

# 超新星爆発に関する光核反応の実験的研究

## Experimental Study of Photonuclear Reactions Relevant to Supernova Explosions

Tatsushi Shima

*Research Center for Nuclear Physics, Osaka University*

- Photonuclear reactions related to SNe
- Laser Compton-scattered  $\gamma$ -ray source
- Recent experiment; ex.  ${}^4\text{He}$
- Summary & Outlook

# Photonuclear Reactions Related to Supernova Explosions

- He-burning ;  $^{12}\text{C}(\alpha, \gamma)^{16}\text{O} \leftrightarrow ^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$  (detailed balance)
- Photodisintegration of iron core ;  
 $(\gamma, \alpha)$ ,  $^4\text{He}(\gamma, \text{p})^3\text{H}$ ,  $(\gamma, \text{n})^2\text{H}$ ,  $^4\text{He}(\gamma, \text{n})^3\text{He}$ ,  $(\gamma, \text{p})^2\text{H}$ ,  $(\gamma, \text{n})^1\text{H}$
- r-, p-,  $\gamma$ -processes ;  $(\text{p}, \gamma)$ ,  $(\text{n}, \gamma)$ ,  $(\gamma, \text{n})$ ,  $(\gamma, \text{p})$ ,  $(\gamma, \alpha)$
- $\nu$ -process ;  $(\nu, \nu'x)$  (neutral current)  $\leftrightarrow (\gamma, x)$   
--- nuclear responses to dipole excitation by  $\tau_3$ ,  $\sigma \cdot \tau_3$

# Photonuclear reactions

## as a tool for nuclear astrophysics

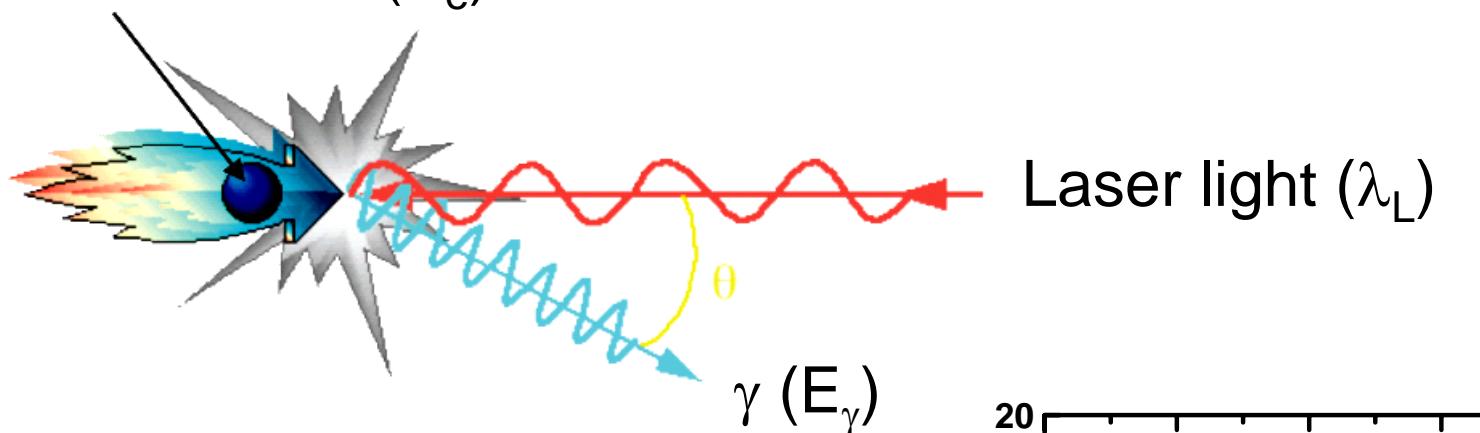
- EM interaction; **well-know !**
- Sensitivity to **multi-body channels**;  
 $X+Y+Z+\dots \rightarrow A+\gamma \Leftrightarrow A+\gamma \rightarrow X+Y+Z+\dots$   
ex.  $2\alpha+n \rightarrow {}^9\text{Be}$ ,  $3\alpha \rightarrow {}^{12}\text{C}$
- Enhancement of cross sections by  
**phase space factor**
- Analogy between  $\gamma\text{-A}$  and  $\nu\text{-A}$  interactions

# Low-energy $\gamma$ -ray sources

- Discrete  $\gamma$ -rays from radioisotopes
- Bremsstrahlung (continuous energy) / tagged photons
- $e^+e^-$  pair annihilation in flight ; monoenergy  $\gamma$  + brems.
- **Laser Compton Scattered  $\gamma$  (LCS- $\gamma$ )**
- Synchrotron photons from super-conducting wiggler
  - quasi-Planck spectrum (Utsunomiya, Nucl. Phys. A777)
- Coulomb Dissociation (CD) ; virtual photons, application to RI  
 ${}^7\text{Be}(p,\gamma){}^8\text{B}$ ,  ${}^{14}\text{C}(n,\gamma){}^{15}\text{C}$ , ...      **RIKEN, GSI**
- $(p,p')$  ; probe for M1 excitation      **RCNP**

# Laser Compton backscattering

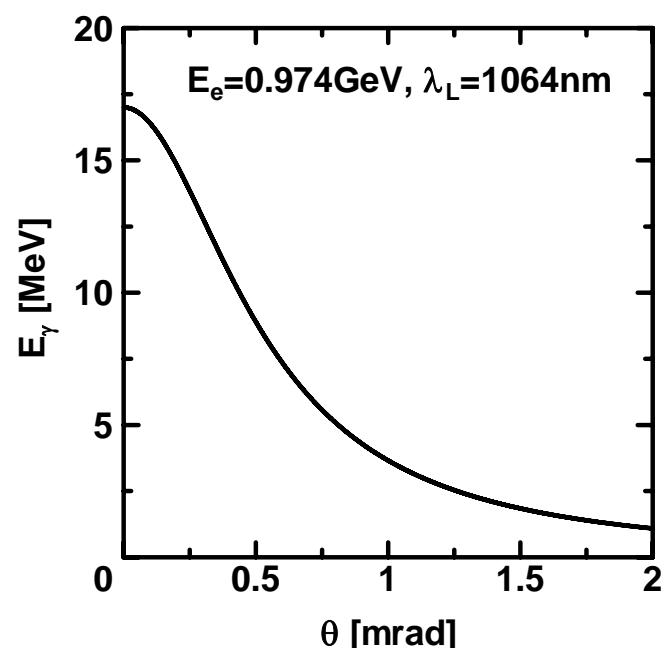
Relativistic electron ( $E_e$ )



$$E_\gamma = \frac{4hc}{\lambda_L} \cdot \frac{\gamma^2}{1 + \gamma^2 \theta^2}, \quad \gamma = \frac{E_e}{m_e c^2}$$

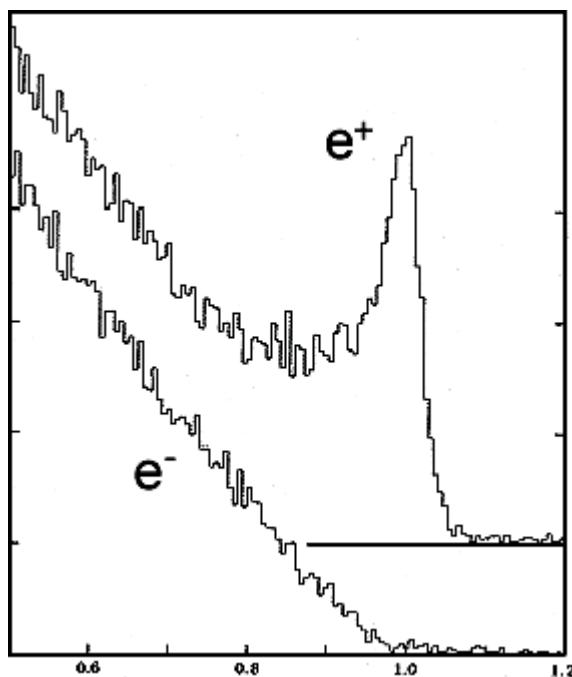
ex.  $\lambda_L=1.064\mu\text{m}$ ,  $E_e=974\text{MeV}$

$\Rightarrow E_\gamma = 17\text{MeV}$

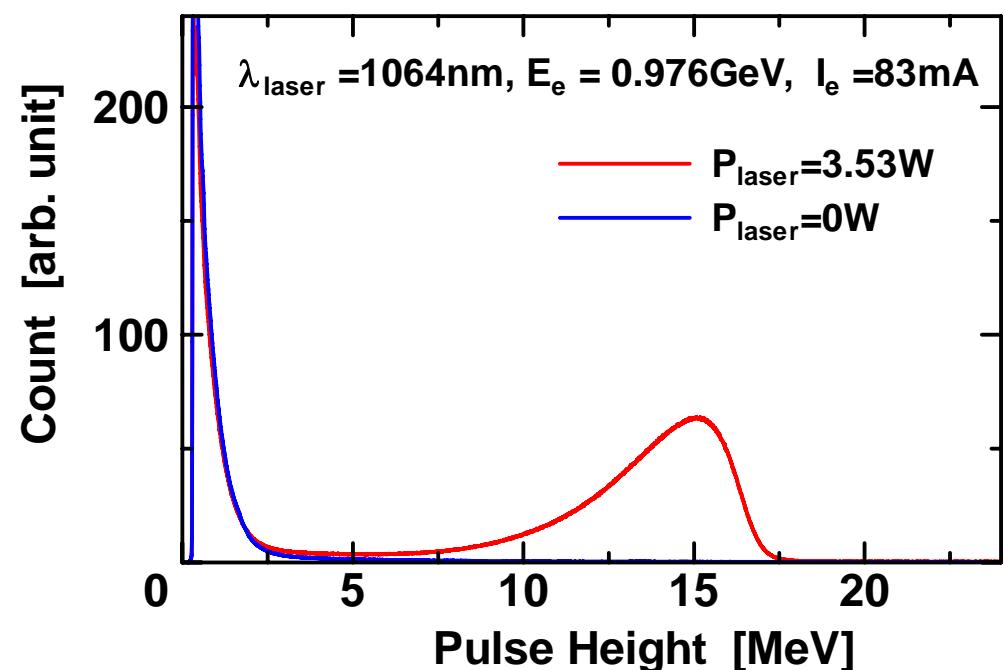


# Energy distributions

Bremsstrahlung,  
 $e^+e^-$  annihilation in flight



Laser Compton-Scattered  $\gamma$   
(PH spectra of GSO scintillator)



BG from low-energy  
component of brems.

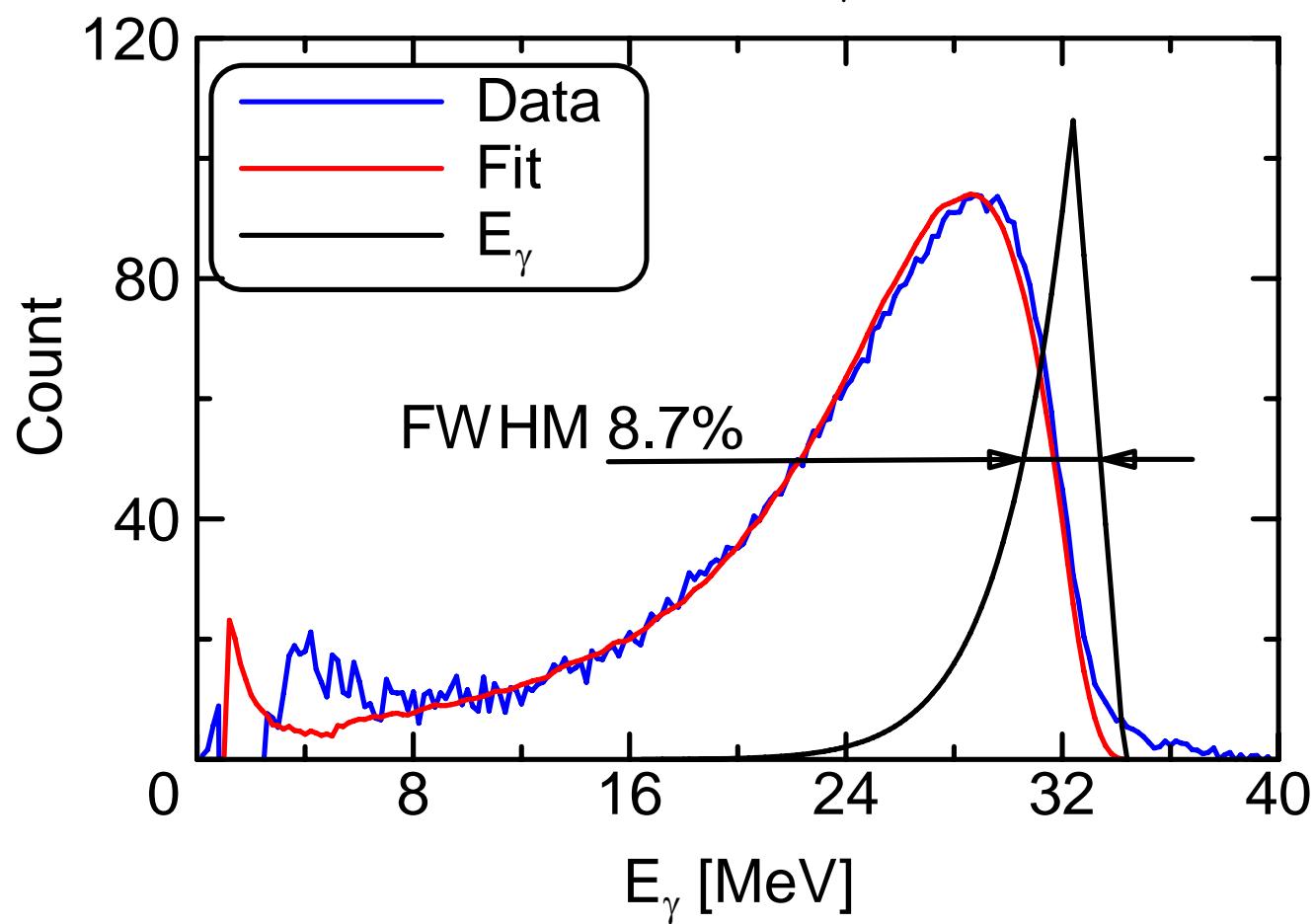
No BG !!

