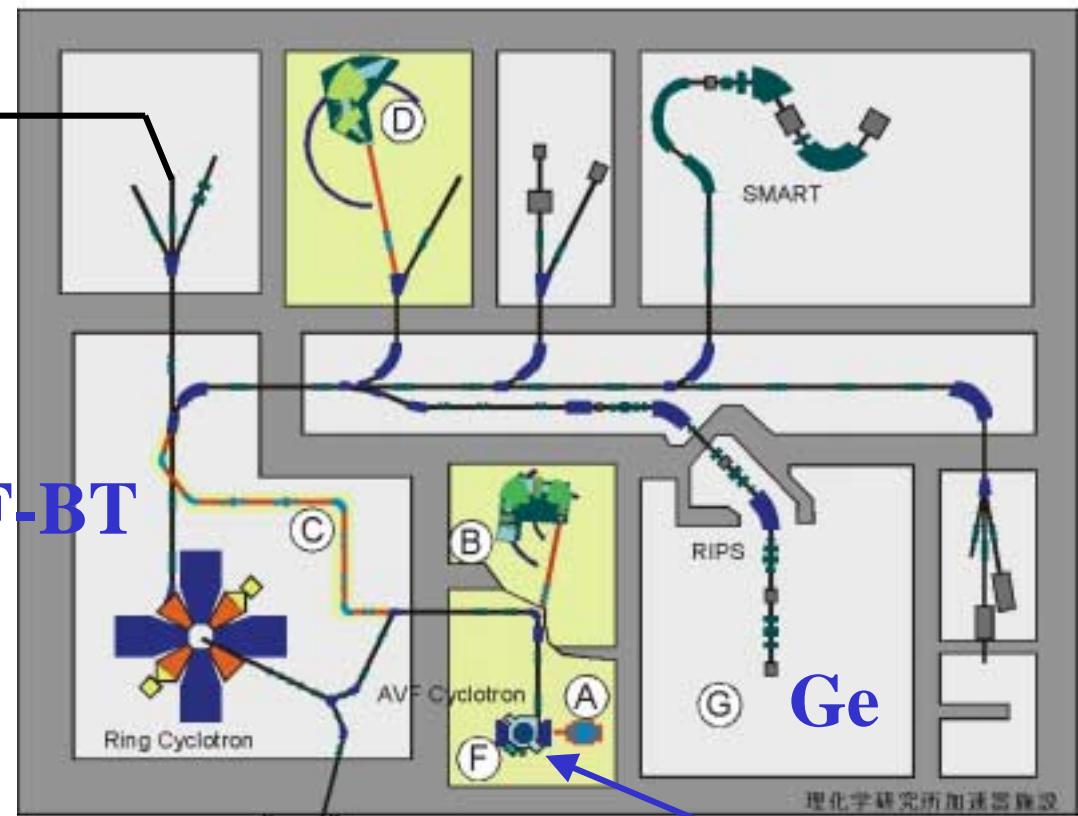
PA

RIBF

CNS Facilities at RIKEN

(Under CNS-RIKEN joint venture)

AVF-BT



CRIB

AVF/ECR

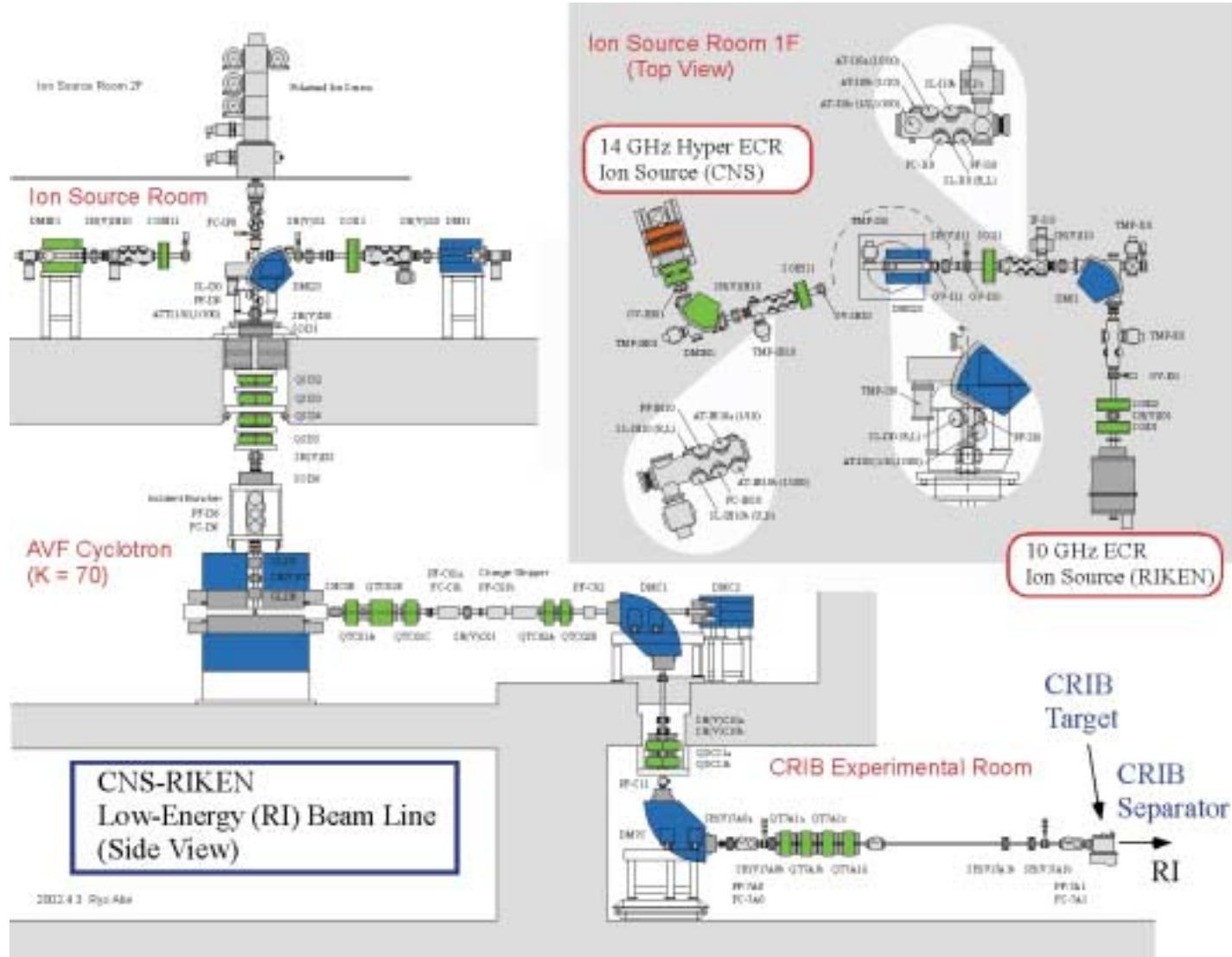
- (A) 大強度重イオン源
- (B) 低エネルギー二次ビーム分離器 CRIB
- (C) AVFビームライン
- (D) 反応粒子磁気分析器
- (E) ビーム反応実験・学生教育実験装置
- (F) AVFサイクロトロンの高性能化(計画中)
- (G) インビーム分光用 Ge ボール(計画中)

CSM

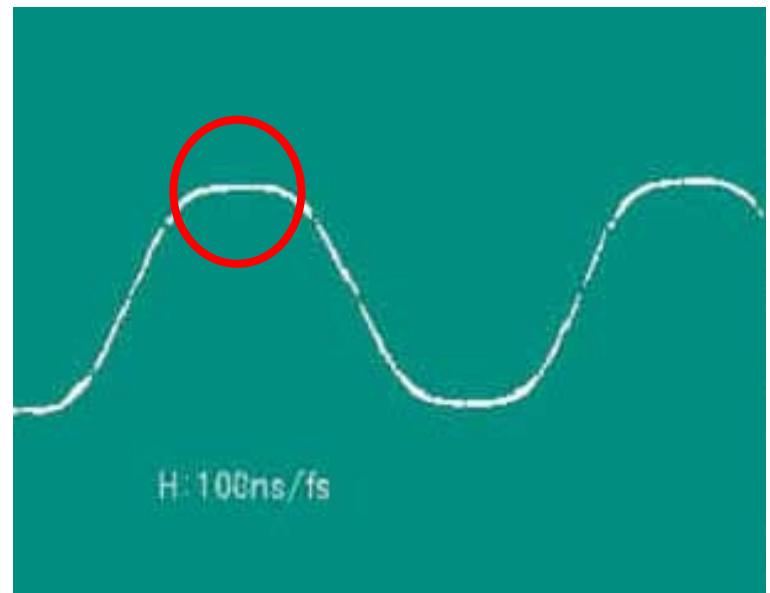
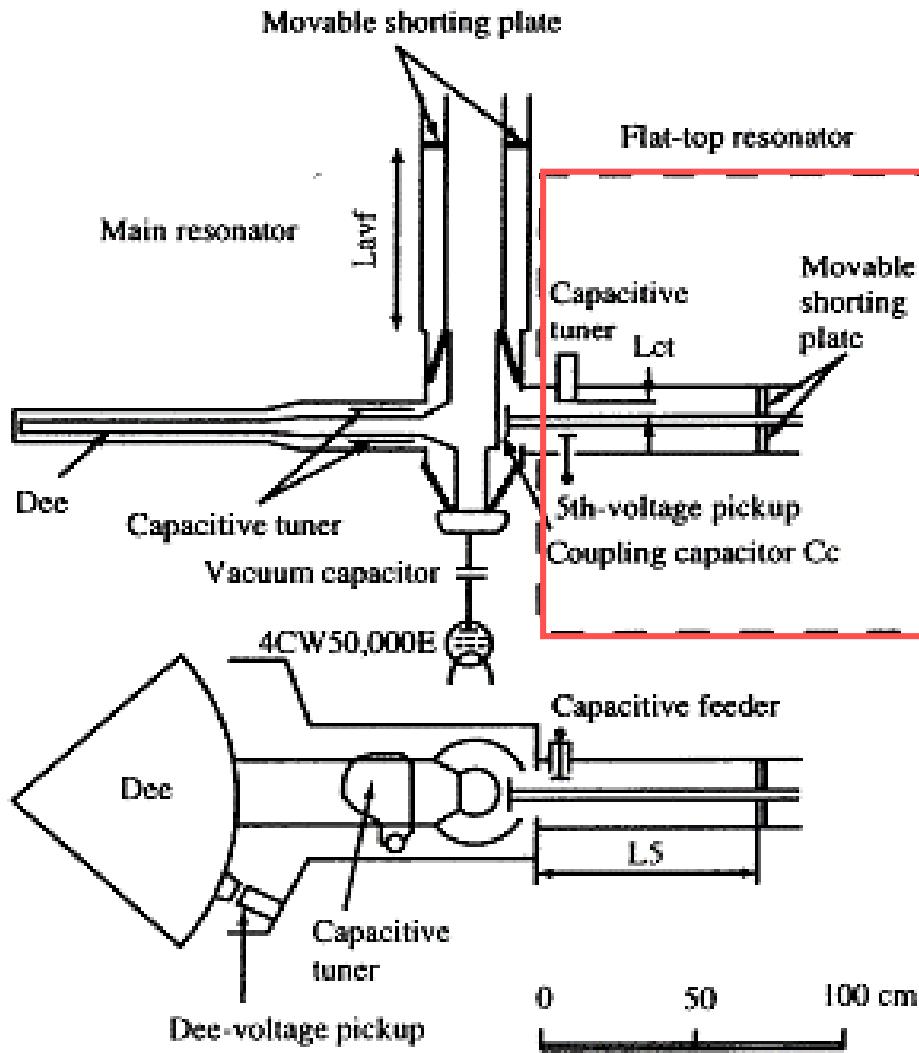
CNS-BT

