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Evolution of Shell Structure and Spin-Isospin Interaction

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Single-particle (or shell) structure is the basis of many nuclear properties such as (sub)magic number, deformation, and even the existence of the nucleus.

The single-particle levels in exotic nuclei can be different from those of stable nuclei due to the following aspects:

Loose bindingWoods-Saxon potentialNeutron skinMean-Field modelsNN interaction, particularly, spin-isospin interactions
(also in nuclei not very close to drip lines)

Major subject :

Such shell evolution due to

tensor and 2-body LS interactions

Something simple enough that everyone can remember

- **1. Shell evolution due to Tensor interaction**
- 2. Tensor effect in shell-model effective interaction
- 3. Tensor-implemented mean field model
- 4. Shell evolution due to 2-body LS interaction
- 5. Example : (partial) origin of carbon mystery
- 6. Summary
- 7. Exercise

Effective single particle energy

Monopole part of the NN interaction

$$V_{ab}^{T} = \frac{\sum_{J} (2J+1) V_{abab}^{JT}}{\sum_{J} (2J+1)}$$

Angular averaged interaction

➡ spherical single particle energies

Effective single-particle energy (ESPE)



Shift of single-particle energies due to interaction with other valence nucleons

Major origin of shell evolution due to spin-isospin *NN* interaction :

tensor interaction



 π meson : dominant source

ρ meson (~ π + π) : minor cancellation for smaller *r*

Important for binding *e.g.* B.S. Pudliner et al., Phys. Rev. C56, 1720 (1997)

Has never shown up directly (or in the first order) in nuclear spectroscopy (e.g. levels, etc)

second-order effect on spin-orbit splitting

T. Terasawa, Prog. Theor. Phys. 23, 87 (1960);

A. Arima and T. Terasawa, Prog. Theor. Phys. 23, 87 (1960)



Major features Opposite signs \implies **spin-orbit splitting varied** T=0: T=1 = 3:1 (same sign) **Only exchange terms** neutron, j'_{\sim} proton, $j_>$ tensor proton, neutron, j' <





Proton effective single-particle levels (relative to $d_{3/2}$)



Systematic variation of proton effective single-particle energies due to the tensor interaction (π + ρ meson) *calculation only*



Systematic variation of neutron effective single-particle energies due to the tensor interaction (π + ρ meson)





Exp. data from J.P. Schiffer et al., Phys. Rev. Lett. 92, 162501 (2004) Also, C. Baktash, Paestum talk.

How the tensor interaction is included in effective shell model interaction ?

1. pf shell GXPF1

(2. sd shell SDPF-M (USD+corr))

3. p shell SFO (CK+corr)

pf shell

GXPF1 interaction M. Honma et al., PRC65 (2002) 061301(R)

G-matrix + polarization correction | + | empirical refinement

Modify realistic G interaction

M. Hjorth-Jensen, et al., Phys. Repts. 261 (1995) 125

- Bonn-C potential *Tensor is here*
- 3rd order Q-box + folded diagram
- Vary 70 well-determined LC's of 195 TBME and 4 SPE
- Fit to 699 experimental energy data of 87 nuclei
- $V(A) = V(A = 42) \times \left(\frac{A}{42}\right)^{-0.3}$ Mass dependence ۲
- Data selection to avoid intruder: $47 \le A$, $Z \le 32$
- Energy evaluation by $FDA^* \longrightarrow 168 \text{keV}$ rms error

G-matrix vs. GXPF1

two-body matrix element <ab; JT / V / cd; JT > $7=f_{7/2}, 3=p_{3/2}, 5=f_{5/2}, 1=p_{1/2}$

- T=0 ... attractive
- T=1 ... repulsive
- Relatively large modifications in V(abab; J0) with large J V(aabb; J1) pairing



Monopole part of various interactions





Monopole interaction after subtraction of tensor part

GXPF1 - Tensor between same *l*



Monopole interaction in p-shell



Tensor interaction is the primary origin of the p-n $j_>-j_<$ coupling also within a major shell (of a fixed parity).



Otsuka et al. Phys. Rev. Lett. 87, 082502 (2001)

Tensor interaction is the primary origin of the σσττ effective interaction within a major shell (of a fixed parity).

> This can be explained analytically at the long-range limit of of the tensor interaction. (A good exercise of Racah algebra)

Tensor interaction is included in G-matrix-type calculations.

Implementation of tensor interaction into mean field calculations

Gogny interaction

 $\begin{array}{ll} (1+\sigma\sigma+\tau\tau+\sigma\sigma\tau\tau) \left(Gauss1+Gauss2\right)+Density \ Dep.\\ finite \ range & zero \ range\\ \hline Tensor \ interaction \ is \ added \end{array}$

All parameters are readjusted

Nuclear matter properties reproduced with improvement of imcompressibility

Gogny-Tokyo interaction - 2 (GT2)

Tensor interaction actually used



Regularized for short distance

Neutron effective single-particle energies of exotic Ni isotopes

Original (D1S)







Proton effective single-particle energies of exotic Ni isotopes

Original (D1S)

GT2 (incl. tensor)



Another origin of shell evolution:

2-body LS interaction



I ntuitive Picture of 2-body LS

Orbital angular momentum of relative motion







other combinations : higher partial waves



2-body *LS* interaction is one of the origins of spin-orbit splitting

It shifts the position of *s* orbit as the inverse mechanism.

In fact, the amount of the effect is largest for *s* orbit due to ³P dominance.

A systematic shell model calculation for Carbon isotopes

Millener-Kurath + modifications





Effective single-particle energies for protons



E2 transition $0^+ \rightarrow 2^+$ in ${}^{16}C$

Exp. : *B*(E2)=3.2 (6) e² fm⁴ Imai et al., Phys. Rev. Lett. (03)

Shell model $e_p = 1.3$ $e_n = 0.5$ $B(E2) = 40 e^2 fm^4$ $e_p = 1.2$ $e_n = 0.1 - 0.2$ $B(E2) = 12 e^2 fm^4$ Sagawa and Asahi, Phys. Rev. C63, 064310 (01) **Summary**

respects

Shell evolution due to spin-isospin interactions

Tensor interaction (long range)

drives $j_{>}$ or $j_{<}$ levels in a specific way This is not necessarily a change of spin-orbit splitting.

is the dominant origin of shell evolution

produces effects of similar magnitude to *neutron skin* (weakening of *ls* splitting)

2-body LS interaction (short range)

special cases (*e.g.* between *s* and *p*)

carbon-oxygen $d_{5/2}$ - $s_{1/2}$ inversion (same mechanism as ||s splitting)

Structure of exotic nuclei in many



Exercise:

Find a combination of three orbits and a way to see the variation of their energies.

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