# Energy distribution

NaI(Tl):  $6''\phi \times 5''$

Collimator:  $\phi 3\text{mm}$

$E_\gamma(\text{max}) = 34\text{MeV}$



# Advantages of LCS- $\gamma$

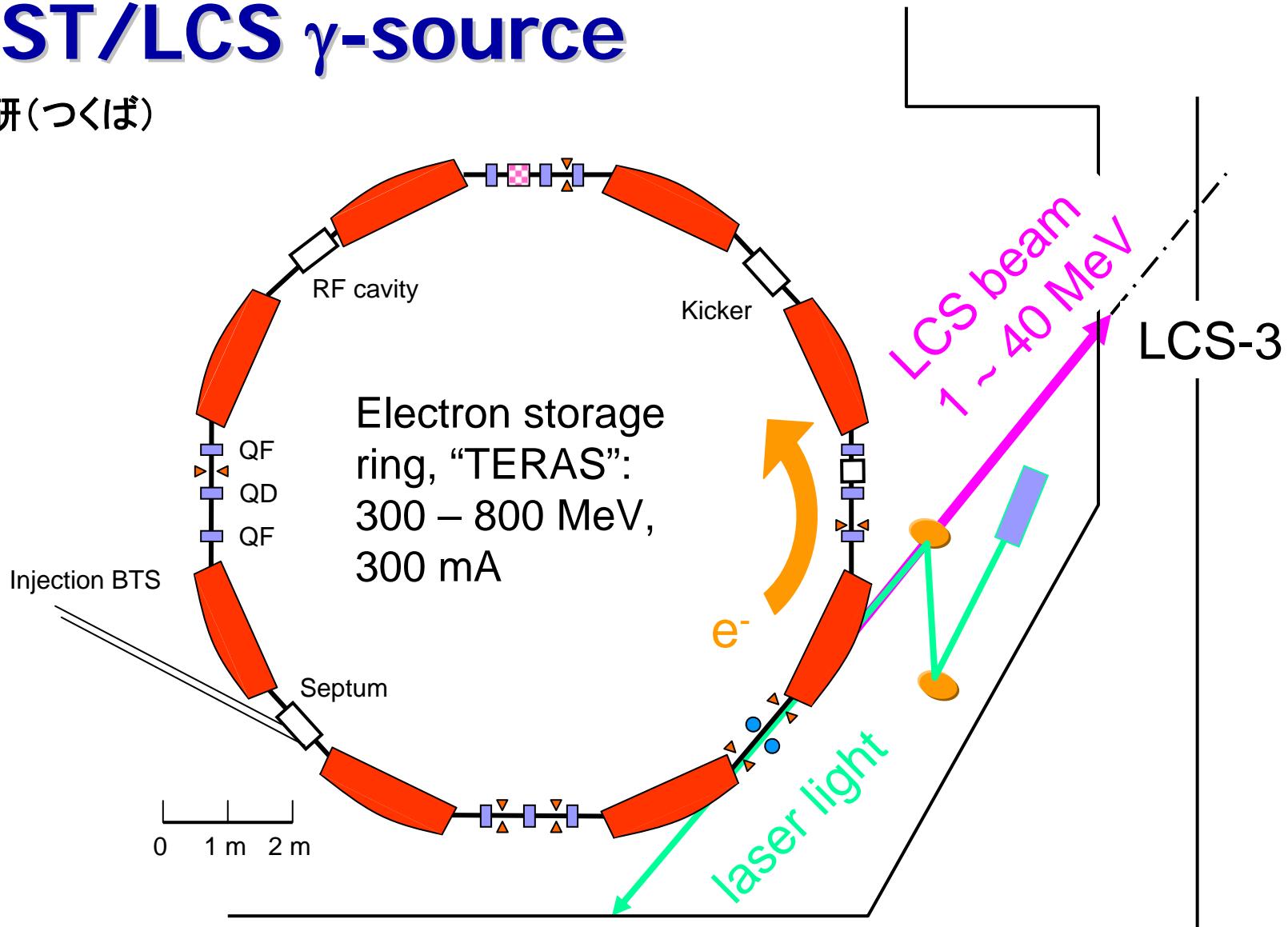
- Quasi-monochromatic;  $\Delta E/E < \sim 10\%$ , little BG
- Well-collimated;  $\Delta\theta < 0.1$  mrad
- Highly polarized; linear or circular,  $P \sim 100\%$
- Continuous or pulsed;  $\Delta t < 10\text{ns}$

# LCS $\gamma$ -ray facilities in the world

Year	Facility	$E_e$ [GeV]	Laser	$E_\gamma$ [MeV]	$\Phi_\gamma$ [/sec/MeV]
1964	Lebedev	0.6	ruby	$\sim 10$	$\sim 10^6$
1965	Harvard / CEA	6	ruby	$\sim 1000$	$5 \times 10^4$
1978	Frascati / LADON	1.5	Ar <sup>+</sup>	5 - 80	$\sim 10^5$
1985	ETL(AIST) / TERAS	0.2-0.8	Nd-YAG	2 - 40	$10^3 \sim 10^5$
1990	BNL / LEGS	2.8	Nd-YAG	150 - 470	$10^4 \sim 10^5$
1996	Grenoble / GrAAL	6	Ar <sup>+</sup>	1500	$\sim 4 \times 10^3$
1999	SPring-8 / LEPS	8	Ar <sup>+</sup>	1500 - 3500	$\sim 10^4$
1999	TUNL(Duke) / H $\gamma$ S	0.2-1.2	FEL	2 - 70	$\sim 10^7$
2004	LASTI(U. Hyogo) / NewSUBARU	1 - 1.5	Nd-YVO <sub>4</sub>	16 - 40	$10^4 \sim 10^5$

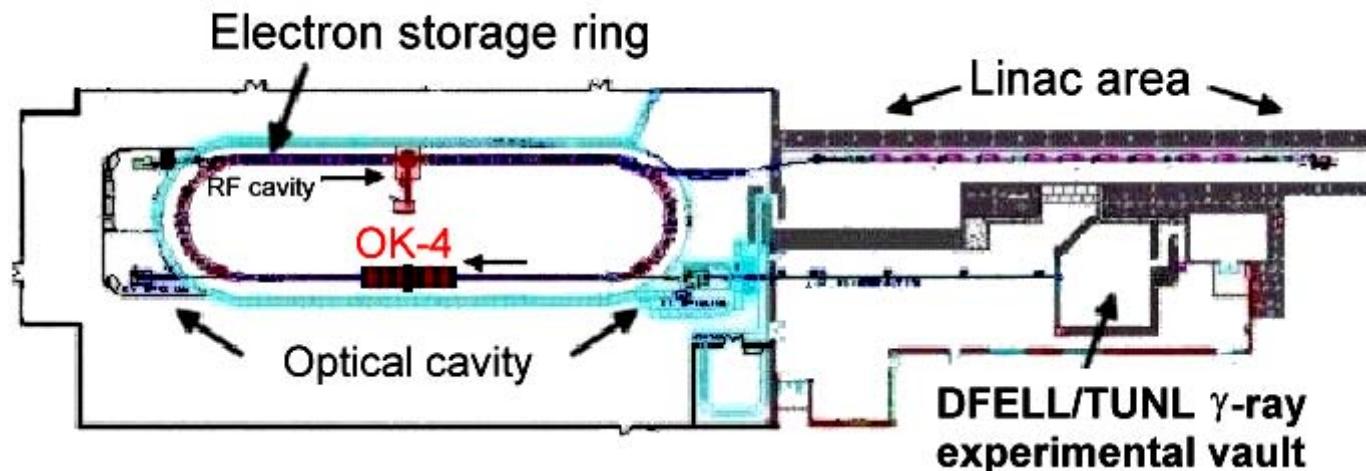
# AIST/LCS $\gamma$ -source

産総研(つくば)



# Duke/HIgS (High Intensity $\gamma$ -ray Source)

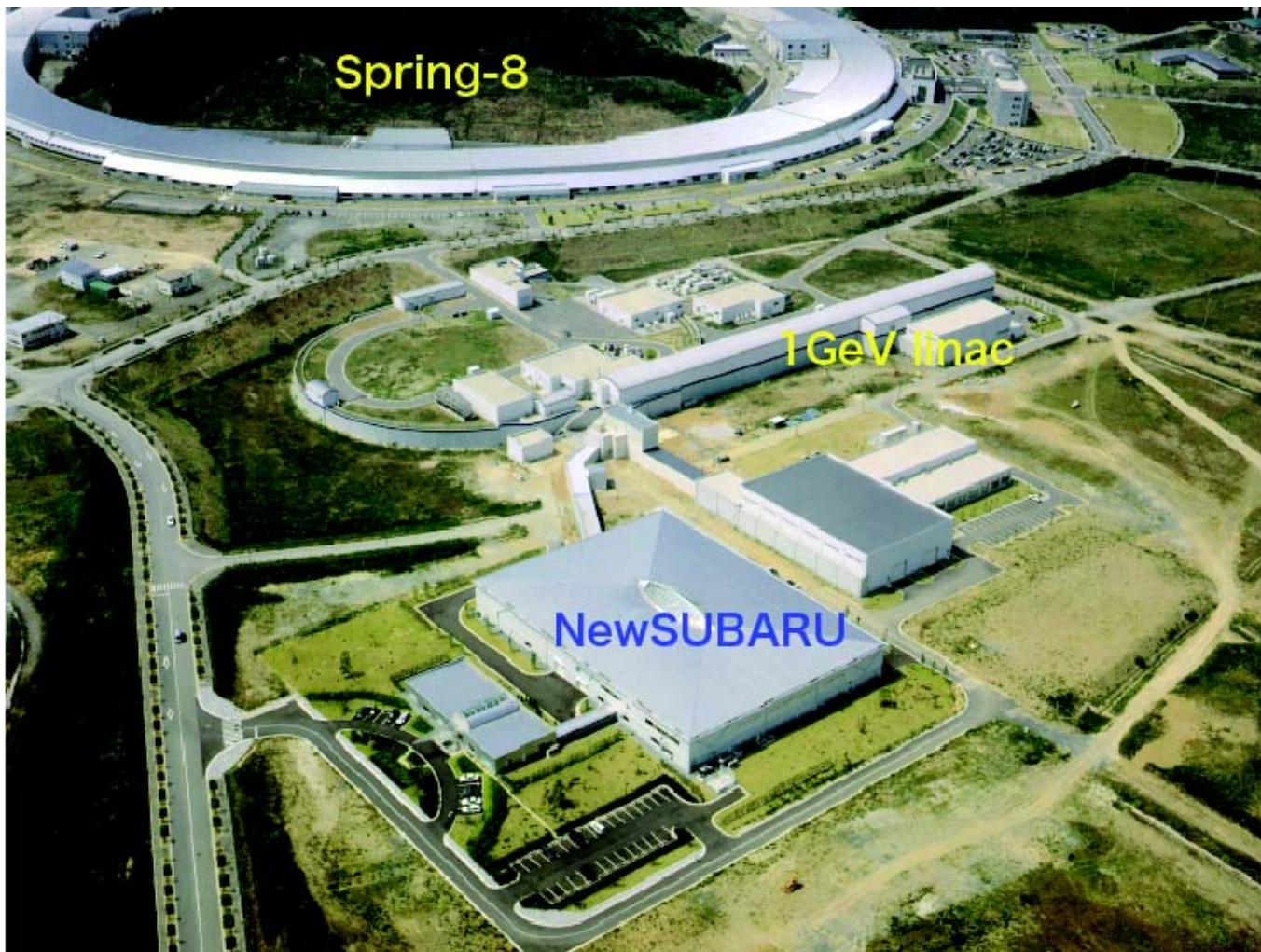
- Intra-cavity Compton Backscattering of **FEL** photons by electrons circulating in the 1.2GeV Duke Storage Ring