- (A) Intense Heavy Ion Source
- (B) Low-energy Secondary Beam Separator CRIB
- (C) AVF Beamline
- (D) Magnetic Spectrograph
- (E) Facility of Application and Educational Experiments
- (F) Gradeup of AVF Cyclotron (plan)
- (G) Ge ball for in-beam spectroscopy (plan)

AVF Cyclotron and the ECRs'



Flat-top Acceleration with AVF



By fundamental + third harmonic

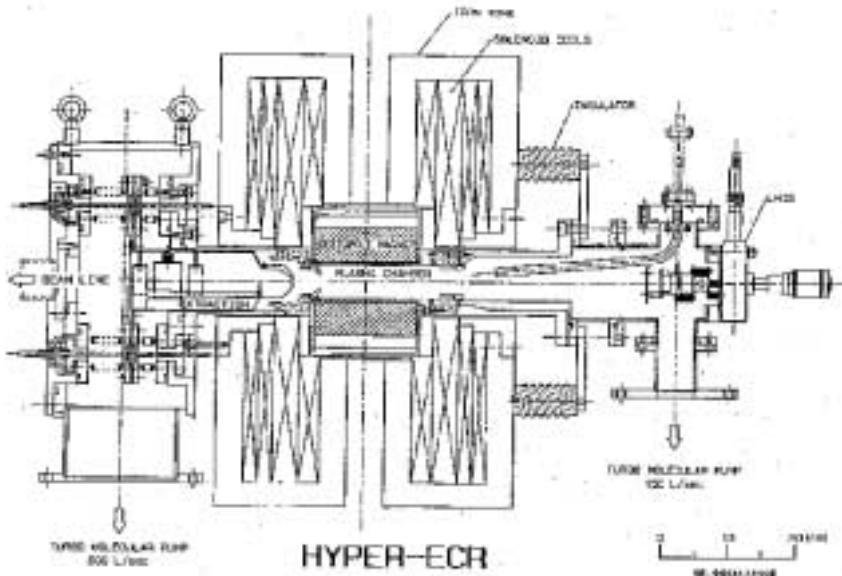
Hyper-ECR

CNS ECR Ion Source (14 GHz)



HyperECR の主なパラメーター

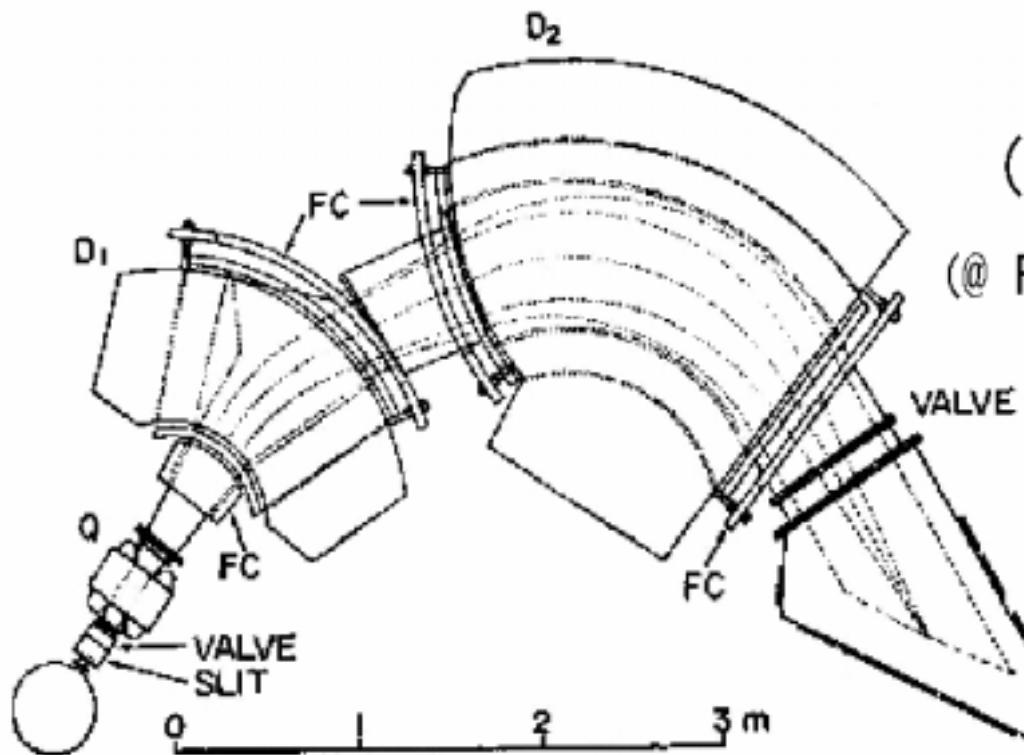
Microwave power source	
frequency	14.25 GHz
max.power	2.0 kW
Plasma chamber	
diameter	50 mm
length	190 mm
Multipole magnet	
multipolarity	Sextupole
Field strength on the surface	10.6 kG
material	Nd-Fe-B
inner diameter	57 mm
length	150 mm
Mirror field	
Max.field strength on axis	12 kG
max.Current	600 A
max.Power	72 kW
Turbo-molecular pumps	500 l/sec
	150 l/sec



Ion Intensities (eμA)

	14N	16O	20Ne	40Ar
4+	315 eμA		240	
5+	800	500 eμA		
6+	57	490	115	
7+		60	65	
8+			32	480
9+			1.5	220
11+				4.3
12+				1.5
13+				1.1
14+				2.2

High-resolution Magnetic Spectrograph PA



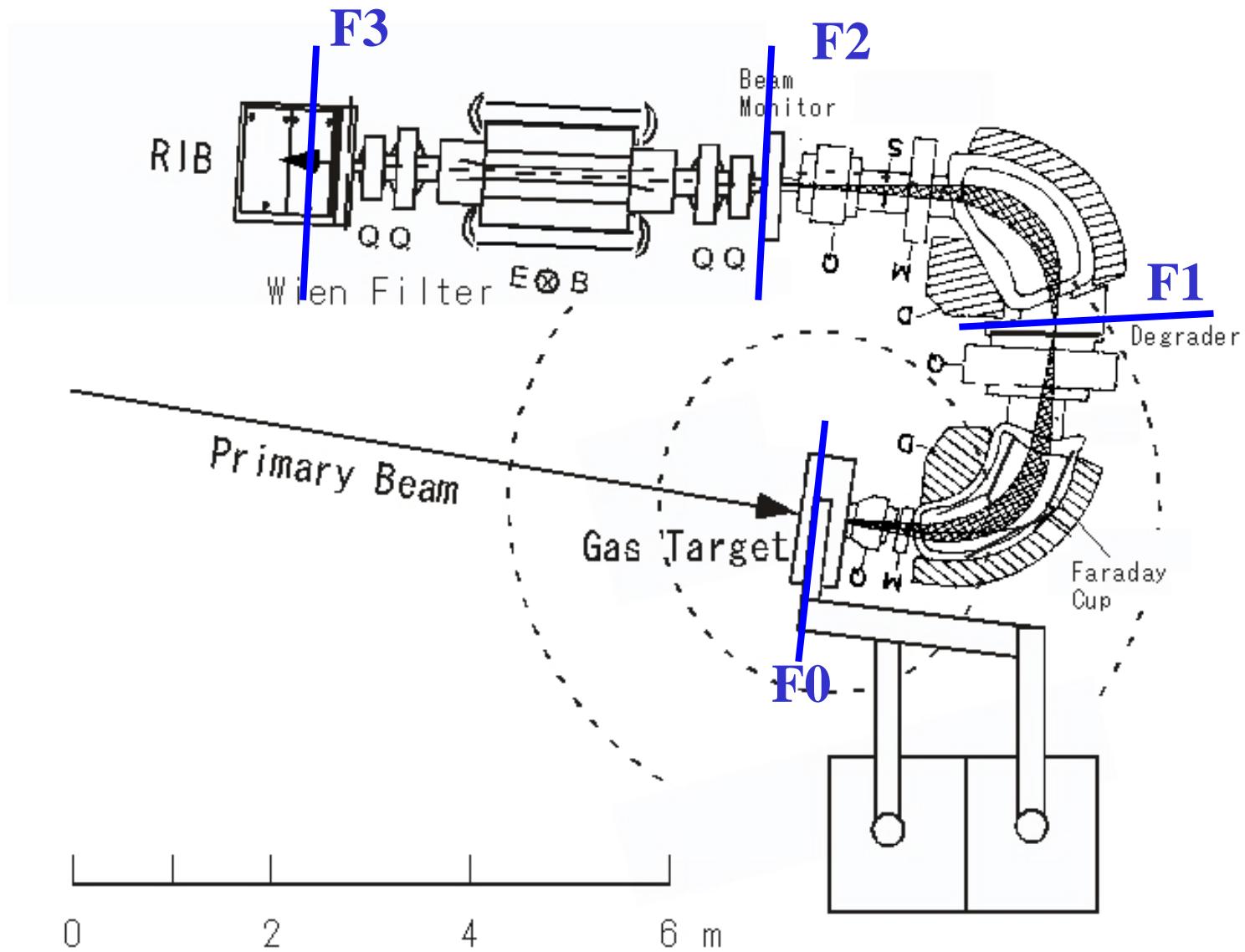
($K = 160$)

(@ RIKEN E2 hall)

性能表 Specifications

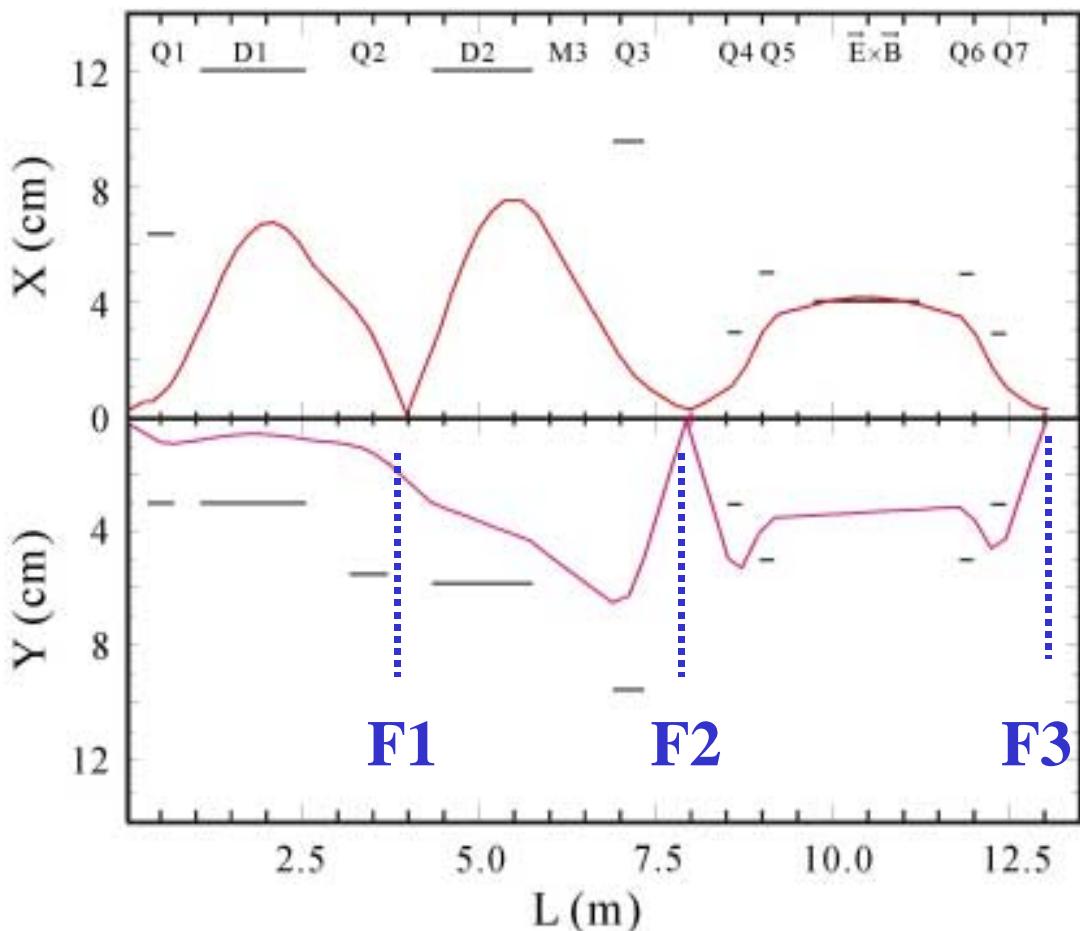
軌道半径 Radius	130 - 150 cm
分析エネルギー範囲 Energy range	30 %
測定角度範囲 Angular range	-20 ~ +135 度
立体角 Solid angle	6.4 msr
横倍率 Horiz. magnification	- 0.37
縦倍率 Vertical magnification	- 4.44
運動量分解能 Momentum resolution	0.01 %
最大電流(双極電磁石) Max. current (Dipole)	880 A
(四極電磁石) (quadrupole)	570 A
総重量 Weight	55 tons

CNS RIB Separator (CRIB)



Beam Property of CRIB

$\Delta X = \Delta Y = 2\text{mm}$
 $\Delta\theta = \Delta\phi = 20\text{ mrad}$



Specifications

Solid Angle $5.6\text{ msr}(75 \times 75\text{ mr}^2)$
Max. Mag. Rigidity 1.28 Tm
Radius of Central Orbit 0.9 m

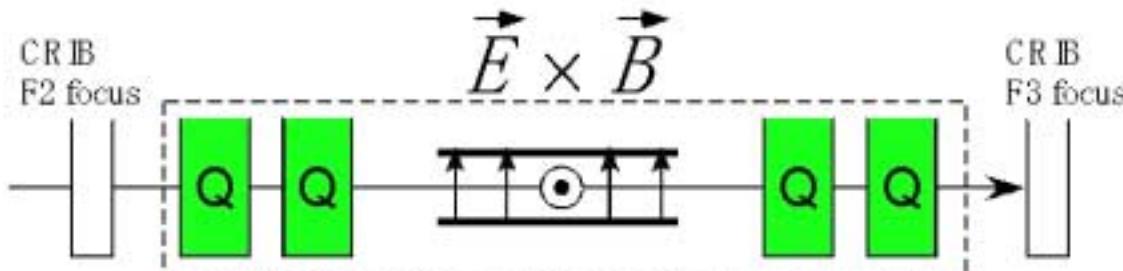
F0 → F1 (dispersive Focal Plane)
Magnification of X 0.3
Momentum dispersion 1.6 m
Momentum acceptance $\pm 7.5\%$
Momentum resolution $P/\Delta p=800$

F0 → F2 (achromatic focal plane)
Magnification of X 1.2
Magnification of Y 0.5
Momentum dispersion 0.0 m

Wien Filter

The last section of CR IB
Velocity filter by $E \times B$

Construction finished
HV test in progress



Total length: 5 m

Length ($E \times B$ part): 1.5 m

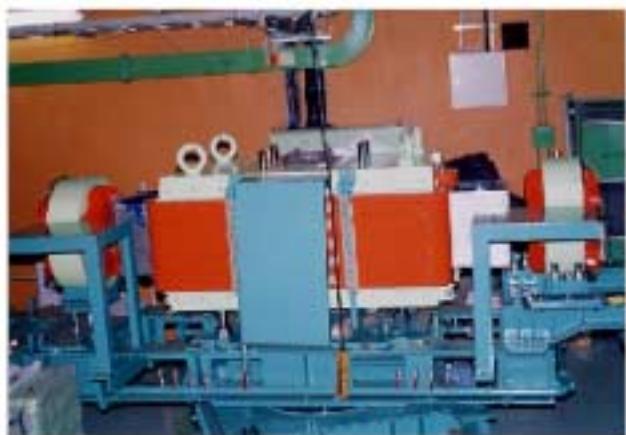
E_{max} : 50 kV/cm (gap 8 cm)

B_{max} : 3 kG gauss

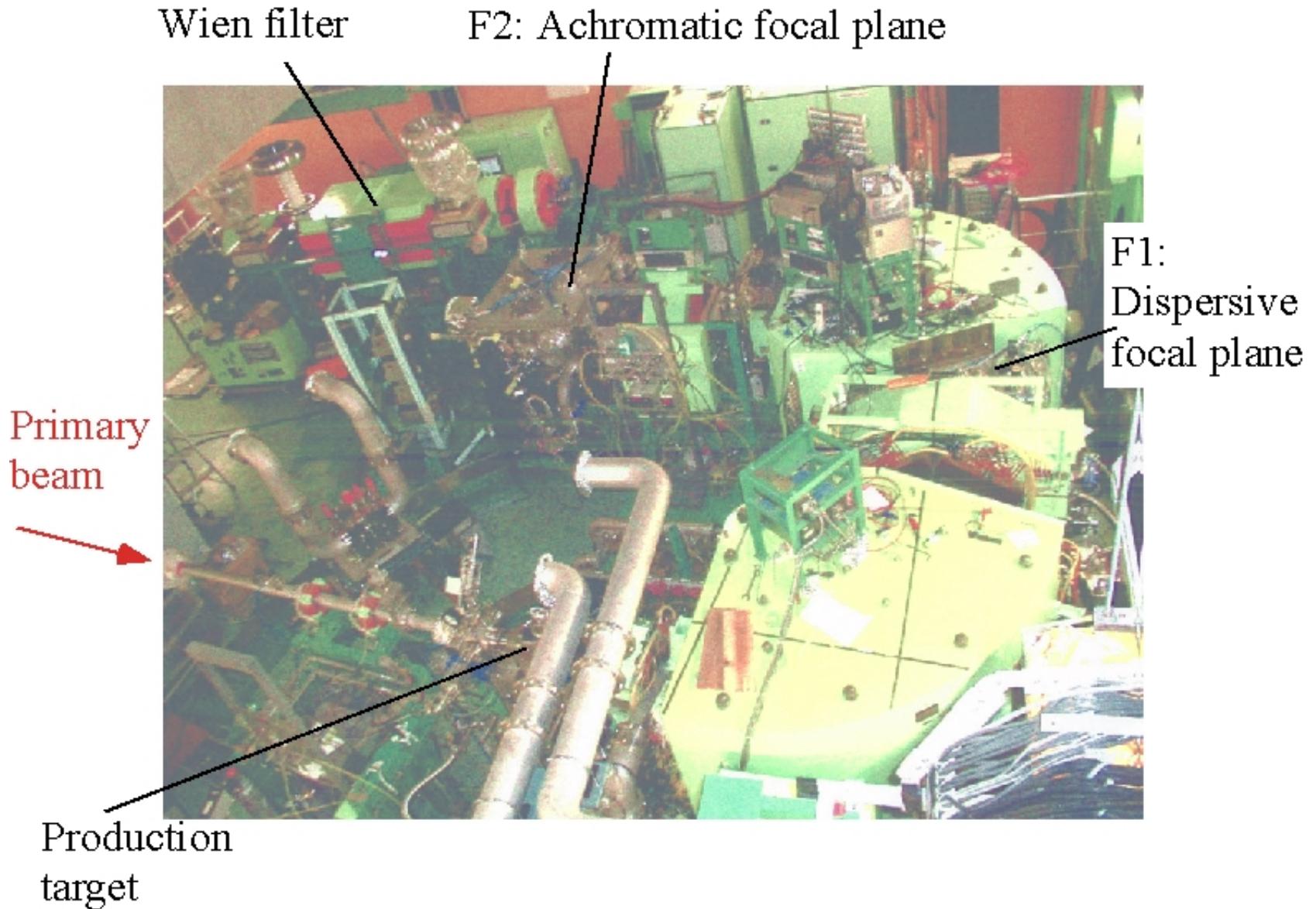
Magnification: 1 (both H & V)