- $E_\gamma = 2\text{--}70 \text{ MeV}$ ,  $\Delta E_\gamma/E_\gamma \sim 1\%$ ,  $\Phi_\gamma \sim 10^7 / \text{MeV/s} (\rightarrow 10^9)$
- Time interval of  $\gamma$ -ray pulses ; 170ns

# NewSUBARU

Lab. of Adv. Sci. and Tech. for Industry,  
University of Hyogo, Japan



**Subaru** is an old Japanese meaning “get together”, and also the Japanese name of the Pleiades Star Cluster.



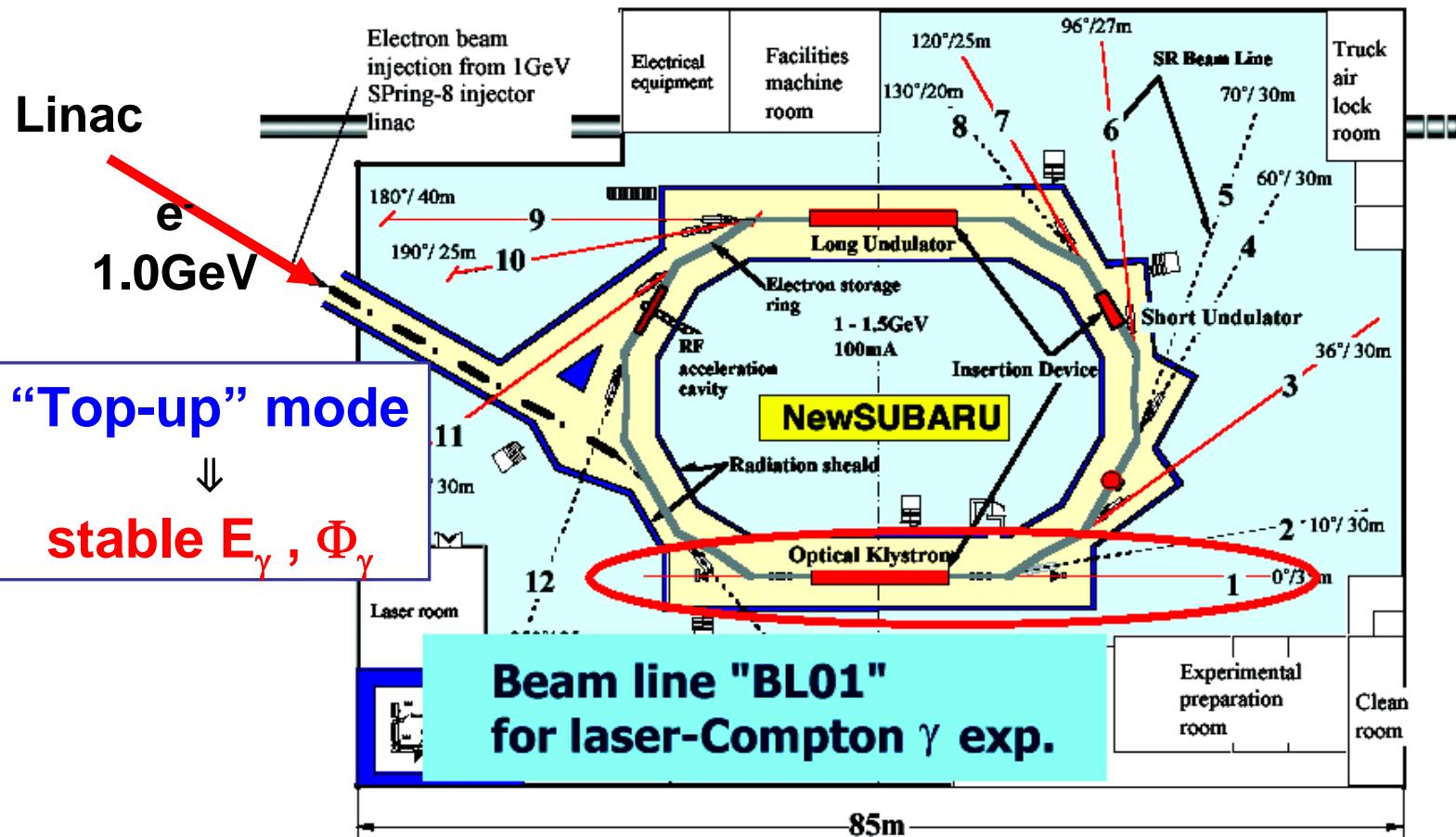
	Subaru	NewSUBARU	SUBARU
Type	Telescope	SR facility	Automaker
Location	Mauna Kea, Hawaii	Hyogo, Japan	Gunma, Japan
Feature	8.2m single mirror with adaptive optics	Race-track shape isochronous ring	Boxer engine + symmetrical AWD
Memorial	First light 1999	First beam 1998	WRC first win 1993





# NewSUBARU/LCS- $\gamma$ source

K. Aoki, S. Miyamoto, et al. NIM A516 (2004) 228-236



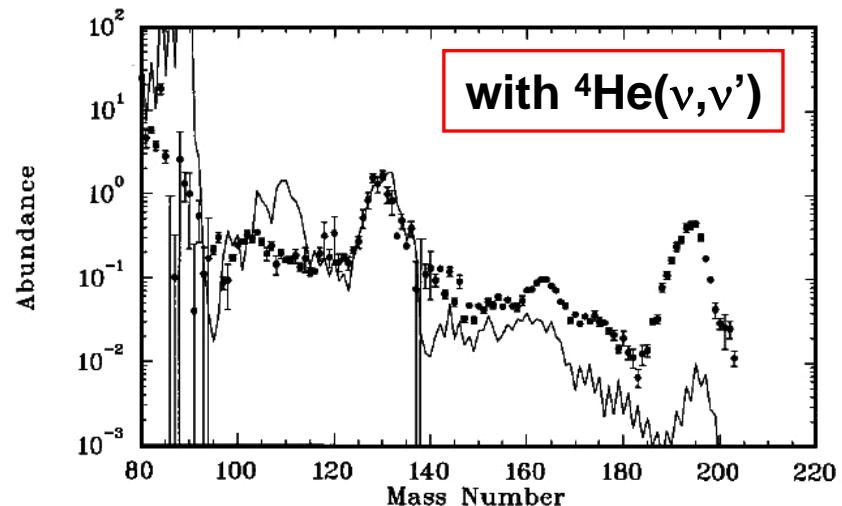
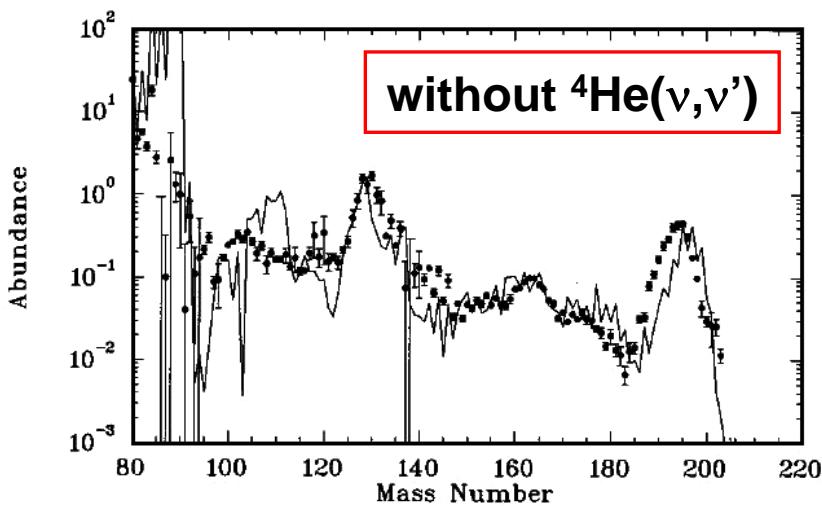
$$E_\gamma = 16 - 40 \text{ MeV}, \Phi_\gamma = 10^{4-5} \text{ photons/MeV/s}, \Delta E_\gamma/E_\gamma = 3 \sim 10\%$$

# **Merits of NewSUBARU LCS $\gamma$ -source**

- Conventional lasers (Nd:YVO<sub>4</sub>, CO<sub>2</sub>, etc.) are used.
  - Various kinds of lasers can be employed.  
Flexibility in energy range, pulse structure, polarization, etc.
- Laser and electron beam collide in a straight section.
  - Little background from bremsstrahlung
- Top-up (top-loaded) operation is available.
  - High flux and excellent stability



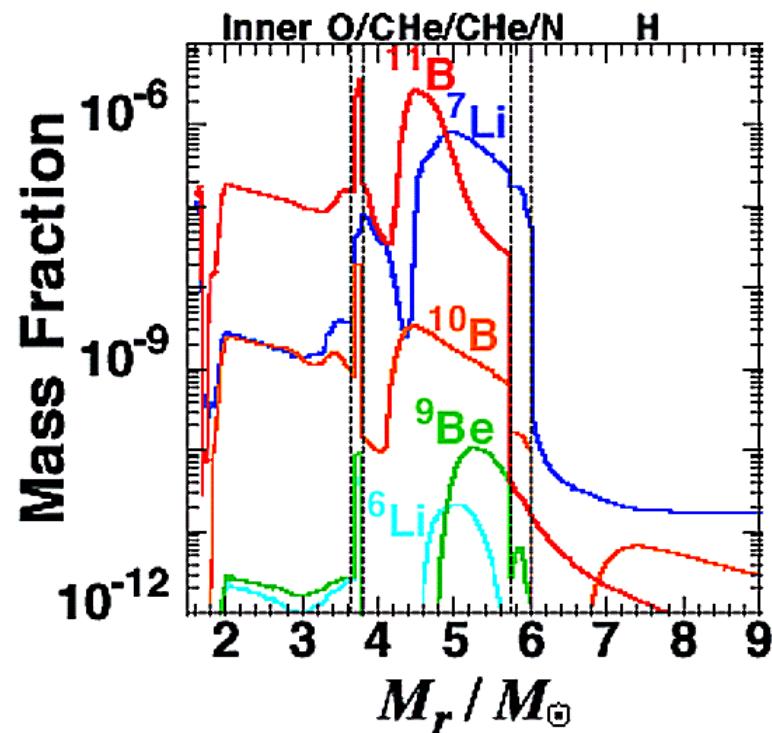
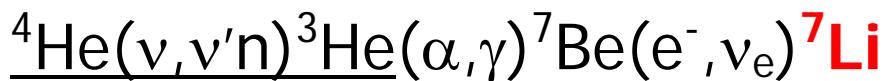
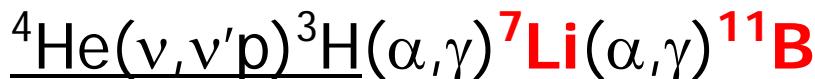
- Neutrino-heating effect  needs CS with 10% accuracy
- R-process in neutrino-driven wind