Velocity resolving power:

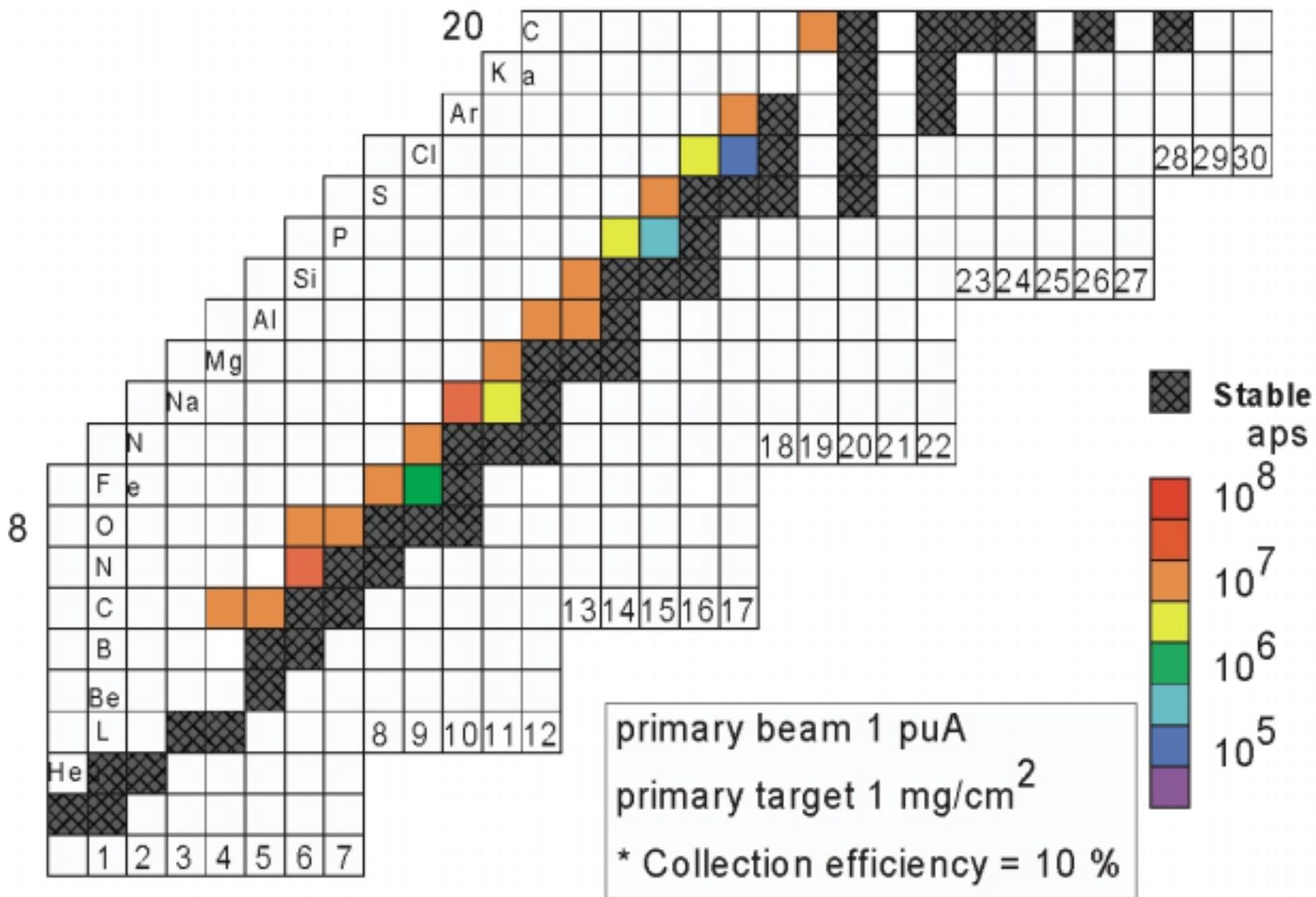
$$\Delta V/V = 2 \% \quad \text{for } A/Q = 2, E/A = 10 \text{ MeV}$$



CNS RIB Separator (CRIB)



Estimated RIB Production Rates with CRIB



Test result of Low-Energy RIB Productions

Used the (p,n) & $(^3He,n)$ reactions in inverse kinematics.
Measured at F2.

^{13}O	^{14}O	^{15}O	^{16}O
^{12}N	^{13}N	^{14}N	^{15}N
^{10}C	^{11}C	^{12}C	^{13}C
^{10}B	^{11}B		

RI beam	Primary beam	Reaction	Cross section	Target	Collection efficiency	Intensity	Purity with degrader
^{10}C 6.1 A MeV	$^{10}B(4+)$ 7.8 A MeV (200 pnA)	$p(^{10}B, ^{10}C)n$	2 mb	CH_4 gas 1.3 mg/cm ²	30 %	$(1.6 \times 10^5 \text{ cps})$	90 %
^{14}O 6.7 A MeV	$^{14}N(6+)$ 8.4 A MeV (500 pnA)	$p(^{14}N, ^{14}O)n$	8 mb	CH_4 gas 1.3 mg/cm ²	50 %	$(1.7 \times 10^6 \text{ cps})$	80 %
^{12}N 3.9 A MeV	$^{10}B(4+)$ 7.8 A MeV 200 pnA	$^3He(^{10}B, ^{12}N)n$	5 mb	3He gas 0.25 mg/cm ²	1 %	$2.5 \times 10^3 \text{ cps}$	3 %
^{11}C 3.4 A MeV	$^{10}B(4+)$ 7.8 A MeV 200 pnA	$^3He(^{10}B, ^{12}N^*)n$ $^{12}N^* \rightarrow ^{11}C + p$	≈ 20 mb	3He gas 0.25 mg/cm ²	≈ 2 %	$1.6 \times 10^4 \text{ cps}$	15 %

$^{17}N, ^{22}Mg > 10^4$ aps, ~ 10 %

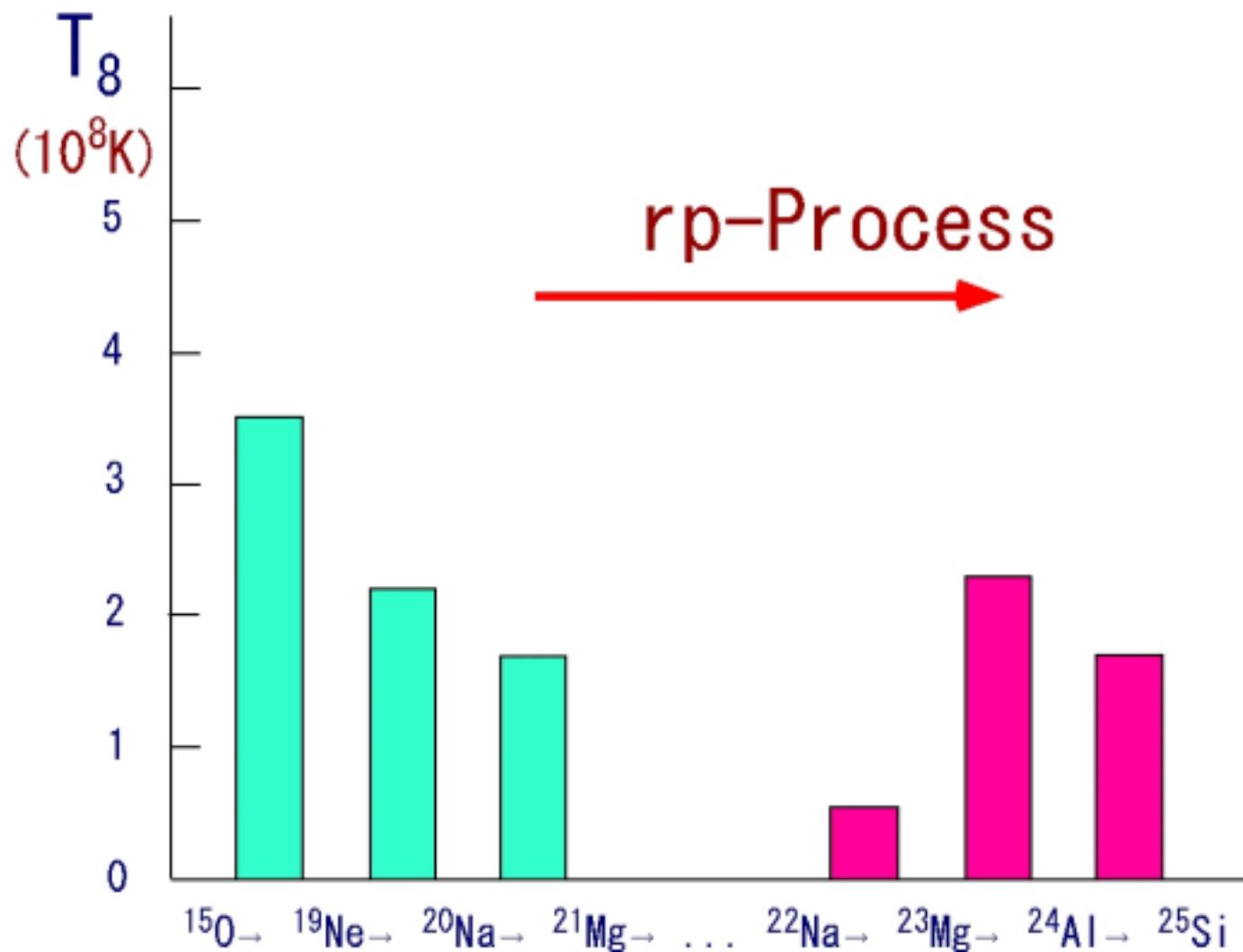
* (); Actual production tests of ^{10}C & ^{14}O were performed at lower intensities.

* Cross-section values are taken from other exp. results.