Meyer, ApJ449 (1995) L55

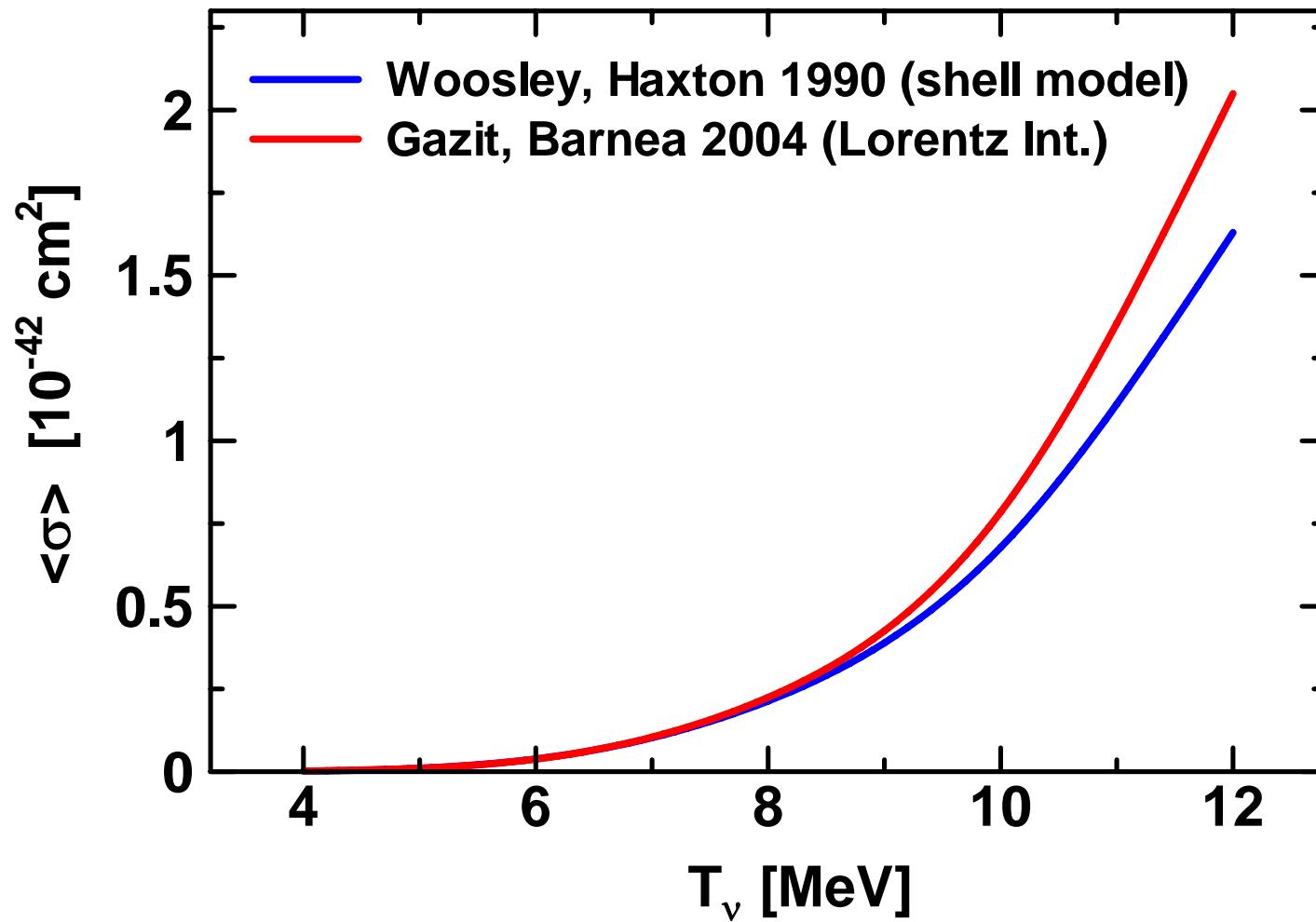
# $^7\text{Li}$ , $^{11}\text{B}$ production by $\nu$ -spallations

Woosley et al., Woosley & Weaver, Rauscher et al., Yoshida et al.



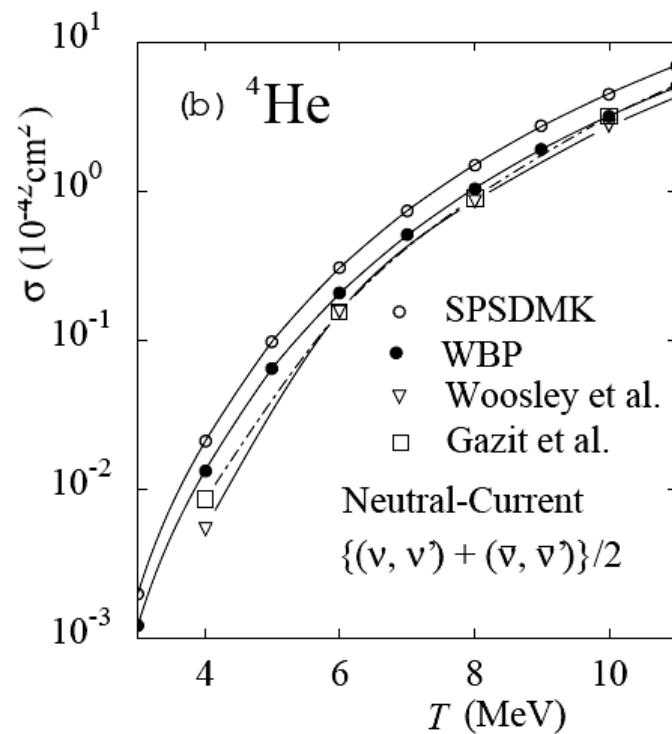
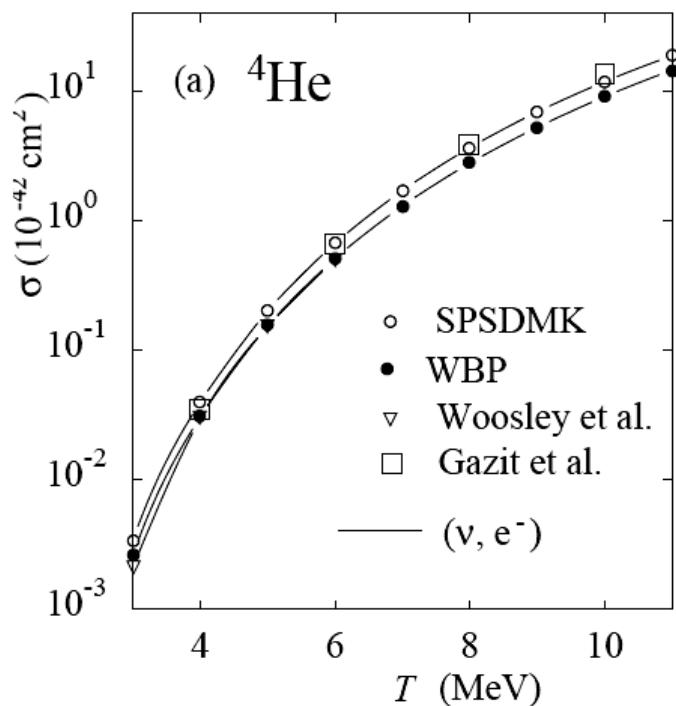
Yoshida, Kajino et al.,  
PRL96, 091101 (2006)