Nuclear Astrophysics Programs at CNS

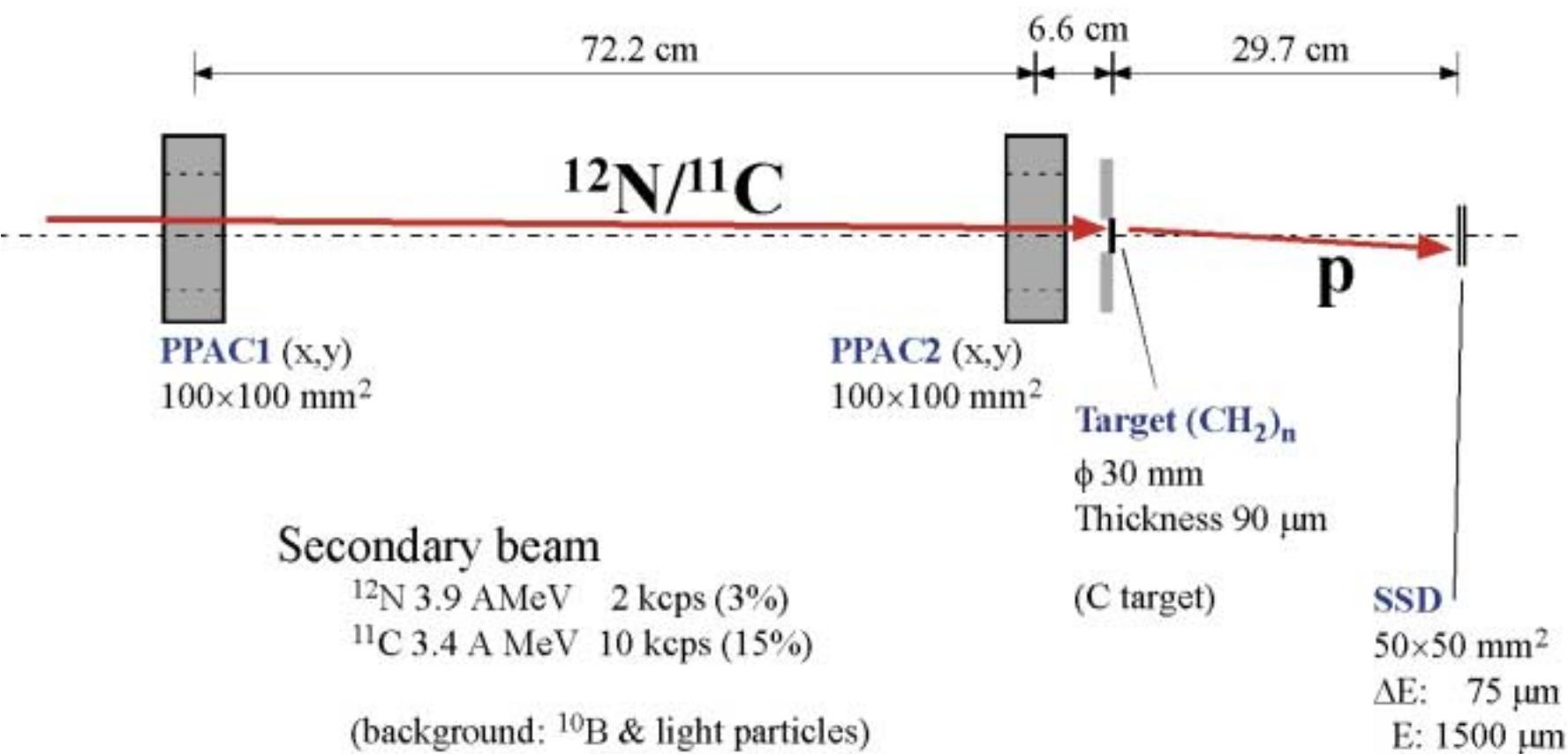
- Primordial Nucleosynthesis
- Hydrogen Burning
 - pp-chain
 - CNO cycle
 - rp-process
- (n,γ) Reaction Study

Ignition Temperature

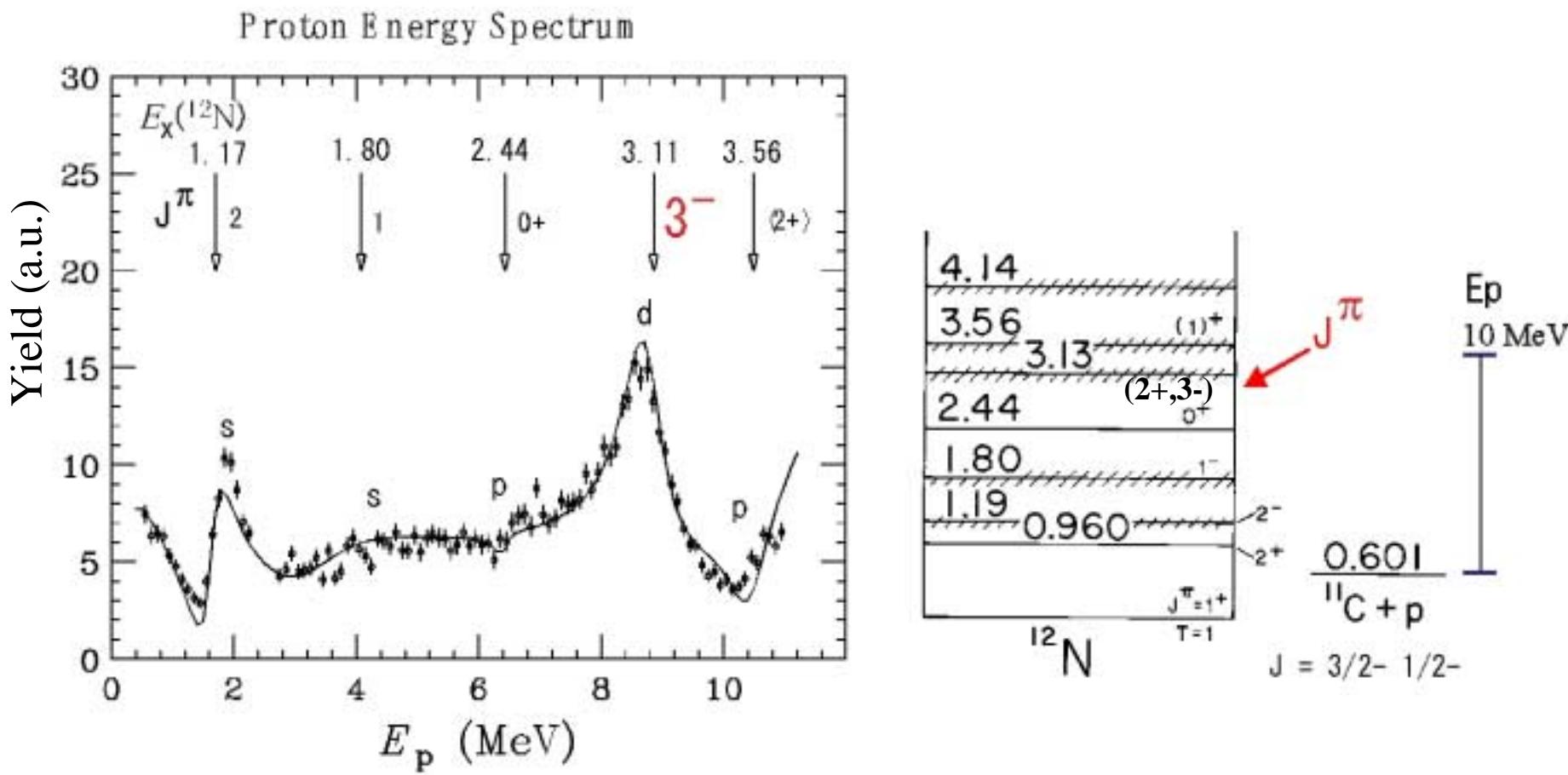


Setup for $^{12}\text{N}+\text{p}$ & $^{11}\text{C}+\text{p}$ elastic resonance scattering