# Neutrino neutral reaction rate on ${}^4\text{He}$

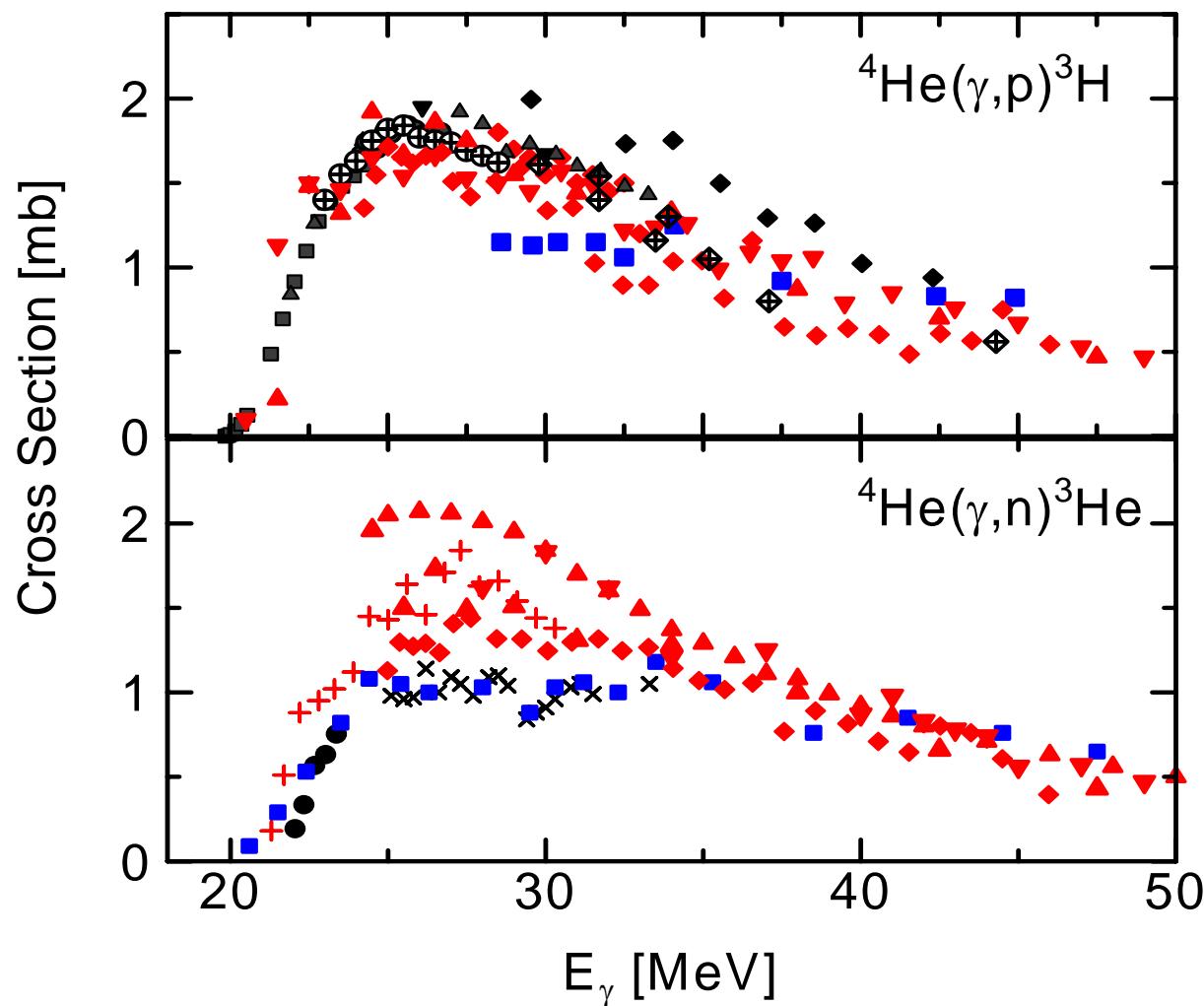


# Shell-model calculation

T. Suzuki et al., PR C74 034307 (2006)

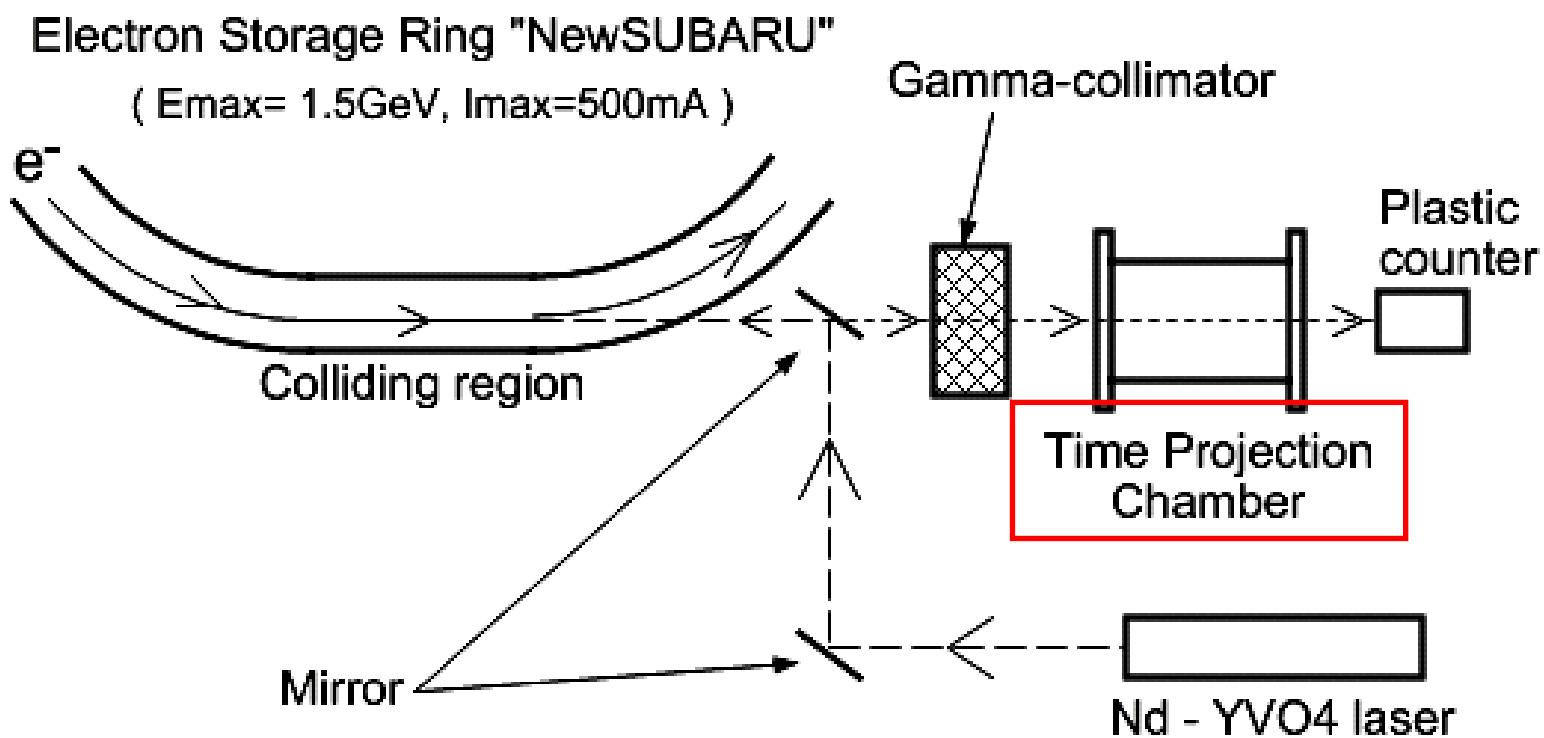


Photodisintegration of  ${}^4\text{He}$  have been studied by means of monochromatic  $\gamma$ , bremsstrahlung, radiative capture...



# Experiment at NewSUBARU LCS- $\gamma$ facility

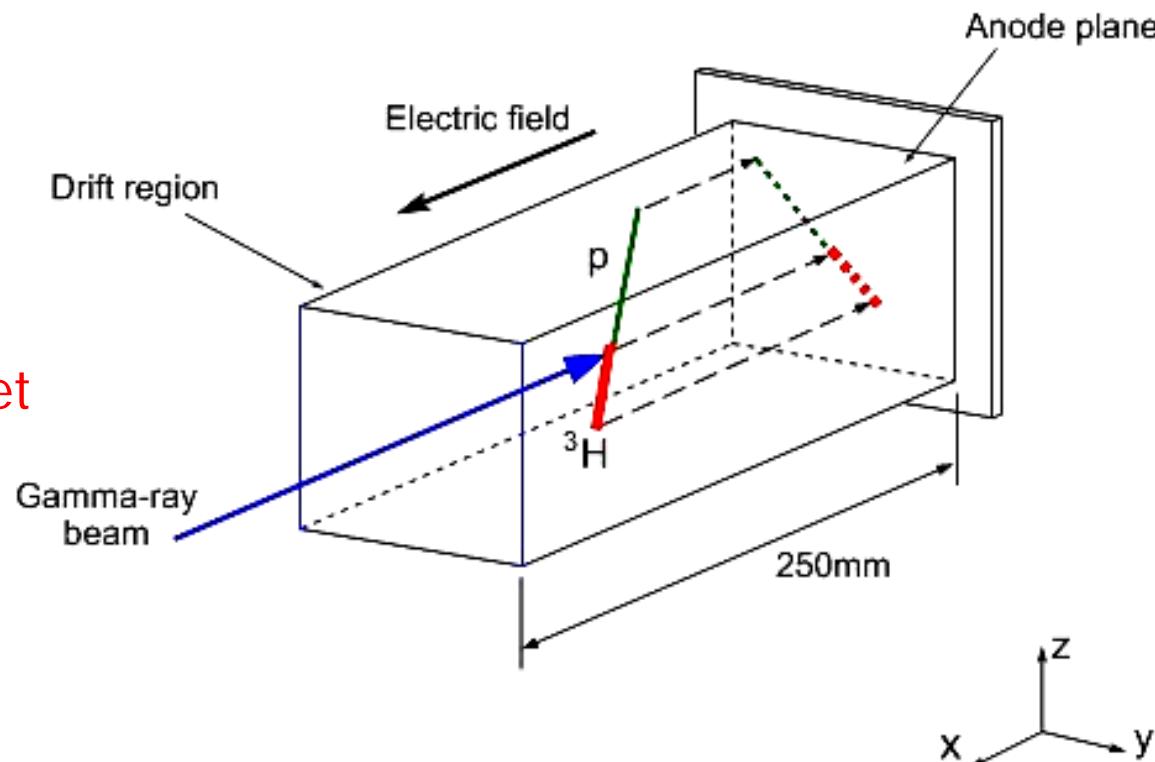
$E_\gamma = 16 \sim 40\text{MeV}$ ,  $\Phi_\gamma = 2\sim 4 \times 10^4 / \text{sec}$ , FWHM~9%, P~100%



# Time Projection Chamber

T. Kii, T. Shima, T. Baba, Y. Nagai, NIM A552 (2005) 329

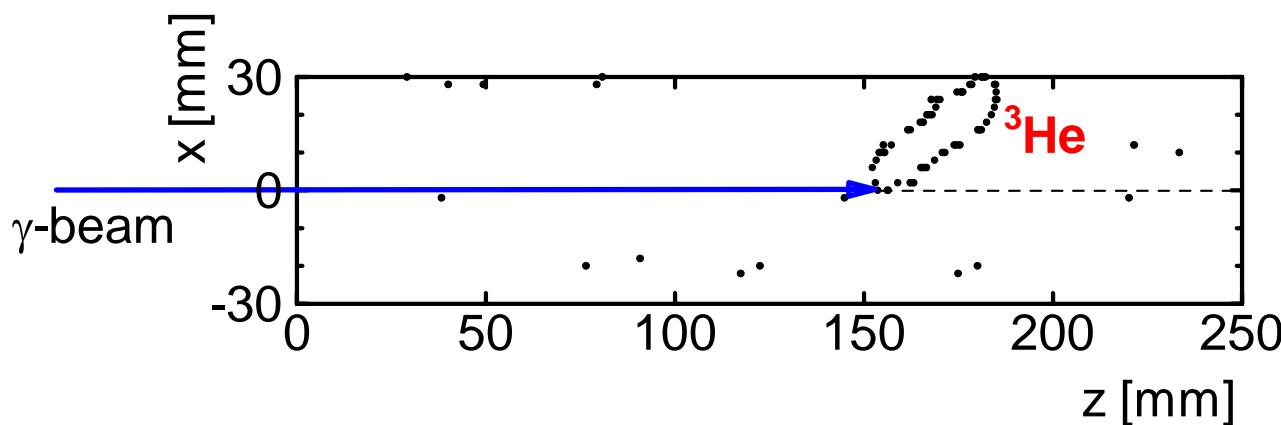
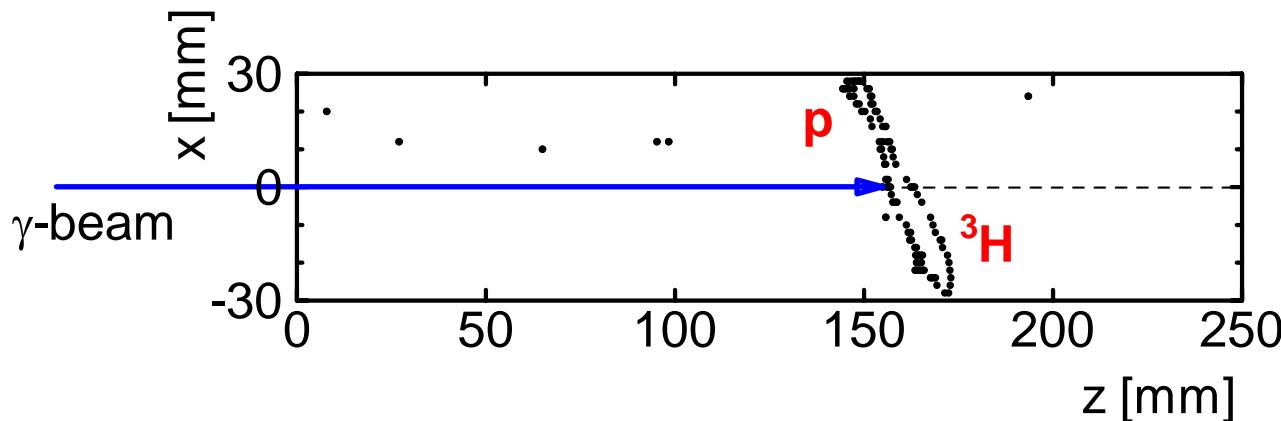
Target gas :  
He + CD<sub>4</sub>  
--- active target



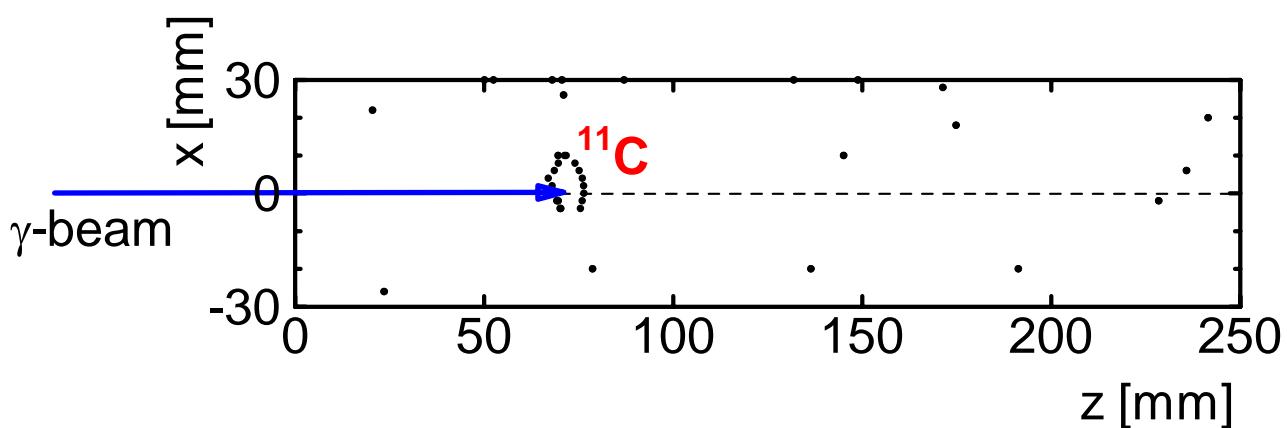
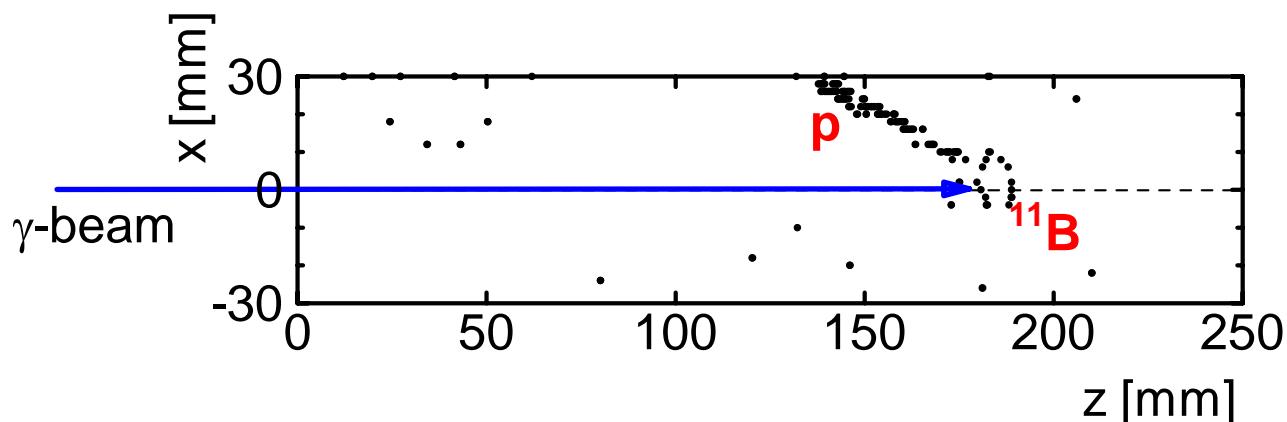
- $\Omega \sim 4\pi$ ,  $\varepsilon > 98\%$  ; high efficiency
- track shape,  $dE/dx \Rightarrow$  event ID,  $d\sigma/d\Omega$ , asymmetry

# Event Identification

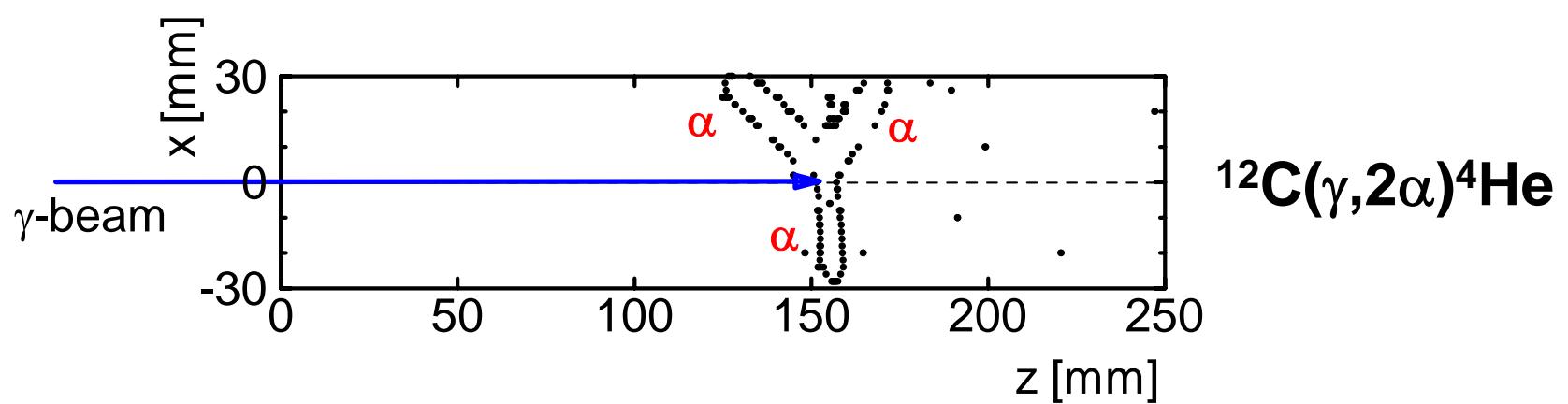
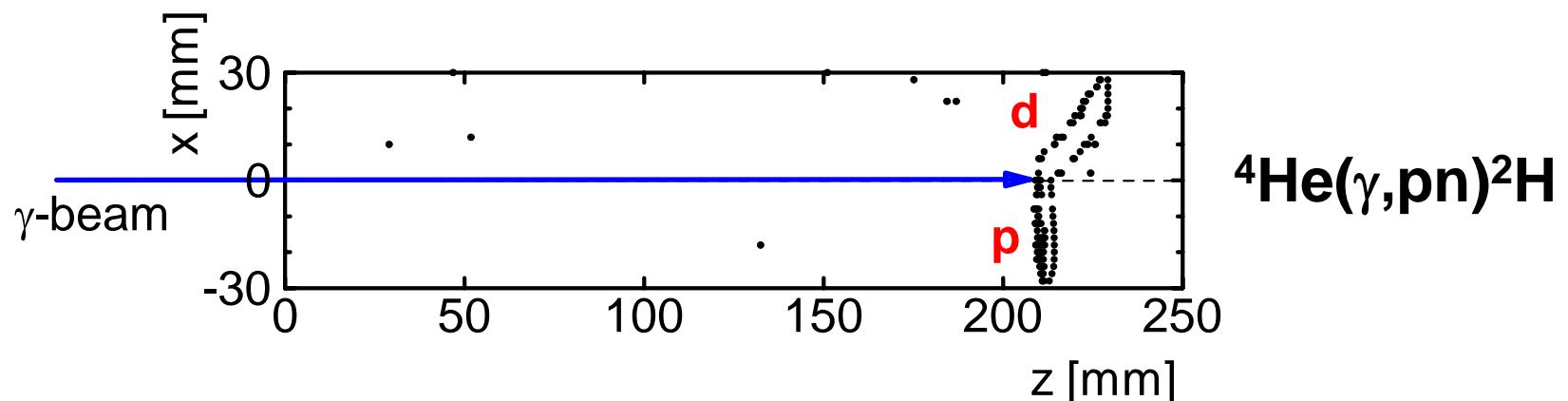
## $^4\text{He}$ photodisintegrations