at CRIB F2

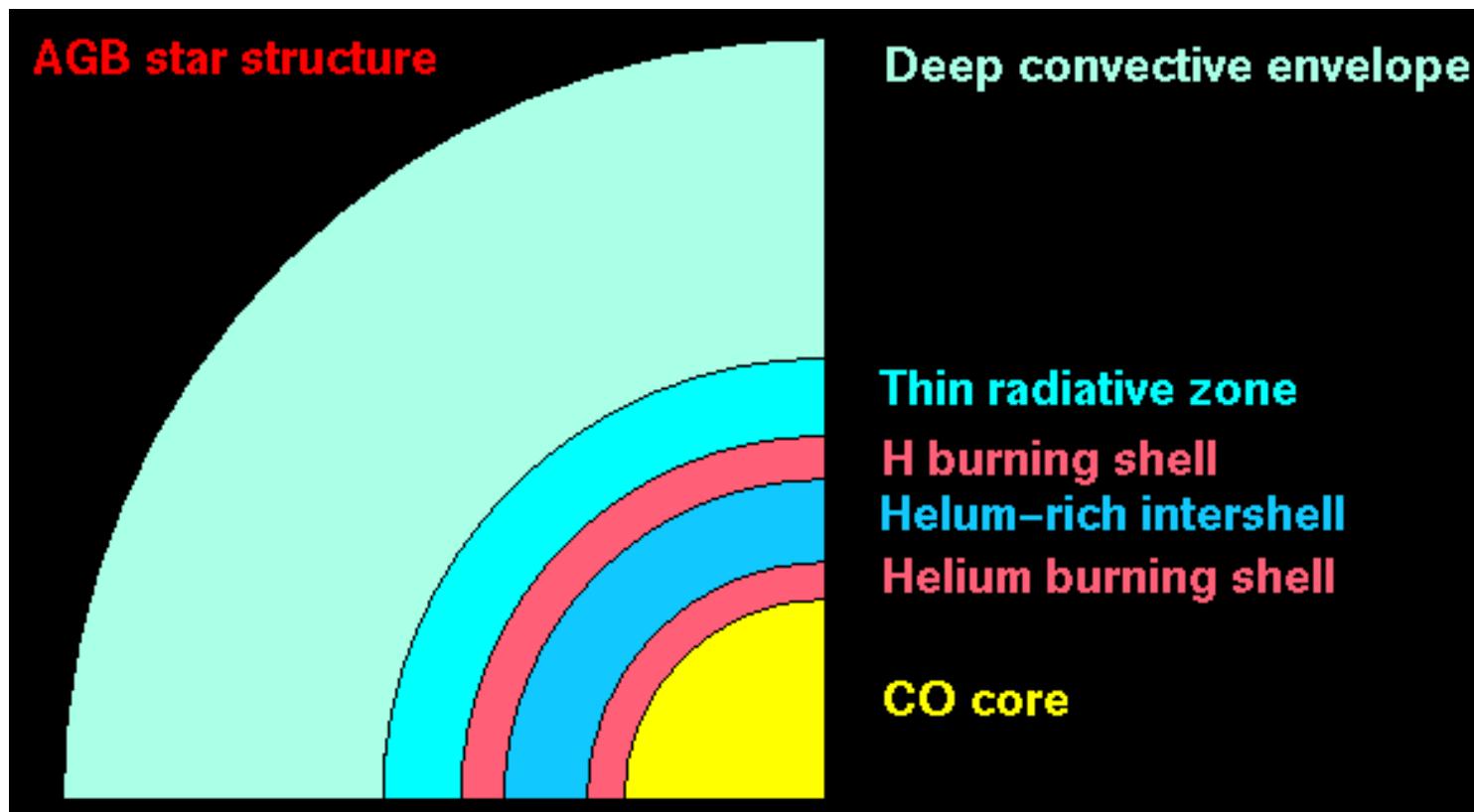


Low-Energy Resonant Elastic Scattering of $^{11}\text{C} + \text{p}$

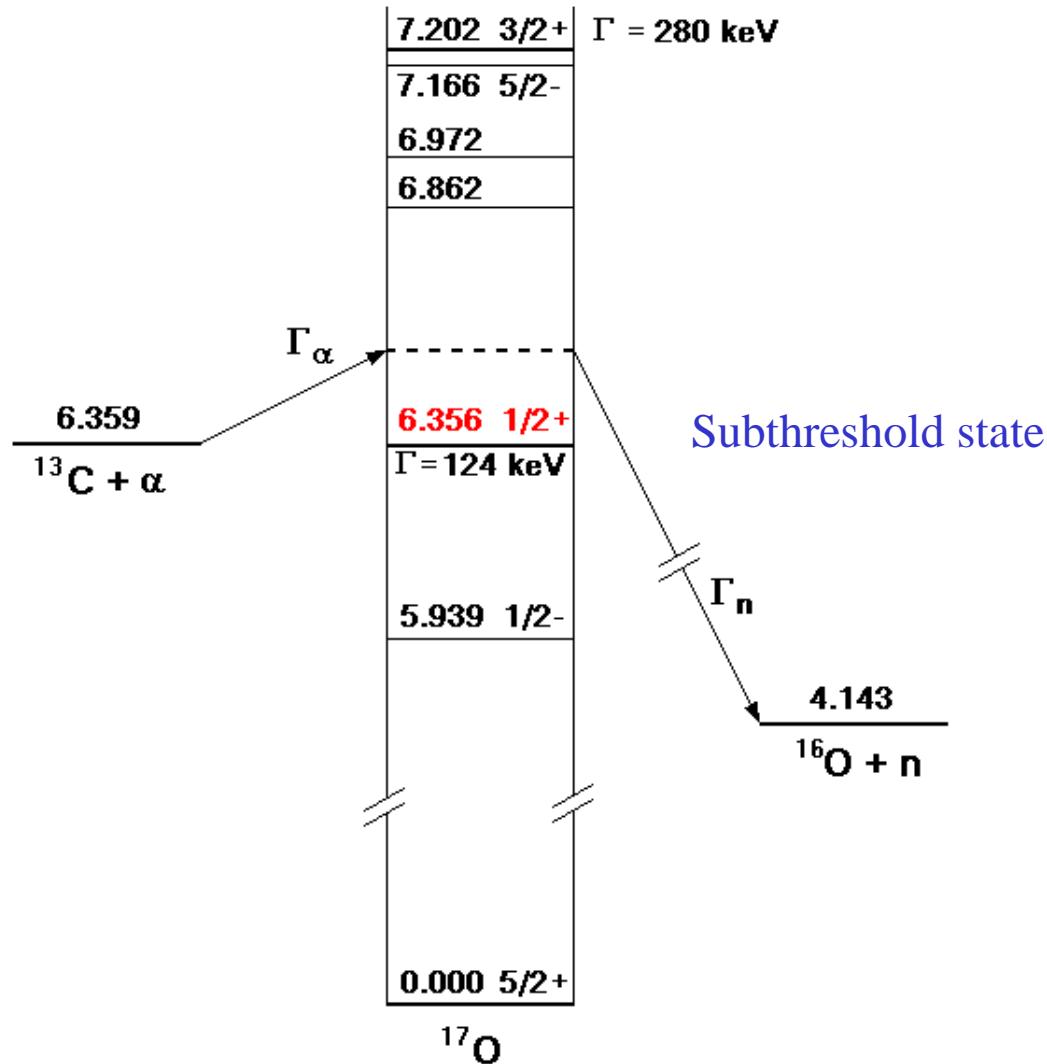


$^{13}\text{C}(\alpha, \text{n})^{16}\text{O}$ Reaction is the Main Neutron Source for the s-Process ?

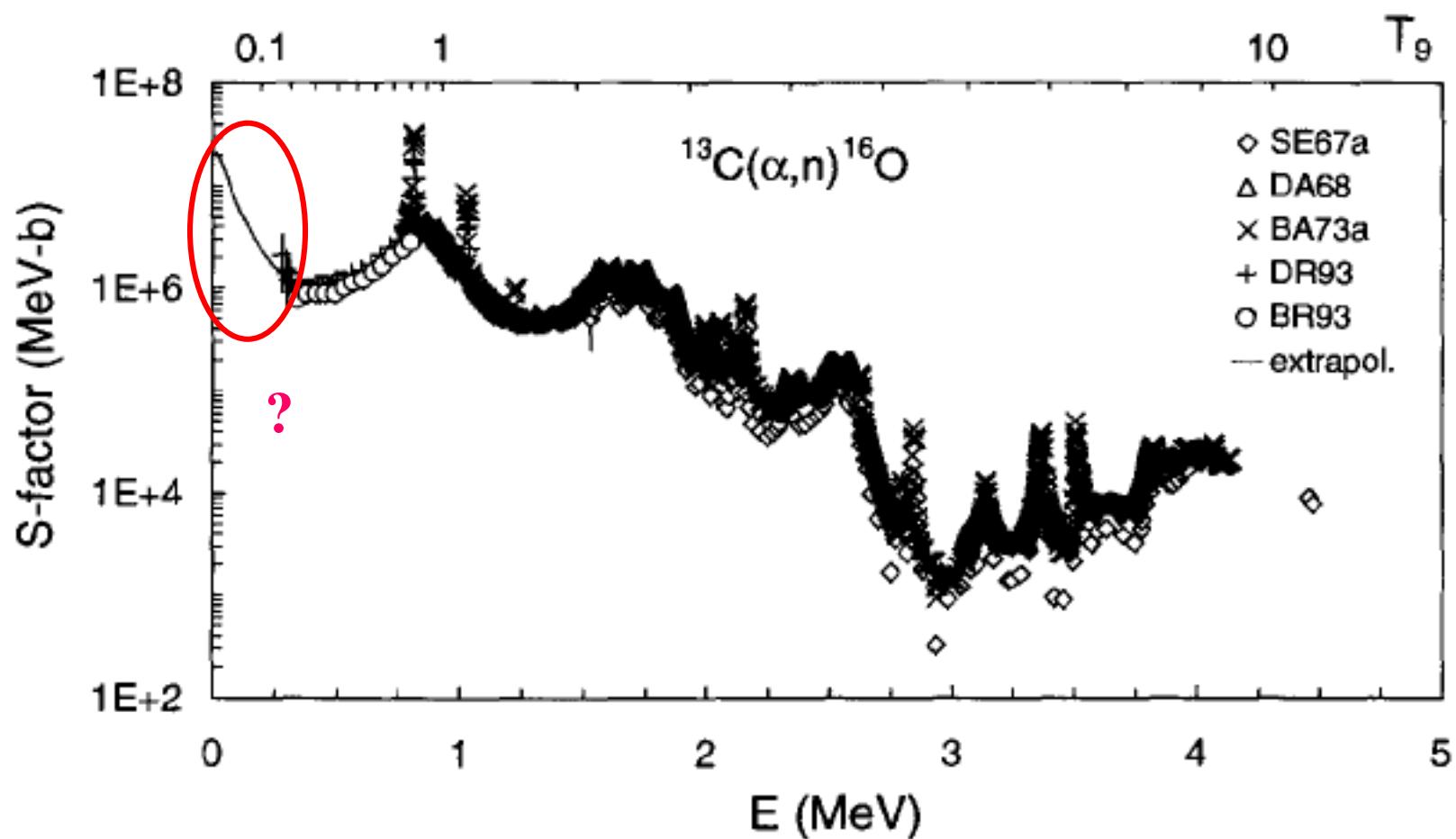
- The $^{13}\text{C}(\alpha, \text{n})^{16}\text{O}$ reaction on He-burning shell (~ 0.1 GK)
- The $^{12}\text{C}(\text{p}, \gamma)^{13}\text{N}(\beta^+)^{13}\text{C}$ reaction on H burning shell
- Low/intermediate-mass AGB stars



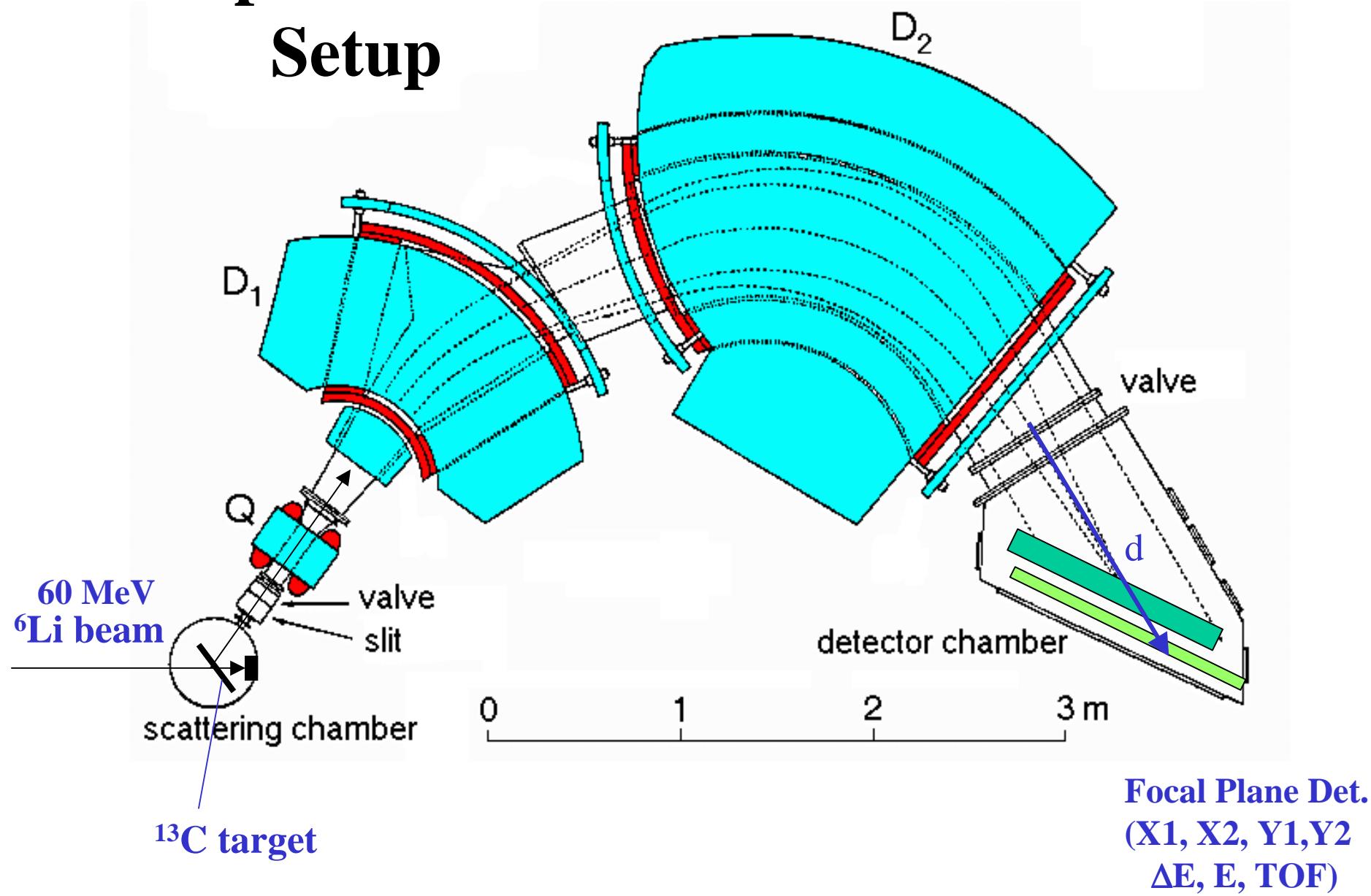
Levels in ^{17}O



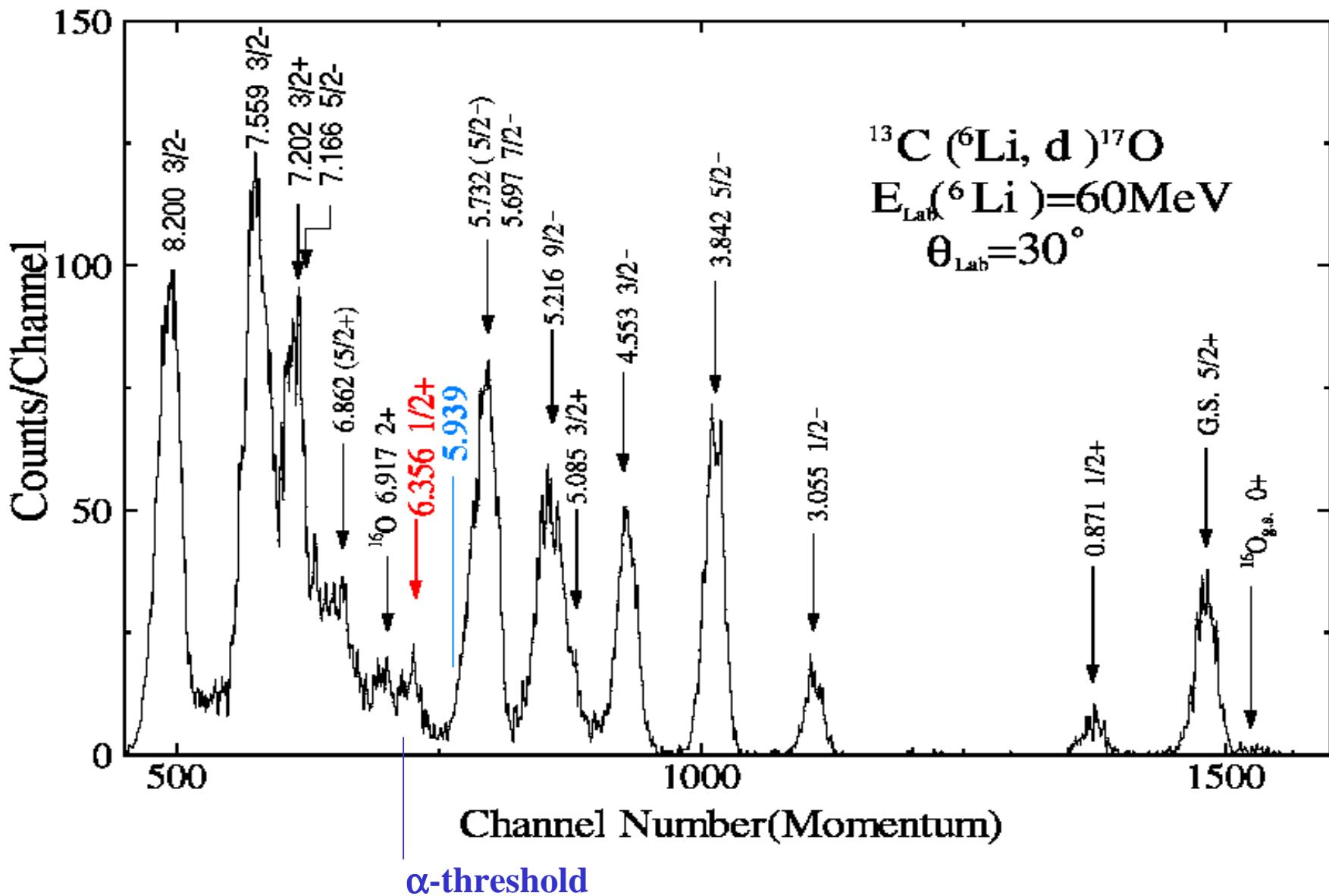
NACRE collaboration (Angulo *et al*, Nucl. Phys. A656(1999)3)