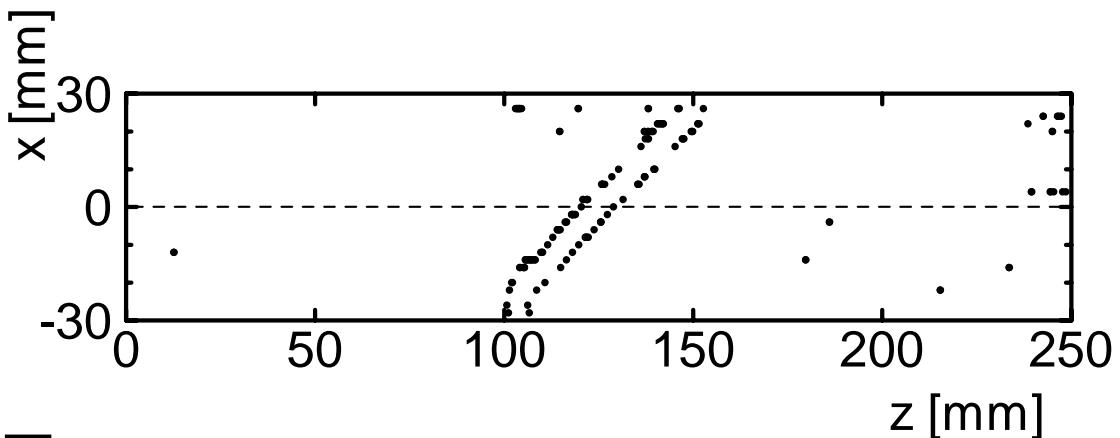
# $^{12}\text{C}$ photodisintegrations



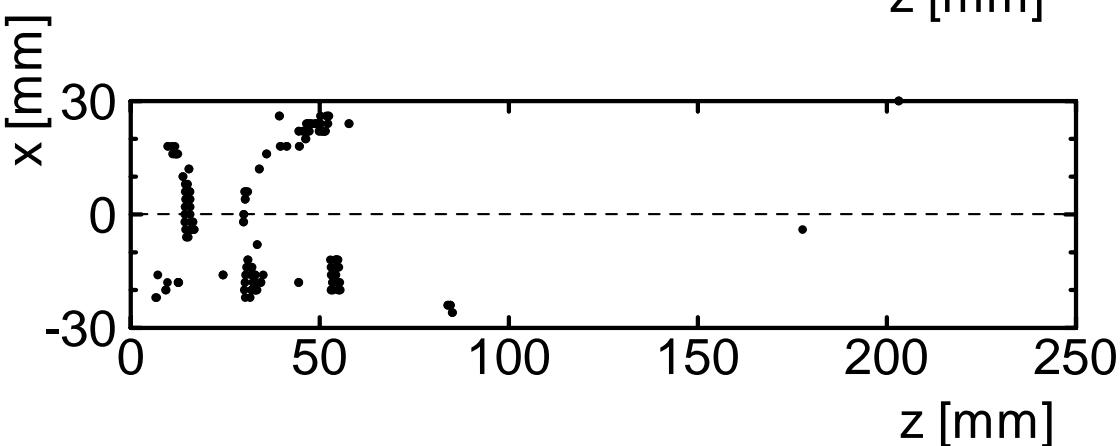
# Three-body decays



# Backgrounds

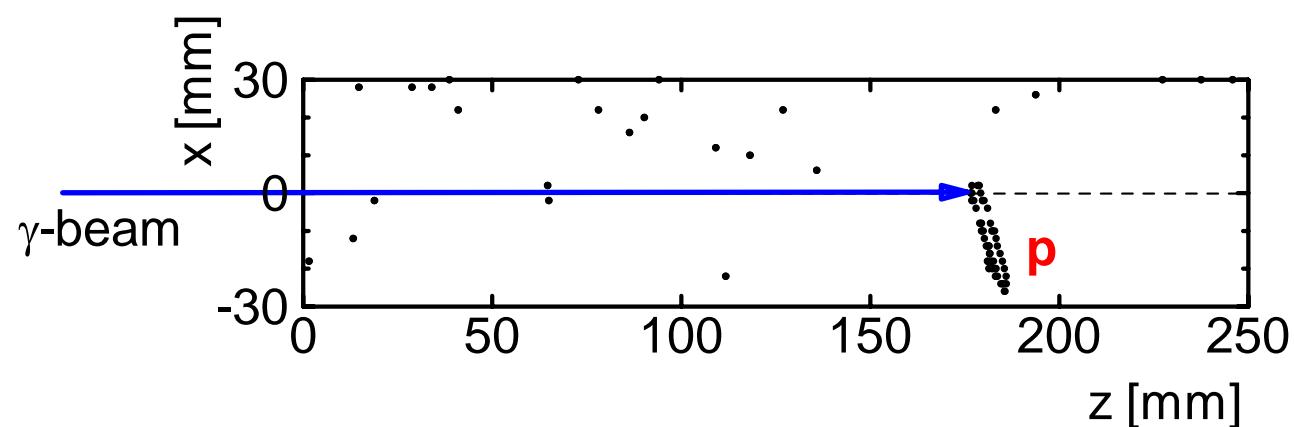


**$\alpha$  from natural RI**

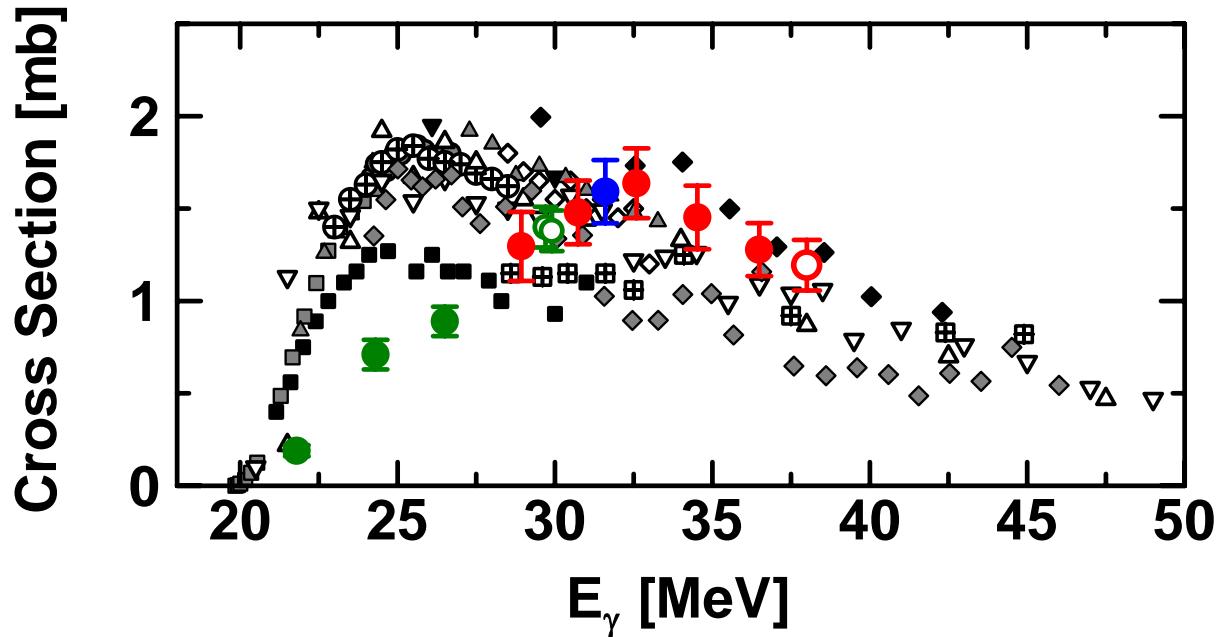


**Photoelectrons**

# $D(\gamma, p)n$

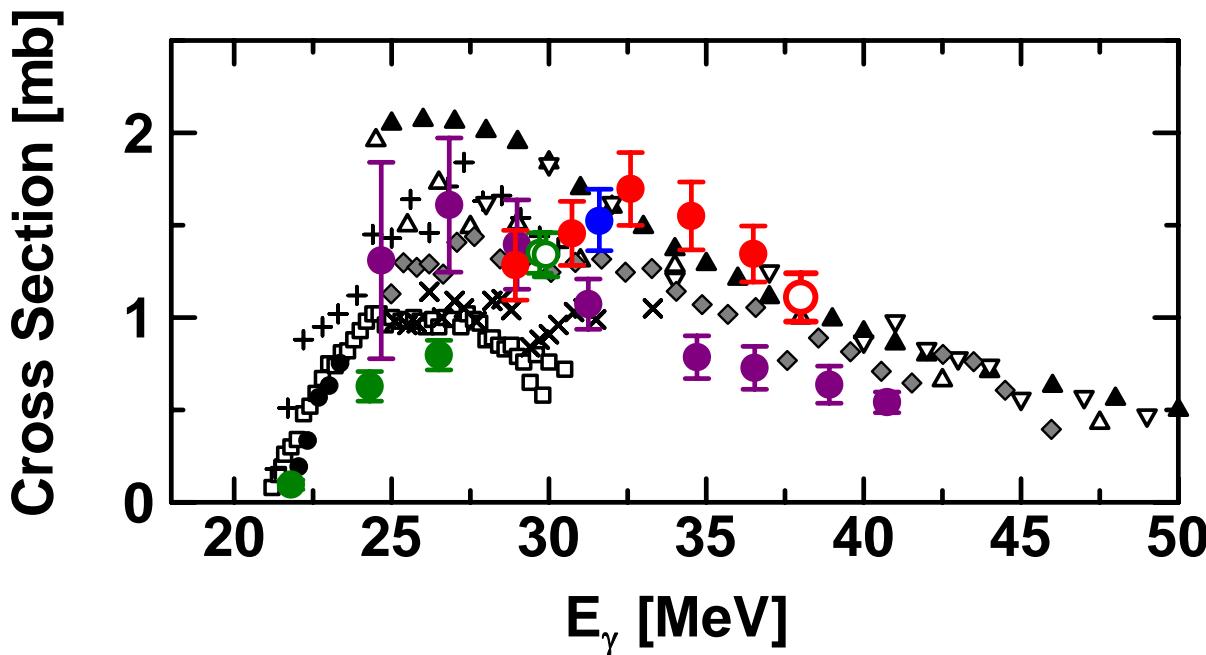


# ${}^4\text{He}(\gamma, \text{p}) {}^3\text{H}$ (preliminary)



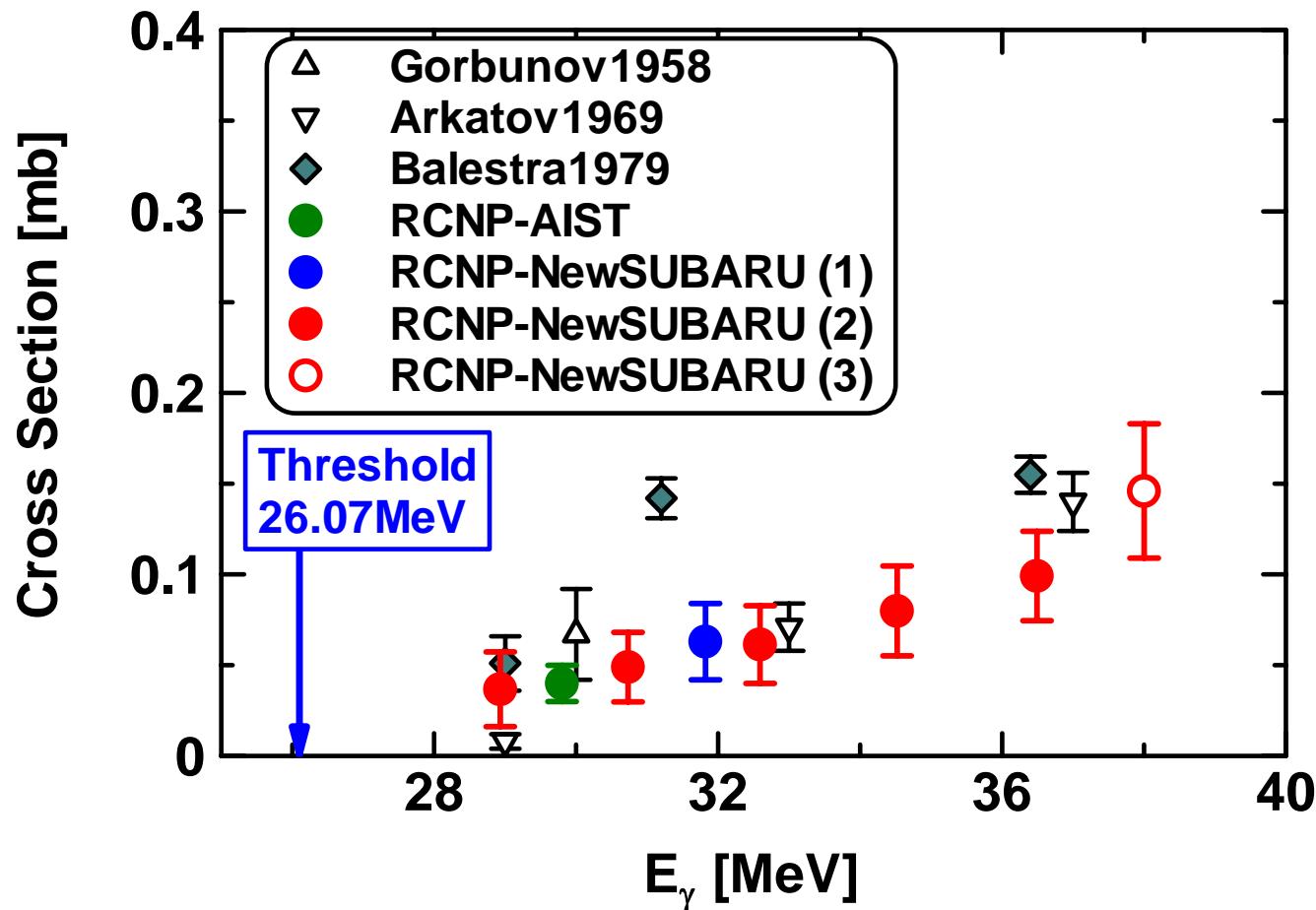
- RCNP-AIST (PRC72, 044004) ;  $\lambda=351\text{nm}$  (3rd),  $E_e=0.8\text{GeV}$
- RCNP-NewSUBARU;  $\lambda=532\text{nm}$  (2nd),  $E_e=0.97\text{GeV}$
- RCNP-NewSUBARU;  $\lambda=1064\text{nm}$  (fund.),  $E_e\leq1.46\text{GeV}$
- RCNP-NewSUBARU;  $\lambda=532\text{nm}$  (2nd),  $E_e=1.06\text{GeV}$

# ${}^4\text{He}(\gamma, \text{n}) {}^3\text{He}$ (preliminary)

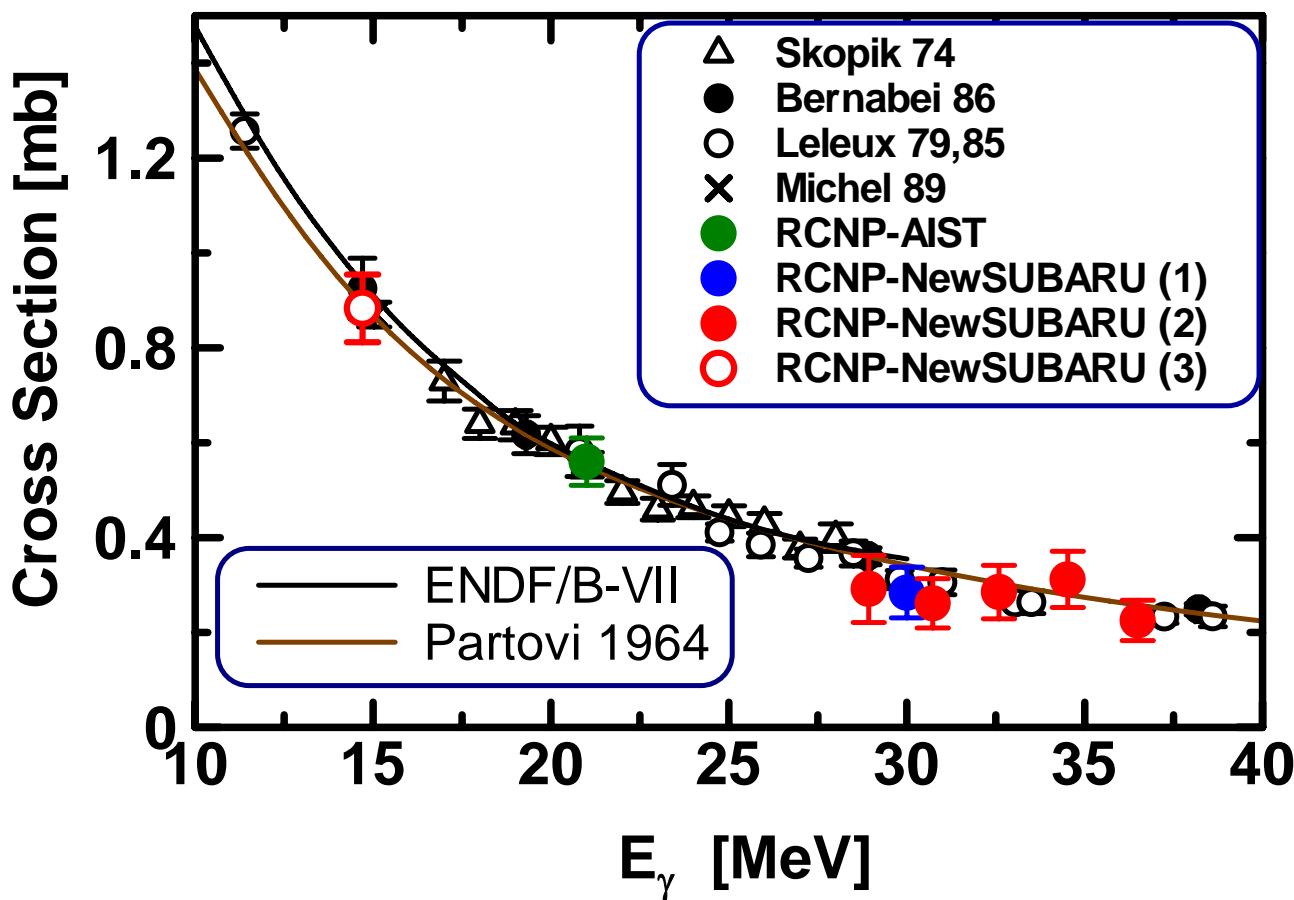


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- RCNP-NewSUBARU;  $\lambda=532\text{nm}$  (2nd),  $E_e=1.06\text{GeV}$
- Lund 2005-2007 (PRC75, 014007) ; tagged photons

# ${}^4\text{He}(\gamma, \text{pn})\text{d}$ (preliminary)



# D( $\gamma$ ,n)p



# Summary

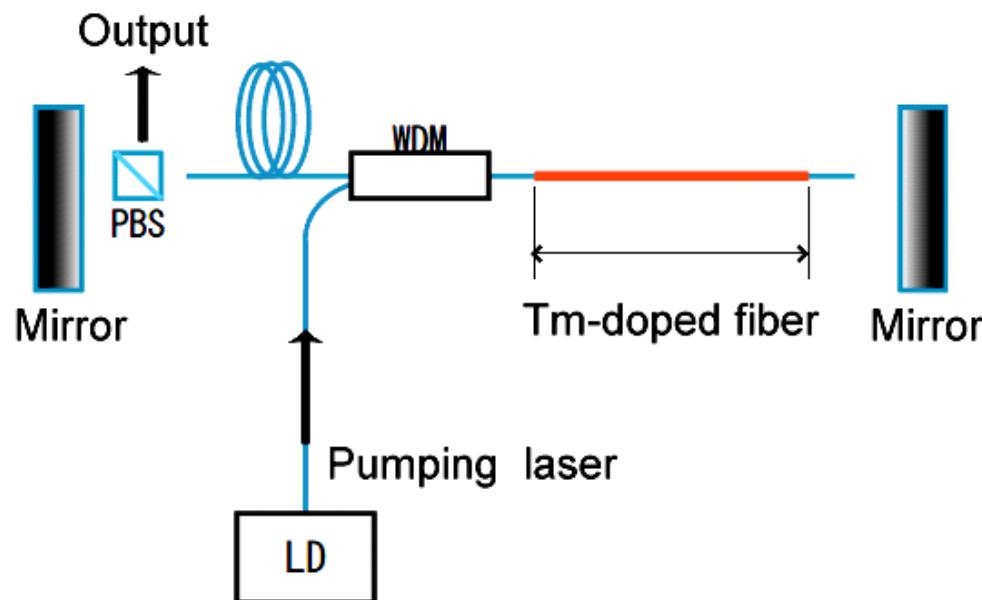
- Laser Compton Scattered  $\gamma$ , combined with new detector techniques, provides a useful tool for high-precision studies of nuclear reactions induced by photons and also neutrinos.
- Complementary to radiative capture, CD, (p,p'), ...
- The LCS  $\gamma$ -source at NewSUBARU is now in operation, and several works for nuclear astrophysics have been performed successfully.
- $^4\text{He}$ : excitation functions are measured up to 37MeV.  
Small cross sections below 30MeV were confirmed.  
GDR peak was found to locate at  $\sim 32.5\text{MeV}$ .  
 $\rightarrow \rightarrow \rightarrow p\text{-}n, 1p1h, 3NF$
- Other light nuclei and heavier p-nuclei are planned to be studied.