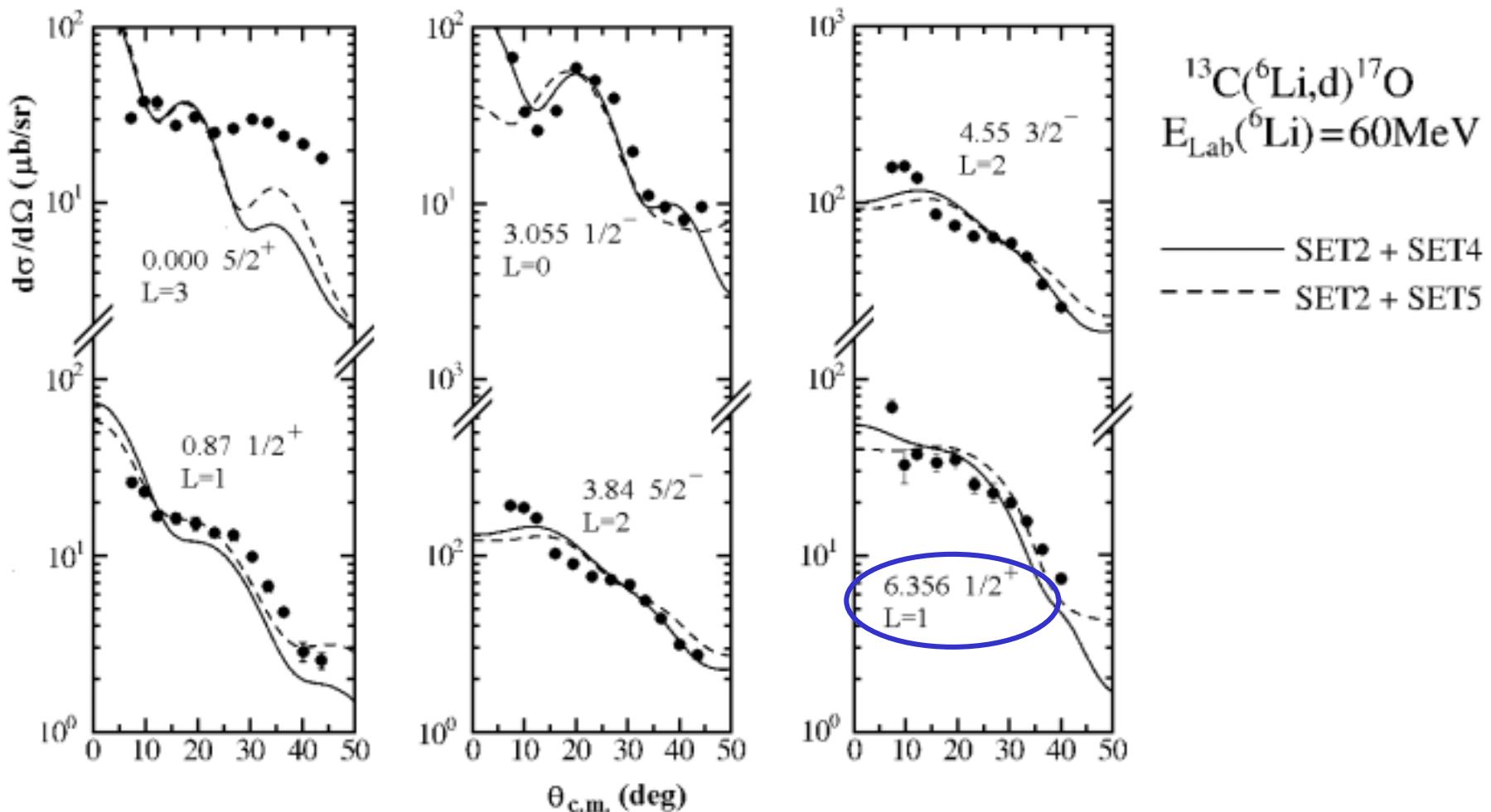
Experimental Setup



$^{13}\text{C}({}^6\text{Li}, \text{d})^{17}\text{O}$

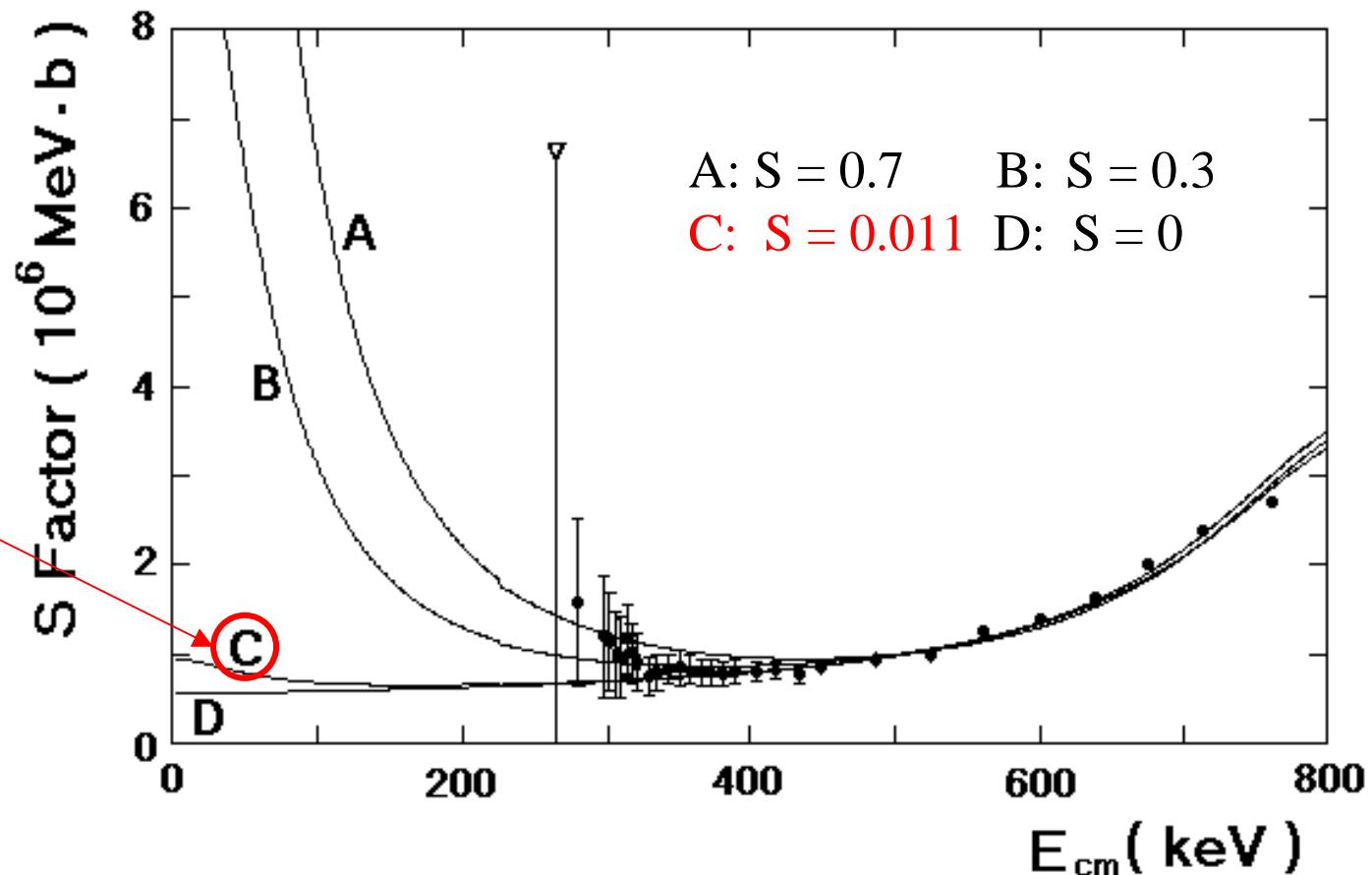


Angular Distributions of $^{13}\text{C}({}^6\text{Li},\text{d}){}^{17}\text{O}$



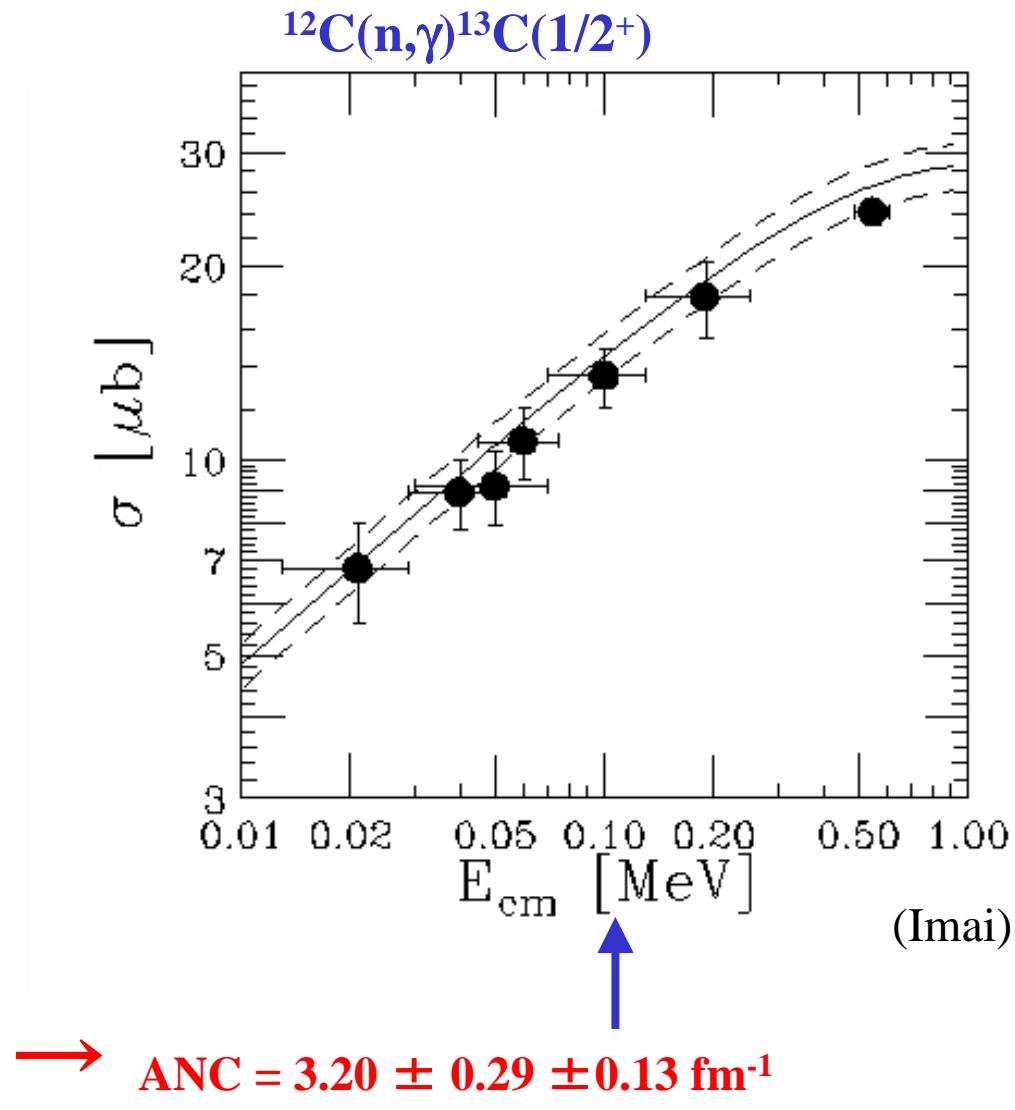
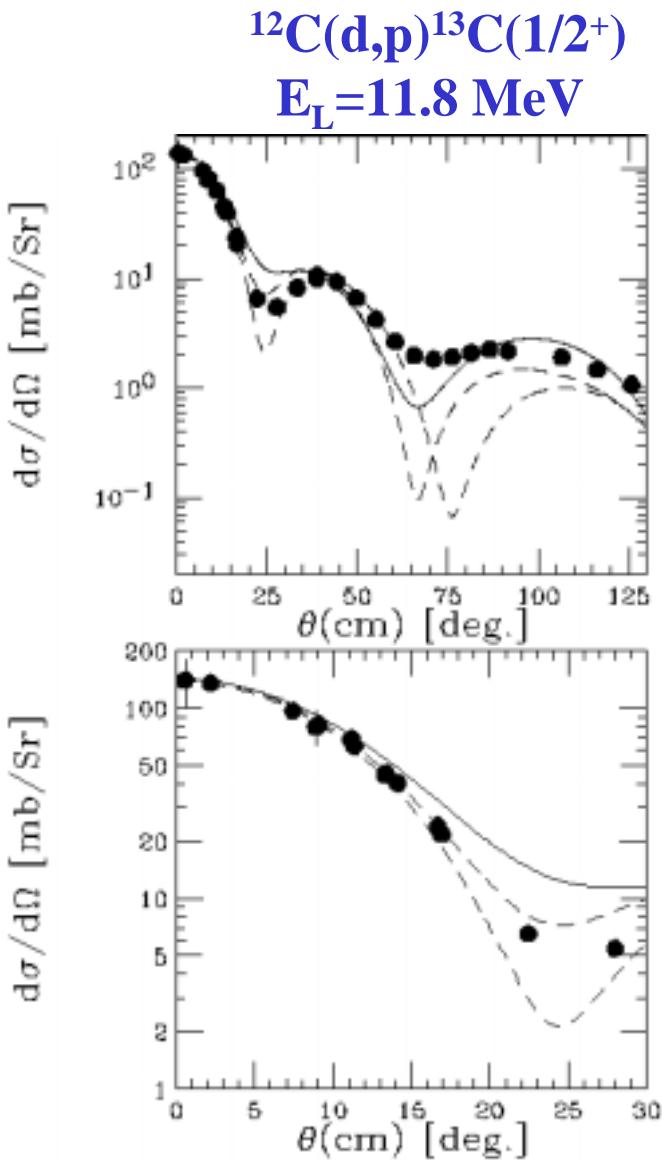
$^{13}\text{C}(\alpha, \text{n})^{16}\text{O}$ Reaction

Present
Result



- The role of the sub-threshold state was found to be very small in the s-process nucleosynthesis !

The $^{12}\text{C}(\text{d},\text{p})$ reaction for the $^{12}\text{C}(\text{n},\gamma)^{13}\text{C}$ stellar reaction



Uncertainties of ANC Method with (d,p) reactions for (n, γ) reactions

1. Choice of optical potential ~ 10 %
 ** Potentials that fit a large angular range give less.
2. Choice of bound state potential ~ 2 %
3. Interior contribution ~ 6 %
4. Breakup effect ~ 4 %

5. Need transfer data of small uncertainty at very
 forward angles. ~ exp %

total > 12 %

Summary

- 1) Many interesting research opportunities under the CNS-RIKEN collaboration.
- 2) Low-energy RIB separator CRIB works well for physics programs.
- 3) Explosive nucleosynthesis (novae, supernovae, etc.) can be investigated.
- 4) Investigation of the r-process nuclei is of great interest.

Please come to CNS for collaboration

- We welcome all of you to come back to CNS for collaboration works.
- We encourage especially young people to come to study a new idea on Physics.