# Outlook

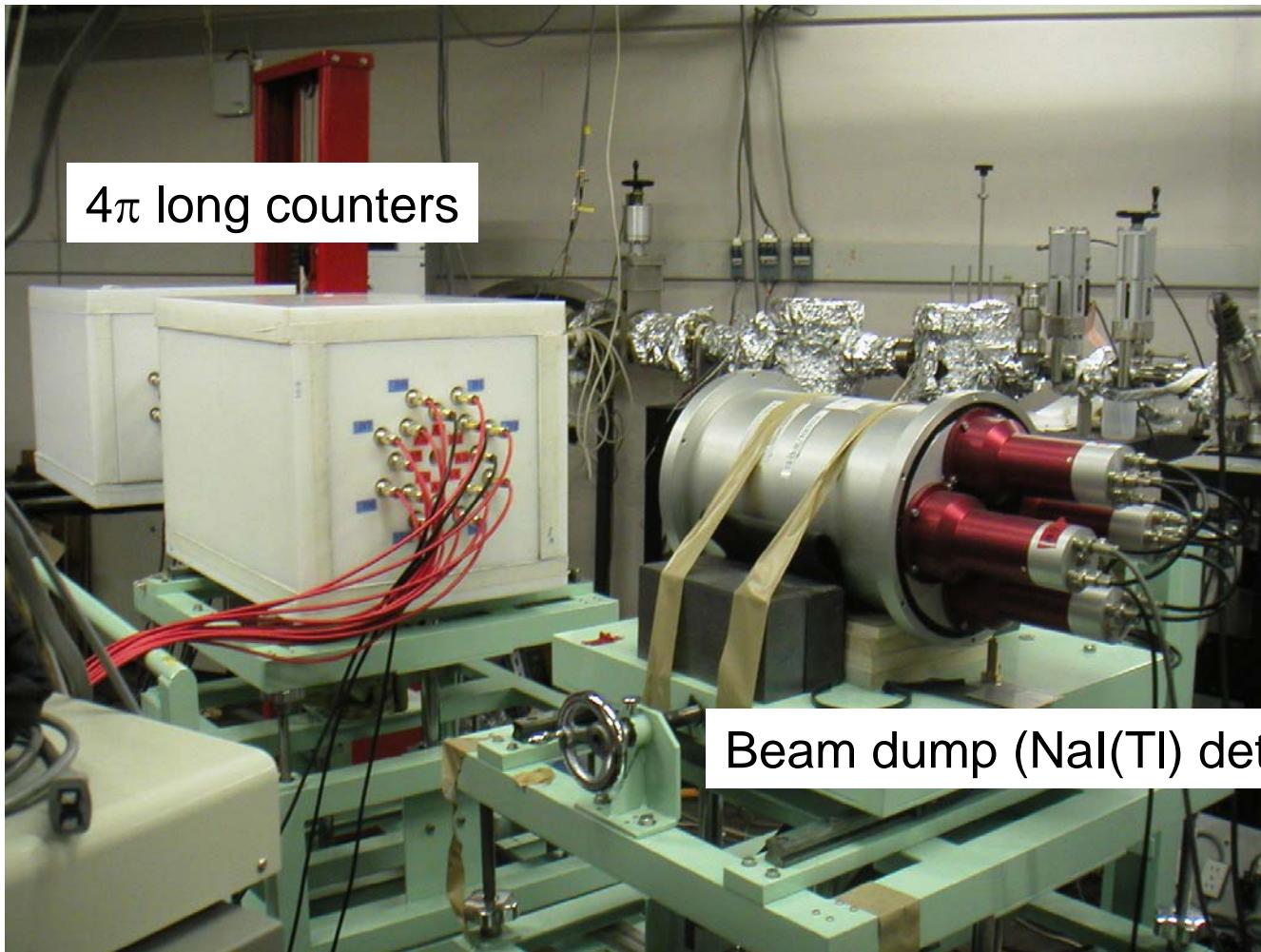
- ( $\gamma, n$ ) cross sections of long-lived radioactive nuclide ;  
$$\Phi_{\gamma} = 10^4 \text{ /s} \Rightarrow 10^6 \text{ /s} ; \text{ target: } 1\text{g} \Rightarrow 10\text{mg}$$
- ( $\gamma, p$ ), ( $\gamma, a$ ) cross sections of p-nuclei ;  
for noble gases --- TPC  
for others --- activation method
- Reference data for indirect methods ; CD, ( $p, p'$ ), ...
- Technical hint for experiment with other neutral beams ; n,  $\nu$

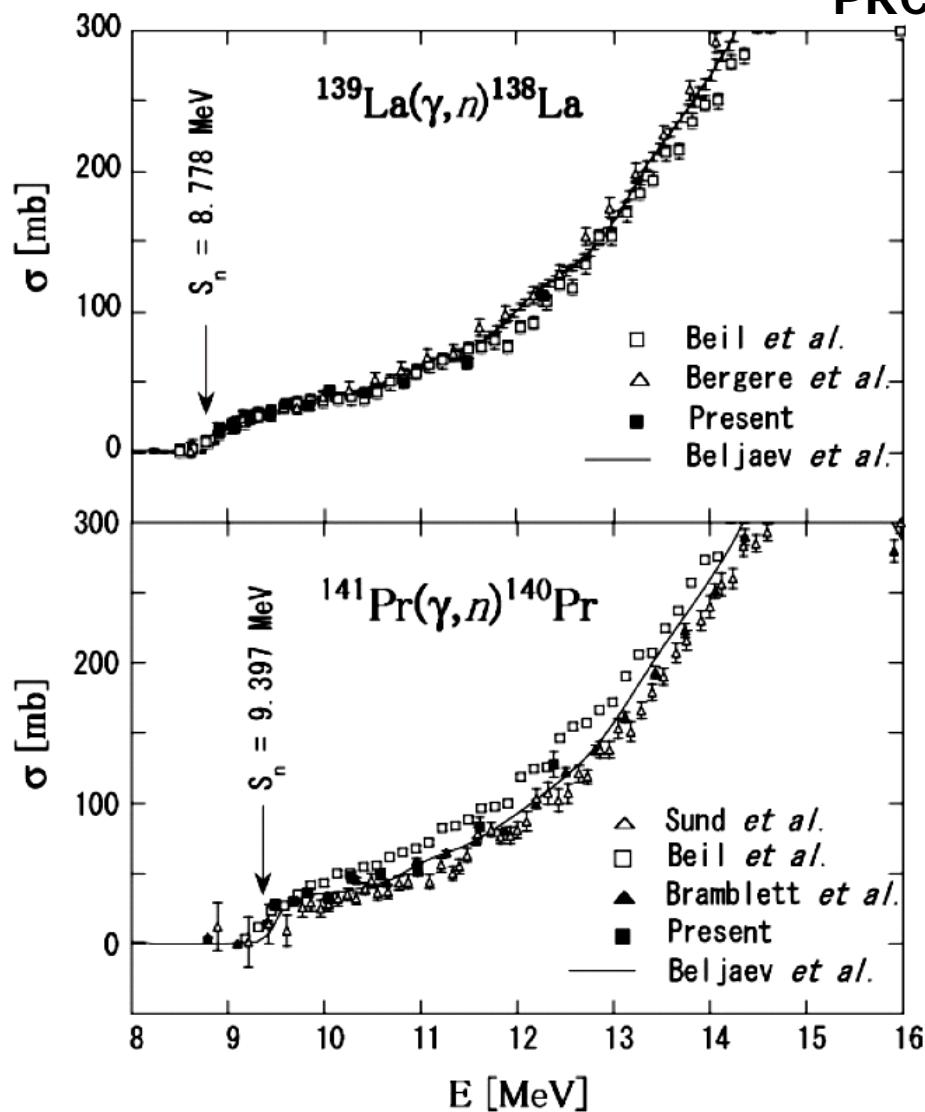
Present; Nd:YVO<sub>4</sub> laser (1.06μm), 1W  
⇒ E<sub>γ</sub> = 16.5MeV, Φ<sub>γ</sub> ~ 3 × 10<sup>4</sup> photons/s

Plan; Tm-doped fiber laser (2.05μm), 100W  
⇒ E<sub>γ</sub> = 8.5MeV, Φ<sub>γ</sub> > 10<sup>6</sup> photons/s



# Neutron detector for ( $\gamma$ ,n) (Konan Univ.)





Radioactivity < 1MBq, 100 $\mu$ mol  $\longleftrightarrow$   $T_{1/2} > \sim 10000\text{yr}$

$\beta^\pm$ - $\gamma$ ,  $T_{1/2} > 10^4\text{y}$  (15 nuclides) :

$^{10}\text{Be}$ ,  $^{26}\text{Al}$ ,  $^{36}\text{Cl}$ ,  $^{40}\text{K}$ ,  $^{59}\text{Ni}$ ,  $^{79}\text{Se}$ ,  $^{93}\text{Zr}$ ,  $^{92}\text{Nb}$ ,  $^{98,99}\text{Tc}$ ,  $^{107}\text{Pd}$ ,  
 $^{135}\text{Cs}$ ,  $^{138}\text{La}$ ,  $^{186}\text{Re}$ ,  $^{208}\text{Bi}$

EC, no  $\gamma$ ,  $T_{1/2} > 100\text{d}$  (12 nuclides) :

$^{41}\text{Ca}$ ,  $^{49}\text{V}$ ,  $^{53}\text{Mn}$ ,  $^{55}\text{Fe}$ ,  $^{93}\text{Mo}$ ,  $^{97}\text{Tc}$ ,  $^{109}\text{Cd}$ ,  $^{137}\text{La}$ ,  $^{179}\text{Ta}$ ,  $^{178}\text{W}$ ,  $^{193}\text{Pt}$ ,  $^{205}\text{Pb}$

$\alpha$ - $\gamma$ ,  $T_{1/2} > 10^4\text{y}$  (16 nuclides) :

$^{210}\text{Bi}$ ,  $^{230,232}\text{Th}$ ,  $^{231}\text{Pa}$ ,  $^{233-236,238}\text{U}$ ,  $^{236,237}\text{Np}$ ,  $^{239,242,244}\text{Pu}$ ,  $^{247,248}\text{Cm}$

$\alpha$ , no  $\gamma$ ,  $T_{1/2} > 100\text{d}$  (4 nuclides) ;  $^{146}\text{Sm}$ ,  $^{148,150}\text{Gd}$ ,  $^{154}\text{Dy}$

# Outlook

- ( $\gamma, n$ ) cross sections of long-lived radioactive nuclide ;  
$$\Phi_{\gamma} = 10^4 \text{ /s} \Rightarrow 10^6 \text{ /s} ; \text{ target: } 1\text{g} \Rightarrow 10\text{mg}$$
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- Technical hint for experiment with other neutral beams ; n,  $\nu$

# Collaborators

**H. Utsunomiya, H. Akimune**

*Department of Physics, Konan University*

**T. Mochizuki, S. Miyamoto, K. Horikawa**

*Laboratory for Advanced Science and Technology for Industry,  
University of Hyogo*

**T. Hayakawa, T. Shizuma**

*Kansai Photon Science Institute, Japan Atomic Energy Agency*

**M. Fujiwara**

*Research Center for Nuclear Physics, Osaka University*

**Y. Nagai**

*Nuclear Science and Engineering Directorate,  
Japan Atomic Energy Agency*